

Regular Article - Theoretical Physics

Energies and radii of light nuclei around unitarity

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Abstract Light nuclei fall within a regime of universal physics governed by the fact that the two-nucleon scattering lengths are large compared to the typical nuclear interaction range set by one-pion exchange. This places nuclear physics near the so-called unitarity limit in which the scattering lengths are exactly infinite. Effective field theory provides a powerful theoretical framework to capture this separation of scales in a systematic way. It is shown here that the nuclear force can be constructed as a perturbative expansion around the unitarity limit and that this expansion has good convergence properties for both the binding energies of A=3,4 nuclei as well as for the radii of these states.

1 Effective field theory for systems near unitarity

Nuclear physics at very low energies hosts a fascinating emergent phenomenon: out of the tremendously complicated dynamics of quarks and gluons, governed by the strong interaction (Quantum Chromodynamics, QCD) that is highly nonperturbative in this regime, ultimately arise strikingly simple features for systems of few nucleons. It was realized many decades ago [1–4] that the low-energy two-nucleon system can be parameterized with a formula that has become famous, in nuclear physics and beyond, as the effective range expansion (ERE):

$$k \cot \delta_0(k) = -\frac{1}{a} + \frac{r}{2}k^2 + \cdots$$
 (1)

Here $\delta_0(k)$ denotes the *S*-wave scattering phase shift for two particles (here, nucleons in a single fixed spin configuration) with relative momentum k. The leading parameter in this expansion, called "scattering length" and denoted by a, governs the nucleon-nucleon (NN) cross section at low energies and completely determines it in the limit where the relative momentum k of the nucleons goes to zero:

$$\sigma = 4\pi a^2 + \mathcal{O}(k^2). \tag{2}$$

Empirically, in the 3S_1 ("t") and 1S_0 ("s") NN spin channels the values are known to be $a_t \simeq 5.4$ fm and $a_s \simeq -23.7$ fm, respectively. Compared to the typical range of the nuclear interaction, set by one-pion exchange providing the longest-range component as $R \sim M_\pi^{-1} \simeq 1.4$ fm, these scattering lengths are unnaturally large, $a_{s,t} \gg R$, qualitatively. Through Eq. (2) this implies that the nuclear force is particularly strong as the energy goes to zero. The fact that this is so can be understood as an accidental "fine tuning" of the QCD parameters [5–9] (the quark masses, in particular) to be close to a critical point where the scattering lengths diverge. This point is called the "unitarity (or unitary) limit," and it is the heart of the emergent simplicity mentioned at the outset.

Systems near the unitarity limit exhibit universal features. As the two-body scattering length becomes large, the details of the underlying interaction largely cease to matter and to a very good approximation the behavior of the system is determined qualitatively by the fact that *a* is large, and quantitatively by how large exactly it is. This phenomenon places low-energy nuclear systems in a common universality class with other systems near unitarity, such as cold atomic gases (where the scattering length can be tuned experimentally *via* Feshbach resonances [10]), nuclei with a halo/cluster structure [11], or certain mesons which can be interpreted as hadronic molecules [12].

In the two-body sector, universality relates scattering parameters to shallow bound and virtual states. This is a consequence of Eq. (1) and the principle of analyticity: the ERE provides an expansion of the S matrix, so whenever poles at complex momenta—in particular bound and virtual states, which reside at purely imaginary momenta—fall within the radius of convergence of the expansion, they are described by the same parameters. Since, schematically, the S-wave S matrix is given by 1 + it with



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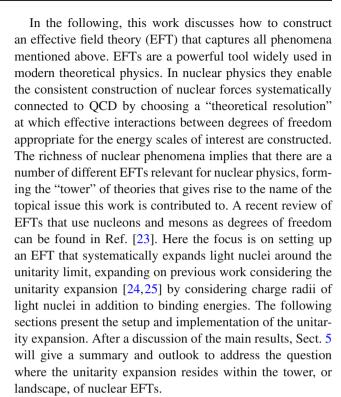
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$$t(k) \sim \frac{1}{k \cot \delta_0(k) - ik},\tag{3}$$

the pole condition is $\cot \delta_0(k) = i$ for some $k = i\kappa$. Keeping only the first term in the ERE this gives $\kappa = 1/a$ and one sees that for positive (negative) a one has a bound (virtual) state in the system. Having a sufficiently large |a| ensures that indeed these momentum scales lie within the radius of convergence of the ERE, which for nucleons is determined by the position of the pion cut, $M_\pi/2$. For the two-nucleon system at the physical point one has the deuteron as a shallow bound state $(B_D \simeq 2.224 \text{ MeV})$, with the difference to $1/(M_N a_t^2) \approx 1.41 \text{ MeV}$ being due to range corrections) in the 3S_1 channel, and a very shallow virtual state at $B_{NN^*} \simeq 0.068 \text{ MeV}$ (with a relatively small range correction since $|a_s|$ is so large). In the unitarity limit, $a_{s,t} \to \infty$, both of these poles become zero-energy S-wave states.

A more striking universal behavior is encountered for three and more particles: in the unitarity limit there exists an infinite tower of three-body bound states, geometrically spaced (the binding energy of each subsequent state is given by a fixed factor times the previous level) and accumulating at zero energy, a phenomenon that has become famous as the Efimov effect [13]. At large but finite scattering length the spectrum is cut off in the infrared due to the existence of a two-body pole in the S matrix. It was shown in Refs. [14–16] that for physical values of the NN scattering lengths the triton can be interpreted as the single remaining bound state of such an Efimov tower. More recently it was established in a model-independent way [17] that a virtual state in the threenucleon (3N) system, known to exist for a long time [18, 19], is in fact an S-matrix pole that would be an excited Efimov state instead if nature were just slightly closer to the unitarity limit. This confirms a relation previously observed in a separable potential model [20].

The phenomenon continues at the four-body level. At unitarity, each three-boson Efimov state (with binding energy B_3) is associated with two four-boson states [21]. One of these is almost five times as deeply bound as the trimer, $B_4/B_3 \simeq 4.611$, while the other resides just below the particle-trimer threshold, $B_{4*}/B_3 \simeq 1.002$ [22]. Universality implies that if the NN scattering lengths were infinite, the ground state of ⁴He would be located at 4.6 times the binding energy of the triton (neglecting Coulomb and other isospin breaking effects). In nature, the ground state is at $B_{\alpha}/B_{H} \simeq 3.66$, and there exists a 0⁺ resonance state just above the proton-triton threshold, i.e., one has $B_{\alpha^*}/B_H \simeq 1.05$, where $B_H \simeq 7.72$ MeV is the ³He binding energy, taken as reference here to at least partially account for isospin breaking effects. The closeness of these ratios to the unitarity-limit values is a strong indication that nature may be perturbatively close to unitarity for systems of at least four nucleons.



As a variant of what has become known as "Pionless EFT," the unitarity expansion is defined in terms of a Lagrange density

$$\mathcal{L} = N^{\dagger} \left(i \mathcal{D}_0 + \frac{\mathcal{D}^2}{2M_N} \right) N$$

$$+ \sum_{\mathbf{i}} C_{0,\mathbf{i}} \left(N^T P_{\mathbf{i}} N \right)^{\dagger} \left(N^T P_{\mathbf{i}} N \right)$$

$$+ D_0 \left(N^{\dagger} N \right)^3 + \cdots, \tag{4}$$

where the notation has been taken over from Refs. [24,26]. The degrees of freedom are nonrelativistic nucleon isospin doublets $N = (p \ n)^T$, coupled to photon fields A_{μ} via the covariant derivative $\mathcal{D}_{\mu} = \partial_{\mu} + ieA_{\mu}(1 + \tau_3)/2$, where e is the proton charge and τ_a is used to label isospin Pauli matrices. Of the electromagnetic interactions, only the static Coulomb potential is relevant up to high orders (defined later), where corrections from transverse photons will eventually enter. The strong interaction is parameterized in Eq. (4) by the "low-energy constants (LECs)" $C_{0,i}$ and D_0 , defining contact (zero-range) interactions without derivatives between two and three nucleons, respectively. The P_i denote projectors onto the NN S waves, $\mathbf{i} = {}^{1}S_{0}$, ${}^{3}S_{1}$. Contact interactions with derivatives as well as higher-body forces are contained in the ellipses in Eq. (4), along with other interactions not shown explicitly here.

The ellipses in Eq. (4) represent a fundamental feature of an EFT, namely that the Lagrangian contains all possible terms which are allowed by the symmetries of the system at



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hand. For the EFT of nucleons considered here these symmetries are inherited from QCD as the underlying theory: each term in Eq. (4) is required to be invariant under Galilean boosts (plus systematic relativistic corrections), rotations, isospin, and other discrete symmetries. It is of course not arbitrary that Eq. (4) explicitly shows some terms but not others. In order to be predictive, each EFT comes with an organizational principle called "power counting," which attributes the various terms to increasingly higher orders. A starting point for this organization is typically a naïve dimensional analysis (NDA): fields and derivatives acting on them are assigned their canonical dimensions, defining the exponent of a typical low-momentum scale Q. In order for each term to have overall dimension four, appropriate powers of the EFT breakdown scale $M_{\rm hi}$ are included in the prefactor. For the standard Pionless EFT expansion, $M_{\rm hi} \sim R^{-1} \sim M_{\pi}$, and this is kept for the construction of the unitarity expansion. However, while standard Pionless EFT assumes $Q \sim 1/a_{s.t}$, Ref. [24] suggested to count these scales separately as $\aleph \sim 1/a_{s,t}$ while assuming that

$$Q \sim Q_A = \sqrt{2M_N B_A/A}. (5)$$

This is a momentum scale associated with the binding energy per nucleon in an A-nucleon system, which for A=2 coincides with the canonical definition of the binding momentum. With this assumption one obtains $\aleph < Q < 1/R$ such that it is possible to set up a *combined* expansion in two parameters \aleph/Q and QR. Coulomb effects are perturbative for momenta of order Q_A as well and are naturally captured by the expansion if one takes into account that the Coulomb momentum scale $k_C = \alpha M_N$, with the fine-structure constant $\alpha \approx 1/137$, is naturally included in the \aleph scale [26].

For a calculation of few-body states it is convenient to switch from the Lagrangian formulation to standard quantum mechanics expressed in terms of potentials. In the two-body sector, it is possible to write

$$V_{2,\mathbf{i}}^{(0)} = C_{0,\mathbf{i}}^{(0)}|g\rangle\langle g|,\tag{6}$$

where $C_{0,\mathbf{i}}^{(0)}$ is the leading-order (LO) piece of the non-derivative contact LEC $C_{0,\mathbf{i}}$ in Eq. (4), which has an expansion of the form

$$C_{0,\mathbf{i}} = C_{0,\mathbf{i}}^{(0)} + C_{0,\mathbf{i}}^{(1)} + \cdots$$
 (7)

Apart from this, $V_{2,i}^{(0)}$ is defined in terms of a separable Gaussian regulator function, given by

$$\langle p|g\rangle = g(p^2) = \exp(-p^2/\Lambda^2) \tag{8}$$

in momentum space. This makes the zero-range theory well defined by regularizing the otherwise divergent interaction via the introduction of a cutoff scale Λ . Both the value of Λ and the particular form of the regulator function are arbitrary

and renormalization, discussed below, ensures that observables are independent of these choices. The separable form is however particularly convenient for the formalism explained in the following. It makes it possible to algebraically solve the Lippmann-Schwinger equations for the LO T matrices $t_{\mathbf{i}}^{(0)}$,

$$t_{\mathbf{i}}^{(0)} = V_{2,\mathbf{i}}^{(0)} + V_{2,\mathbf{i}}^{(0)} G_0 t_{\mathbf{i}}^{(0)}, \tag{9}$$

where

$$G_0(z) = \frac{1}{z - H_0} \tag{10}$$

defines, for an arbitrary energy z, the two-body Green's function in terms of the free (purely kinetic) Hamiltonian H_0 . The result is:

$$t_{\mathbf{i}}^{(0)}(z; \mathbf{k}, \mathbf{k}') = \langle \mathbf{k} | t_{\mathbf{i}}^{(0)} | \mathbf{k}' \rangle = g(k^2) \tau_{\mathbf{i}}(z) g(k'^2), \tag{11}$$

$$\tau_{\mathbf{i}}(z) = \left[1/C_{0,\mathbf{i}}^{(0)} - \langle g|G_0|g\rangle \right]^{-1}.$$
 (12)

The regulator ensures that $\langle g|G_0|g\rangle$ is finite. At the on-shell point, $E=k^2/M_N$ and $\mathbf{k}=\mathbf{k}'$, this solution can be matched directly to the ERE, yielding

$$C_{0,\mathbf{i}}^{(0)} \to C_{0,\mathbf{i}}^{(0)}(\Lambda) = \frac{1}{2\pi^2 M_N} \left(\frac{1}{a_{\mathbf{i}}} - \theta_0 \Lambda\right)^{-1}.$$
 (13)

With this running coupling appearing in Eq. (6), the two-body sector of the theory is renormalized. The number θ_0 in general depends on the choice of regulator; for the Gaussian form used here one finds $\theta_0 = 1/\sqrt{2\pi}$. In the unitarity limit, $1/a_i = 0$, such that the leading-order two-body interaction is parameter free:

$$C_{0,\mathbf{i}}^{(0)}(\Lambda) = \frac{-1}{2\pi^2 M_N} \frac{1}{\theta_0 \Lambda}.$$
 (14)

Perturbative higher orders are defined by formally expanding the full T matrices t_i as

$$t_{\mathbf{i}} = t_{\mathbf{i}}^{(0)} + t_{\mathbf{i}}^{(1)} + t_{\mathbf{i}}^{(2)} + \cdots,$$
 (15)

where $t_{\mathbf{i}}^{(0)}$ is defined by Eq. (12). The corrections $t_{\mathbf{i}}^{(n)}$ for n > 0 can conveniently be obtained by solving similar integral equations [25,27]. For the unitarity expansion, corrections from the finite scattering length enter at NLO *via* $C_{0,\mathbf{i}}^{(1)}$, yielding a separable potential $V_{2,\mathbf{i}}^{(1)}$ with the same form as Eq. (6). For $t_{\mathbf{i}}^{(1)}$, this gives rise to

$$t_{\mathbf{i}}^{(1)} = V_{2,\mathbf{i}}^{(1)} + V_{2,\mathbf{i}}^{(1)} G_0 t_{\mathbf{i}}^{(0)} + V_{2,\mathbf{i}}^{(0)} G_0 t_{\mathbf{i}}^{(1)}, \tag{16}$$

which, just like the LO equation, can be solved algebraically (see Ref. [25] for explicit details). From this procedure one obtains



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$$C_{0,\mathbf{i}}^{(1)}(\Lambda) = \frac{-2\pi^2 M_N}{a_{\mathbf{i}}} \left[C_{0,\mathbf{i}}^{(0)}(\Lambda) \right]^2. \tag{17}$$

Range corrections enter at NLO together with $C_{0,i}^{(1)}$. They are generated by contact interactions involving quadratic derivatives acting on the nucleon fields, included in the ellipses in Eq. (4). The corresponding potential can be written in momentum space as $C_{2,i}^{(1)}g(k^2)\left(k^2+k'^2\right)g(k'^2)$. By virtue of this still being a separable interaction, the corresponding version of Eq. (16) with

$$\langle \mathbf{k} | V_{2,\mathbf{i}}^{(1)} | \mathbf{k}' \rangle$$

$$= C_{0,\mathbf{i}}^{(1)} g(k^2) g(k'^2) + C_{2,\mathbf{i}}^{(1)} g(k^2) \left(k^2 + k'^2\right) g(k'^2) \quad (18)$$

can still be solved algebraically. Matching the result to the ERE (1) up to the quadratic term gives

$$C_{2,\mathbf{i}}^{(1)}(\Lambda) = \pi^2 M_N \left(\frac{r}{2} - \frac{1}{\theta_2 \Lambda}\right) \left[C_{0,\mathbf{i}}^{(0)}(\Lambda)\right]^2,$$
 (19a)

$$C_{0,\mathbf{i}}^{(1)}(\Lambda) = 4\pi^2 M_N \theta_2 \Lambda^3 C_{0,\mathbf{i}}^{(0)}(\Lambda) C_{2,\mathbf{i}}^{(1)}(\Lambda), \tag{19b}$$

with $\theta_2 = \theta_0/4$ for the Gaussian regulator used here. Going to higher orders is straightforward, proceeding in the same way *via* integral equations that can be solved algebraically, recursively using the solutions of previous orders [25,27]. At second order, the *T*-matrix correction is obtained from

$$t_{\mathbf{i}}^{(2)} = V_{2,\mathbf{i}}^{(2)} + V_{2,\mathbf{i}}^{(2)} G_0 t_{\mathbf{i}}^{(0)} + V_{2,\mathbf{i}}^{(1)} G_0 t_{\mathbf{i}}^{(1)} + V_{2,\mathbf{i}}^{(0)} G_0 t_{\mathbf{i}}^{(2)}.$$
(20)

The treatment of Coulomb contributions (which are neglected in this work), in particular the matching between perturbative and nonperturbative Coulomb regimes within Pionless EFT, is discussed in Refs. [25,26].

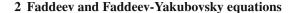
Leading order for A > 2 is not complete with only the $C_{0,i}^{(0)}$ interactions. It is a distinct feature of Pionless EFT, intimately related to the Efimov effect [14–16], that a three-nucleon interaction enters at LO. Naïvely it would be expected to contribute only much later in the power counting because the larger number of fields, according to NDA, implies more inverse powers of $M_{\rm hi}$ in the prefactor. Analogously to the two-body interactions, the potential induced by the term involving D_0 in Eq. (4) can be written in a separable form,

$$V_3^{(0)} = D_0^{(0)} |^3 H \rangle |\xi\rangle \langle \xi| \langle |^3 H|$$
 (21)

at LO, where $|{}^{3}\text{H}\rangle$ projects onto a J=T=1/2 three-nucleon state and the regulator $|\xi\rangle$ is defined, for Jacobi momenta $\mathbf{u}_{1}=\frac{1}{2}(\mathbf{k}_{1}-\mathbf{k}_{2})$ and $\mathbf{u}_{2}=\frac{2}{3}[\mathbf{k}_{3}-\frac{1}{2}(\mathbf{k}_{1}+\mathbf{k}_{2})]$, as

$$\langle \mathbf{u}_1 \mathbf{u}_2 | \xi \rangle = g \left(u_1^2 + \frac{3}{4} u_2^2 \right).$$
 (22)

The \mathbf{k}_i label the individual nucleon momenta. An NLO correction $V_3^{(1)}$ has the same form as Eq. (21), but involves the LEC $D_0^{(1)}$. Both $D_0^{(0)}$ and $D_0^{(1)}$ are determined by the triton binding energy and then enter in other calculations of $A \geq 3$ observables. These are described in the following.



This section gives an overview of the three- and four-body formalism, implementing a unified framework to solve, respectively, the Faddeev and Faddeev-Yakubovsky equations used to obtain the results presented in the following sections. The main aspects are explained in broad strokes, referring the reader to the references given for more background. Developments needed to calculate charge radii along with perturbative corrections are, however, elaborated on further, with key results explained in the main text and additional details provided in Appendices A and B.

The basis for a description of the three-nucleon system are Jacobi momenta

$$\mathbf{u}_1 = \frac{1}{2}(\mathbf{k}_1 - \mathbf{k}_2),\tag{23a}$$

$$\mathbf{u}_2 = \frac{2}{3} \left[\mathbf{k}_3 - \frac{1}{2} (\mathbf{k}_1 + \mathbf{k}_2) \right], \tag{23b}$$

where the \mathbf{k}_i label the individual nucleon momenta, conjugate to position vectors \mathbf{x}_i . Projecting these momenta onto partial waves yields states $|u_1u_2; s\rangle$, where

$$|s\rangle = |\left(l_2\left((l_1s_1)j_1\frac{1}{2}\right)s_2\right)J;\left(t_1\frac{1}{2}\right)T\rangle \tag{24}$$

collects angular momentum, spin, and isospin quantum numbers. They are coupled such that $(l_1s_1)j_1$ and t_1 describe the two-nucleon subsystem, whereas l_2 denotes the orbital angular momentum associated with the Jacobi momentum u_2 and s_2 is an intermediate quantum number. For the trinucleon bound states, the total spin and isospin are J=T=1/2. These states are determined by solving the Faddeev equation [28]

$$|\psi^{(0)}\rangle = G_0 t^{(0)} P |\psi^{(0)}\rangle + \frac{1}{3} (G_0 + G_0 t^{(0)} G_0) V_3^{(0)} (1+P) |\psi^{(0)}\rangle, \qquad (25)$$

where $|\psi^{(0)}\rangle = |\psi^{(0)}_{(12)3}\rangle$ is one of three equivalent two-body Faddeev components. As already done in the discussion of the two-body sector, explicit superscripts "(0)" are used to denote leading-order quantities. Alternatively, one can incorporate the three-body interaction $V_3^{(0)}$ by writing [29]

$$|\tilde{\psi}^{(0)}\rangle = G_0 t^{(0)} P |\tilde{\psi}^{(0)}\rangle + G_0 t^{(0)} |\psi_3^{(0)}\rangle,$$
 (26a)

$$|\psi_3^{(0)}\rangle = G_0 t_3^{(0)} (1+P) |\tilde{\psi}^{(0)}\rangle,$$
 (26b)

where $|\psi_3^{(0)}\rangle$ is an auxiliary amplitude, and

$$t_3^{(0)} = V_3^{(0)} + V_3^{(0)} G_0 t_3^{(0)}. (27)$$

The tilde is used to distinguish $|\tilde{\psi}^{(0)}\rangle$ from $|\psi^{(0)}\rangle$. In either form of the Faddeev equations, G_0 denotes the free three-body Green's function and $P=P_{12}P_{23}+P_{13}P_{23}$ generates the non-explicit components through permutations. $t^{(0)}$ collectively denotes the two-body T-matrices $t_{\bf i}^{(0)}$. Note that



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 $|\psi_3^{(0)}\rangle$ can be eliminated by inserting Eq. (26b) into Eq. (26a), yielding an equation of the form

$$|\psi^{(0)}\rangle = K^{(0)}|\psi^{(0)}\rangle \tag{28}$$

with

$$K^{(0)} = G_0 t^{(0)} P + G_0 t^{(0)} G_0 t_3^{(0)} (1+P),$$
 (29)

and alternatively a similar kernel can be obtained from Eq. (25) in terms of $V_3^{(0)}$. Either form of the Faddeev equations is solved by representing it within the space of states $|u_1u_2;s\rangle$, discretizing the momenta $u_{1,2}$ on a quadrature mesh. The binding energy is determined by varying the energy E, entering as an argument to both G_0 and $t^{(0)}$ until the kernel K has a unit eigenvalue. At that energy, $E=-B_0$, $|\psi^{(0)}\rangle$ can then be determined as the corresponding eigenstate and finally one constructs the full wavefunction as

$$|\Psi^{(0)}\rangle = (1+P)|\psi^{(0)}\rangle + |\psi_3^{(0)}\rangle.$$
 (30)

If one starts from Eq. (25), there is no explicit three-body Faddeev component and one simply has $|\Psi^{(0)}\rangle = (1+P)|\psi^{(0)}\rangle$.

Fundamentally, the permutation operator P leads to a coupling of different partial waves (see Ref. [30] for an excellent pedagogical discussion of both this and the Faddeev equations in general), and for the construction of the full wavefunction (30) it is important to include higher partial waves: the proper antisymmetry of $|\Psi^{(0)}\rangle$ is only recovered as more and more states are included, which means that in principle all observables calculated from $|\Psi^{(0)}\rangle$ have to be checked for convergence.

The full wavefunctions $|\Psi^{(0)}\rangle$ are used to calculate both perturbative shifts for the binding energy,

$$B_1 = \langle \Psi^{(0)} | V^{(1)} | \Psi^{(0)} \rangle, \tag{31}$$

as well as the radius at LO (see Sect. 4). Note that $|\Psi^{(0)}\rangle$ is assumed here to be properly normalized,

$$\langle \Psi^{(0)} | \Psi^{(0)} \rangle = 1.$$
 (32)

In calculating the matrix elements in Eqs. (31) and (32) it is advantageous to exploit the antisymmetry of $|\Psi^{(0)}\rangle$ as much as possible because that will speed up convergence of results with respect to the number of partial-wave channels. For example, the two-body part of the full potential $V^{(1)}$ can be expressed through permutations in terms of only the potential between the pair of nucleons 1 and 2, and it holds that (1+P)(1+P)=3(1+P).

Note furthermore that Eq. (26), and similarly Eq. (25), can be significantly simplified by exploiting the fact that the twoand three-body interactions are separable and act only within *S* waves. As a result, it suffices to work with merely two coupled equations for the triton [29], and using the procedure described in the previous paragraph it is furthermore possible to eliminate all intermediate higher partial-wave components in the calculation of B_1 . These simplifications are used in the practical implementation to fit the three-body LECs $D_0^{(0)}$ and $D_0^{(1)}$.

Describing the four-nucleon system requires an additional Jacobi momentum

$$\mathbf{u}_3 = \frac{3}{4} \left[\mathbf{k}_4 - \frac{1}{3} (\mathbf{k}_1 + \mathbf{k}_2 + \mathbf{k}_3) \right],\tag{33}$$

as well as an alternative set of momenta $(\mathbf{v}_1, \mathbf{v}_2, \mathbf{v}_3)$, describing a 2+2 cluster setup, *i.e.*, $\mathbf{v}_1 = \mathbf{u}_1$, \mathbf{v}_3 denotes the relative momentum in the (34) system, and \mathbf{v}_2 is defined as the relative momentum between the (12) and (34) subsystems.

Including the remaining quantum numbers, this leads to channel states

$$|a\rangle = |(l_2((l_1s_1)j_1\frac{1}{2})s_2)j_2, (l_3\frac{1}{2})j_3, (j_2j_3)J; ((t_1\frac{1}{2})t_2\frac{1}{2})T\rangle,$$
(34a)

$$|b\rangle = |(\lambda_2(\lambda_1\sigma_1)\iota_1)\iota_2, (\lambda_3\sigma_3)\iota_3, (\iota_2\iota_3)J; (\tau_2\tau_3)T\rangle.$$
(34b)

The $|a\rangle$ are a straightforward extension of three-nucleon states (24), including the angular momentum l_3 associated with \mathbf{u}_3 as well as spin and isospin $\frac{1}{2}$ for the fourth nucleon. For the b states, $(\lambda_1, \sigma_1, \tau_1)$ and $(\lambda_3, \sigma_3, \tau_3)$ are two-body quantum numbers for the (12) and (34) subsystems, respectively, whereas $\lambda_{1,2,3}$ are the angular momenta associated with $v_{1,2,3}$. For ⁴He the total spin and isospin are J=T=0.

Following Refs. [29,31], the Faddeev-Yakubovsky equations can be written as

$$|\psi_A^{(0)}\rangle = G_0 t^{(0)} P \left[(1 - P_{34}) |\psi_A^{(0)}\rangle + |\psi_B^{(0)}\rangle \right]$$

$$+ \frac{1}{3} (1 + G_0 t^{(0)}) G_0 V_3^{(0)} |\Psi^{(0)}\rangle$$
(35a)

$$|\psi_B^{(0)}\rangle = G_0 t^{(0)} \tilde{P} [(1 - P_{34})|\psi_A^{(0)}\rangle + |\psi_B^{(0)}\rangle],$$
 (35b)

corresponding to the decomposition

$$|\Psi^{(0)}\rangle = (1+P)\Big[(1-P_{34}-P_{34}P)|\psi_A^{(0)}\rangle + (1+\tilde{P})|\psi_B^{(0)}\rangle\Big]$$
(36)

of the full four-body wavefunction. The two distinct Fadde-ev-Yakubovsky components $|\psi_A\rangle$ and $|\psi_B\rangle$ correspond to, respectively, 3+1 and 2+2 cluster configurations of the four-body system, with the former naturally expressed in terms of the Jacobi momenta u_i and states $|a\rangle$, and the latter in terms of v_i and $|b\rangle$. Note that the same notation is used here as for the three-body case, and G_0 in Eqs. (35a, b) now represents the free four-body Green's function. In addition to the operator P already encountered in the three-body system, Eqs. (35a, b) include the further permutations P_{34} and $\tilde{P}=P_{13}P_{24}$ to ensure proper antisymmetry. In fact, the overall symmetry is determined by the sign in front of P_{34} : to study a bosonic system, one would use $(1+P_{34})$ acting on $|\psi_A^{(0)}\rangle$ to construct fully symmetric states.



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The structure of Eqs. (35a, b) can be made clearer by rewriting them in a generic matrix form,

$$\left(\mathbb{1} - \hat{K}^{(0)}\right) | \boldsymbol{\psi}^{(0)} \rangle = 0, \tag{37}$$

with $|\pmb{\psi}^{(0)}\rangle=(|\psi_A^{(0)}\rangle,|\psi_R^{(0)}\rangle)^T$ and the kernel

$$\hat{K}^{(0)} = G_0 t^{(0)} \hat{P} + \frac{1}{3} (G_0 + G_0 t^{(0)} G_0) V_3^{(0)} \hat{P}_3.$$
 (38)

In Eq. (38), G_0 and $t^{(0)}$ are understood to be diagonal matrices, and the permutation operators are collected in

$$\hat{P} = \text{diag}(P, \tilde{P}) \otimes \begin{pmatrix} (1 - P_{34}) & 1 \\ (1 - P_{34}) & 1 \end{pmatrix},$$
 (39a)

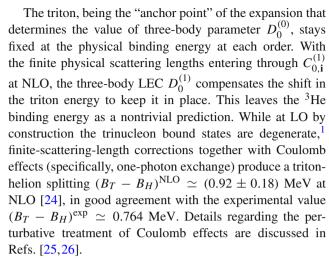
$$\hat{P}_{3} = \begin{pmatrix} (1+P)(1-P_{34}-P_{34}P) & (1+P)(1+\tilde{P}) \\ 0 & 0 \end{pmatrix},$$
(39b)

where "\ointilde{\text{"}}" represents matrix multiplication. From this form the structural analogy to the three-body Faddeev case becomes obvious.

Just like for the Faddeev equations, Eqs. (35a, b) are solved by projecting onto states $|u_1u_2u_3;a\rangle$, $|v_1v_2v_3;b\rangle$ [31], discretizing all momenta on a grid, and looking for a unit eigenvalue of the resulting kernel matrix as a function of the energy. However, the set of coupled equations does not naturally truncate even if all interactions are pure S wave. This means that already for a determination of the binding energies it is necessary to truncate the sums in Eq. (34) (by choosing all total angular momenta j_i and ι_i less than some j_{max}) and study the numerical convergence of results as j_{max} is increased.

3 Binding energies of light nuclei

By construction, the unitarity expansion renders the deuteron a zero-energy bound state at leading order. Since the expansion is set up in powers of the inverse scattering lengths, it corresponds to the zero-range binding momentum $\kappa_t = 1/a_t$ in the ${}^{3}S_{1}$ channel. As demonstrated explicitly in Ref. [25] by using the perturbative formalism discussed in Sect. 1, this implies that the deuteron remains at zero energy at NLO and only moves to $1/(M_N a_t^2)$ in an N²LO calculation. This is so for both the pure expansion in $1/a_t$, neglecting range corrections, but interestingly also for the full unitarity expansion that includes, via $C_{2,\mathbf{i}}^{(1)}$, range corrections starting at NLO. This is so because the unitarity LO shifts all corrections that mix ERE parameters to a higher order compared to where they would be with a finite scattering length at LO. Overall, the dominant source of uncertainty for the deuteron energy comes from the $1/(Q_2a_t)$ expansion, which still amounts to a 50% effect at N²LO. Conservatively taking the experimental binding energy as reference value for the uncertainty estimate yields $B_d^{\rm N^2LO} = 1.41 \pm 1.12$ MeV.



The ⁴He nucleus provides a more serious test of the unitarity expansion. Since $Q_AR \sim 0.8$ for ⁴He, it is the standard Pionless EFT part of the unitarity scheme which naïvely one might doubt to work, while the pure unitarity expansion, \Re/Q_A , should indeed work better with increasing Q_A . Figure 1 shows the ⁴He binding energy as a function of the momentum cutoff Λ . The observed convergent behavior as Λ increases indicates that the EFT calculation is properly renormalized, as established originally in Refs. [29,32,33]. Results for the standard pionless LO, given by the (blue) diamonds in Fig. 1, are consistent with this earlier work.

While any Λ above the breakdown scale (of order M_π) is a valid cutoff choice in principle, polynomials in $1/\Lambda$ are fitted to the points in Fig. 1 to quantitatively assess the convergence and conveniently extrapolate $\Lambda \to \infty$. This procedure gives $B_\alpha = 39(12)$ MeV in the unitarity limit, with the uncertainty estimated as $\mathcal{O}(r_{s,t}/a_{s,t}) \simeq 30\%$ based on the expectation that range effects 3S_1 are the dominant correction. Including the finite-scattering lengths as NLO corrections gives the (green) circles in Fig. 1, very close to the standard pionless LO, indicating that the $1/(Q_4a_{s,t})$ expansion appears to converge remarkably well up to this order, and indeed the extrapolated result 30(9) MeV comes out very near the standard pionless LO value of 31(9) MeV.²

It should be stressed that an NLO including only finitea corrections is incomplete: in the full unitarity expansion, range corrections and Coulomb effects enter at the same time.



In Table 1, it is indicated that the ${}^{3}\text{H}$ energy is fixed at 8.48 MeV as renormalization condition to determine $D_{0}^{(0)}$, while the ${}^{3}\text{He}$ energy is then a prediction with 30% uncertainty estimate at LO.

² Note that the NLO data points at finite cutoffs shown in Fig. 1 differ slightly from the results shown in Refs. [24,34]. This is due to a small error that has been fixed in the numerical implementation. Incidentally, the NLO result in the limit $\Lambda \to \infty$ is almost unaffected by this, with the value merely changing from the 29.5 MeV reported in Ref. [24] to 30.2 MeV. The difference is negligible compared to the \simeq 9 MeV uncertainty estimated at this order. Therefore, all conclusions of the previous work remain unchanged.

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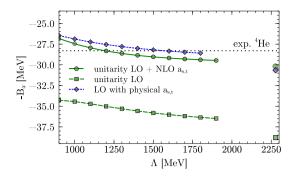


Fig. 1 ⁴He binding energy as function of the Gaussian cutoff parameter Λ . The (blue) diamonds and (green) squares show, respectively, the results for standard Pionless EFT and the unitarity expansion at LO. Inclusion of first-order corrections in $1/a_{s,t}$ (i.e., an incomplete NLO that neglects range and Coulomb effects) gives the (green) circles. The closeness to the standard leading order demonstrates how well this part of the expansion convergences. Large symbols on the right edge indicate results for an extrapolation $\Lambda \to \infty$ (see text)

While the latter are expected to be small given that ⁴He is rather deeply bound, it turns out that the inclusion of range effects actually has a profound consequence: in Ref. [35] it is shown that a four-body interaction is required to renormalize the universal four-boson system once range corrections are included at NLO, and universality implies that this conclusion carries over directly to ⁴He in Pionless EFT. The implication is that a four-nucleon input datum is required to fix the unknown four-body parameter at NLO, and this is most naturally taken to be ⁴He ground-state energy. Other properties, such as the ground-state radius (discussed below) or the position of the 0+ excited state, will remain predictions at NLO – unless it turns out that additional many-body forces are required for a renormalized NLO calculation of these observables. While it may seem unlikely and is certainly not to be expected based on NDA, such a possibility cannot a priori be excluded: each calculation needs to be carefully checked if it is properly renormalized.

The rapid convergence of the pure unitarity expansion persists off the physical point. Figure 2 shows the correlation between 3N and 4N binding energies, known as the Tjon line [36]. Its existence is explained by the three-body parameter largely governing the physics of the system [33]. It is seen that the result starting from unitarity is shifted very close to having the exact scattering lengths at LO over a significant range or triton energies. This observation provides further evidence that the unitarity expansion converges well and that the results found at the physical point are not merely accidental.

The unitarity expansion for ground-state energies up to A = 4 is summarized in Table 1. Observables fixed as input data at a given order are shown as underlined text. This in particular includes the ³He binding energy at N²LO since according to Refs. [25,26] an isospin-breaking three-body

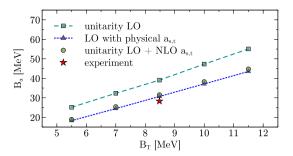


Fig. 2 Tjon line: correlation between the ⁴He and ³H binding energies. (Blue) dotted curve: standard pionless LO result; (green) dashed upper curve: unitarity limit at LO. Additional points nearly on top of the blue curve: inverse scattering lengths added in first-order perturbation theory. Star: experimental point

Table 1 Unitarity expansion convergence pattern. Underlined values indicate energies which are used as input values to determine three-body LECs. An asterisk superscript indicates an incomplete NLO result which only includes the finite-scattering length but no contributions from effective ranges or electromagnetic interactions

State	$E_B^{ m LO}/{ m MeV}$	$E_B^{ m NLO}/{ m MeV}$	$E_B^{ m N^2LO}/{ m MeV}$	$E_B^{\text{exp.}}/\text{MeV}$
² H	0	0	1.41 ± 1.12	2.22
^{3}H	8.48	<u>8.48</u>	<u>8.48</u>	8.48
³ He	8.5 ± 2.5	7.6 ± 0.2	<u>7.72</u>	7.72
⁴ He	39 ± 12	$30 \pm 9^*$		28.3

force is required once perturbation theory mixes Coulomb effects and range corrections.³ Taking into account the findings of Ref. [35], the ⁴He binding energy should be underlined for a complete NLO as well.

4 Charge radii and form factors

It has so far been established that the unitarity expansion describes well the ground-state energies of light nuclei. While certainly impressive given how simple the LO of the expansion is, it is still merely a first step towards showing that the scheme comprehensively captures the properties of light nuclei.

Further insight can be gained by considering charge radii of the A=3,4 systems as well. Within the momentum-space framework employed in this work, this is achieved by calculating charge form factors

$$F_C(q^2) = \langle \Psi | \hat{\rho}(\mathbf{q}) | \Psi \rangle \tag{40}$$

³ Note that if Coulomb effects are included nonperturbatively already at LO, which is not necessary for light nuclei, an isospin-breaking three-body force enters already at NLO [37].



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for the ³H and ⁴He ground states, from which one obtains point charge radii as

$$\langle r_0^2 \rangle = -\frac{1}{6} \frac{\mathrm{d}}{\mathrm{d}(q^2)} F_C(q^2) \Big|_{q^2 = 0} , \langle r_0 \rangle \equiv \sqrt{\langle r_0^2 \rangle}. \tag{41}$$

In Eq. (40), the total charge operator $\hat{\rho} \equiv J_0$, *i.e.*, the zero component of the electric current J_{μ} , is given by the sum of the individual nucleon contributions,

$$\hat{\rho} = \sum_{i=1}^{A} \hat{\rho}_i. \tag{42}$$

The (anti-)symmetry of the wavefunction makes it possible to replace this sum by $A\hat{\rho}_i$ for any fixed i. A particularly convenient choice for three nucleons is i=3 because it holds that

$$\mathbf{x}_3 = \mathbf{R}^{(3)} + \frac{2}{3}\mathbf{r}_2,\tag{43}$$

where \mathbf{r}_2 is the relative distance conjugate to \mathbf{u}_2 and $\mathbf{R}^{(3)}$ is the overall center-of-mass coordinate. With this choice, the momentum-space expression for the current operator involves a momentum transfer only onto the Jacobi momentum \mathbf{u}_2 . Likewise, for four nucleons a good choice is i=4 because with

$$\mathbf{x}_4 = \mathbf{R}^{(4)} + \frac{3}{4} \mathbf{r}_3 \tag{44}$$

one obtains a momentum transfer only onto \mathbf{u}_3 .

To use $\hat{\rho}$ within the Faddeev-Yakubovsky framework, it is necessary to represent it within the appropriate partial-wave basis. Since $\hat{\rho}$ does not depend on spin, two-body matrix elements of $\hat{\rho}$ are given by

$$\langle u; (ls)jm|\hat{\rho}(\mathbf{q})|u'; (l's')j'm'\rangle$$

$$= \delta_{jj'}\delta_{mm'}\langle u; l||\hat{\rho}(\mathbf{q})||u'; l\rangle, \tag{45}$$

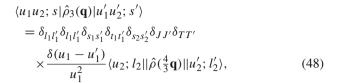
where the reduced matrix element on the right-hand side is given by

$$\langle u; lm | \hat{\rho}(\mathbf{q}) | u'; l'm' \rangle = \delta_{ll'} \delta_{mm'} \frac{1}{2} \sum_{k} \sqrt{\binom{2l}{2k}} C_{k0,(l-k)0}^{l0} \times \int_{-1}^{1} dx \, P_{k}(x) \frac{\delta(u' - \iota(u, q, x))}{u'^{2}} \frac{u^{l-k} \left(-\frac{1}{2}q\right)^{k}}{\iota(u, q, x)^{l}}, \quad (46)$$

without the $\delta_{mm'}$, where

$$\iota(p, q, x) = \sqrt{p^2 - pqx + q^2/4}. (47)$$

A detailed derivation of Eq. (46) is provided in Appendix A. Embedding $\hat{\rho}$ into the three- and four-body bases merely leads to additional Dirac and Kronecker deltas, as well as to kinematic prefactors multiplying the momentum transfer q which can be read off from Eqs. (43) and (44):



$$\langle u_{1}u_{2}u_{3}; a|\hat{\rho}_{4}(\mathbf{q})|u'_{1}u'_{2}u'_{3}; a'\rangle$$

$$= \delta_{l_{1}l'_{1}}\delta_{s_{1}s'_{1}}\delta_{j_{1}j'_{1}}\delta_{t_{1}l'_{1}}\delta_{s_{2}s'_{2}}\delta_{j_{2}j'_{2}}\delta_{t_{2}t'_{2}}\delta_{JJ'}\delta_{TT'}$$

$$\frac{\delta(u_{1} - u'_{1})}{u_{1}^{2}}\frac{\delta(u_{2} - u'_{2})}{u_{2}^{2}}\langle u_{3}; l_{3}||\hat{\rho}(\frac{3}{2}\mathbf{q})||u'_{3}; l'_{3}\rangle. \tag{49}$$

At leading order in the unitarity expansion, the form factor is given by

$$F_C^{(0)}(q^2) = \langle \Psi^{(0)} | \hat{\rho}(\mathbf{q}) | \Psi^{(0)} \rangle, \tag{50}$$

and analogously Eq. (41), with added superscripts "(0)," yields the LO point charge radii. At NLO, the correction to the form-factor is⁴

$$F_C^{(1)}(q^2) = 2 \langle \Psi^{(1)} | \hat{\rho}(\mathbf{q}) | \Psi^{(0)} \rangle, \tag{51}$$

where the factor 2 follows from symmetry. Perturbatively expanding Eq. (41) gives

$$\langle r_0 \rangle^{(1)} = \frac{1}{2} \frac{\langle r_0^2 \rangle^{(1)}}{\langle r_0 \rangle^{(0)}},$$
 (52)

with $\langle r_0^2 \rangle^{(1)}$ calculated from the slope of $F_C^{(1)}(q^2)$ at $q^2=0$. Evaluating the perturbative radius shifts defined above requires the NLO correction to the wavefunctions that enter in Eq. (51). Is is possible to obtain these for three- and four-body systems from inhomogeneous versions of the Faddeev and Faddeev-Yakubovsky equations, respectively. As is derived in Appendix B, for three particles one has $|\Psi_1\rangle = (1+P)|\psi_1\rangle$, with the NLO Faddeev component $|\psi_1\rangle$ defined as a solution of 5

$$\left[1 - G_0 t^{(0)} P - (G_0 + G_0 t^{(0)} G_0) V_3^{(0)}\right] |\psi_1\rangle
= (G_0 + G_0 t^{(0)} G_0) \left[V_2^{(1)} (1 + P) + V_3^{(1)} + B_1\right] |\psi_0\rangle.$$
(53)

In Eq. (53), G_0 and $t^{(0)}$ are understood to be evaluated at the LO binding energy, $E = -B_0$. Similarly, for four nucleons NLO Faddeev-Yakubovsky equations can be written as

$$\left(\mathbb{1} - \hat{K}^{(0)}\right)|\boldsymbol{\psi}^{(1)}\rangle = \hat{K}^{(1)}|\boldsymbol{\psi}^{(0)}\rangle,\tag{54}$$

with the kernel $\hat{K}^{(0)}$ as defined in Eq. (38) and



⁴ In principle, the current operator should also be expanded perturbatively, $\hat{\rho} = \hat{\rho}^{(0)} + \hat{\rho}^{(1)} + \cdots$, but there is no separate NLO contribution in the pure 1/a expansion considered here.

⁵ Note that Eq. (53) makes explicit use of the fact that the three-body force considered here is symmetric under all permutations, such that $V_3^{(n)}(1+P) = 3V_3^{(n)}$.

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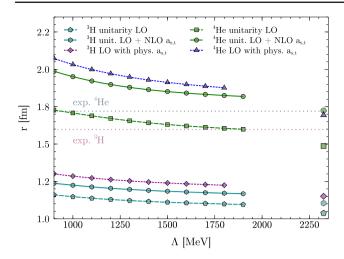


Fig. 3 Point-charge radii for ${}^3\mathrm{H}$ and ${}^4\mathrm{He}$ as function of the Gaussian cutoff parameter Λ . The (purple) diamonds and (cyan) pentagons show, respectively, ${}^3\mathrm{H}$ results for standard Pionless EFT and the unitarity expansion at LO. The (cyan) hexagons are obtained by perturbatively including 1/* correction on top of the unitarity LO. For ${}^4\mathrm{He}$, results are shown in the upper part, with symbols as in Fig. 1. Large symbols on the right edge indicate results for an extrapolation $\Lambda \to \infty$ (see text)

$$\hat{K}^{(1)} = B_1(G_0 + G_0 t^{(0)} G_0) + G_0 t^{(1)} \hat{P} + G_0 t^{(1)} G_0 V_3^{(0)} \hat{P}_3 + (G_0 + G_0 t^{(0)} G_0) V_3^{(1)} \hat{P}_3.$$
(55)

Note that as explained in Appendix B special care has to be taken when solving Eqs. (53) and (54) to account for the fact that the operators on the left-hand sides are singular at $E=-B_0$. From the components $|\psi^{(1)}\rangle=(|\psi_A^{(1)}\rangle,|\psi_B^{(1)}\rangle)^T$ the full NLO correction $|\Psi^{(1)}\rangle$ is obtained analogously to Eq. (36). For practical calculations Eq. (54) is simplified based on the fact that all interactions are chosen to be separable. This can be achieved with the same factorization as used in Ref. [29] at LO.

As for the ⁴He energy discussed in the previous section, the focus here is on the 1/a part of the unitarity expansion while the inclusion of range corrections is postponed to future work. Results for the ground-state radii of both ³H and ⁴He are shown in Fig. 3 as a function of the UV cutoff Λ . Convergence as Λ increases is evident from the plot, and just like it was done for the binding energies, polynomials in $1/\Lambda$ are fitted to the data points in order to extrapolate $\Lambda \to \infty$. The horizontal lines in Fig. 3 show the experimental values of the point charge radii, which, following Ref. [38], are defined as

$$\langle r_0^2 \rangle_{^3H} = \langle r^2 \rangle_{^3H} - \langle r^2 \rangle_p - 2\langle r^2 \rangle_n \tag{56}$$

for the triton, and

$$\langle r_0^2 \rangle_{^4\text{He}} = \langle r^2 \rangle_{^4\text{He}} - 2\langle r^2 \rangle_p - 2\langle r^2 \rangle_n \tag{57}$$

for ⁴He. That is, contributions from the root-mean-square radii of the individual nucleons are subtracted from the experimental nuclear charge radii.

Using experimental values from Ref. [39] for the quantities appearing on the right-hand sides of the above definitions one obtains, with error bars negligible compared to those of the present theoretical calculation, $\langle r_0^2 \rangle_{^{3}\text{H}}^{\text{exp}} = 1.59 \text{ fm}$ and $\langle r_0^2 \rangle_{^{4}\text{He}}^{\text{exp}} = 1.72 \text{ fm}$.

The lower part of Fig. 3 shows results for the triton. In the limit $\Lambda \to \infty$, indicated as points on the right border of the plot, the ³He radius comes out as $\langle r_0 \rangle_{^3\text{H}}^{(0)} = 1.15(35)$ fm for the standard pionless LO, and $\langle r_0 \rangle_{^3\text{H}}^{(0)} = 1.04(31)$ fm at unitarity. Perturbative corrections shift the unitarity LO result more than half way towards the value obtained for physical scattering lengths at leading order,

$$\langle r_0 \rangle_{3H}^{(0)} + \langle r_0 \rangle_{3H}^{(1)} = 1.10(33) \text{ fm},$$
 (58)

indicating that the unitarity expansion works well for this observable. This is in line with the results of Ref. [40], where good convergence is found for a perturbative expansion of 3 H and 3 He radii around an SU(4) symmetric leading order (of which the unitarity limit is a special case). The result obtained for standard pionless LO is furthermore in excellent agreement with the calculation of Ref. [38], at unitarity the radius satisfies well the universal relation [38,40,41]

$$M_N B_{^3\text{H}} \langle r_0^2 \rangle_{^3\text{H}} = (1 + s_0^2)/9 \approx 0.224.$$
 (59)

As done for binding energies it is assumed here that the Q_AR part dominates the overall expansion, yielding a 30% uncertainty both at LO as well as NLO. Indeed, from Ref. [38] it is known that range corrections contribute significantly to the triton radius in Pionless EFT and shift the result close to the experimental value once they are included. This uncertainty assignment places the experimental value outside the error band of the unitarity LO result. Since it is purely based on omitted range corrections, this is not actually a reason for concern and merely indicates that to be yet more conservative one should consider adding the uncertainties from the Q_AR and \aleph/Q_A expansions coherently.

Results for ⁴He radius, shown in the upper part of Fig. 3, look equally good. In fact, consistent with what is found for the binding energy, the result obtained from a standard pionless LO calculation, $\langle r_0 \rangle_{^4\text{He}}^{(0)} = 1.69(51)$ fm, comes out surprisingly close to the experimental value. At unitarity the radius comes out smaller, $\langle r_0 \rangle_{^4\text{He}}^{(0)} = 1.49(45)$ fm, consistent with the observed overbinding in the unitarity limit. Inclusion of perturbative 1/a corrections shifts this value to

$$\langle r_0 \rangle_{^{4}\text{He}}^{(0)} + \langle r_0 \rangle_{^{4}\text{He}}^{(1)} = 1.73(52) \text{ fm},$$
 (60)

in excellent agreement with the standard pionless LO result and therefore providing yet more evidence for the good convergence properties of the unitarity expansion. In particular, the fact that convergence appears to be somewhat faster for ⁴He than for the triton is in good agreement with



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 $\aleph/Q_4 < \aleph/Q_3$ obtained from Eq. (5), and it therefore reinforces confidence in this estimate.

5 Summary and perspectives

Superficially, the unitarity expansion may seem like merely a minor departure from standard Pionless EFT. It is rather well known that Pionless EFT, unlike Chiral EFT, is the ideal EFT to describe few-nucleon systems at low energies since its expansion explicitly embraces implications from the scattering lengths being large, basing its power counting explicitly on this fact. Chiral EFT is limited at low energies by its simultaneous expansion in both momenta and around the chiral limit, with $M_{\pi} \neq 0$ parametrizing the distance from it. This combination yields a power counting for $Q \sim M_{\pi}$ which does not easily capture the physics of the regime $Q \ll M_{\pi}$. Notably, one-pion exchange only contributes to the NN scattering lengths through loop effects.

However, the unitarity expansion does in fact constitute a significant paradigm shift in the EFT-based description of light nuclei: it goes as far as saying that the details of the two-body sector, represented by the experimental values of the scattering lengths, do not actually matter much to describe the gross properties of light nuclei. Instead, it fully embraces universality and uses the three-body sector as anchor point, constructing a leading order with just a single parameter and an exact manifestation of the Efimov effect. In this work it has been shown that the 1/a expansion of the unitarity scheme works well not only for binding energies of up to four nucleon systems, but that similarly good convergence is obtained for the ³H and ⁴He point charge radii as well. This finding solidifies the picture drawn in the introduction of this work, placing few-nucleon systems in a universal regime perturbatively close to the unitarity limit.

It is an important next step to include range corrections and Coulomb effects, thus considering the full unitarity scheme that pairs the expansions in \aleph/Q_A and $Q_A/M_{hi} = Q_A R$. So far, this has been investigated only for the ³H-³He energy splitting, where by construction range corrections cancel at NLO [24]. An isospin-breaking three-nucleon force is required once range corrections mix with Coulomb contributions at N²LO [25]. Range corrections are known to significantly contribute to the triton point charge radius [38]. For the ⁴He radius, the closeness to the experimental point already without range corrections found in this work leaves little room (about 0.01 fm in the central value) for a significant shift at full NLO, suggesting that several of the contributions neglected here might cancel. From Ref. [35] it is known that full NLO requires a four-nucleon force to be included along with range corrections, which is most conveniently fit to reproduce the ⁴He energy at the experimental point at NLO. It is conceivable that this fixing of the energy will maintain a good reproduction of the radius, just like it is observed for the standard Pionless LO result. Coulomb contributions should also be included in a complete NLO calculation, along with isospin breaking in the 1S_0 NN scattering lengths. This is expected to be a small effect for a bound state as deep as 4 He, but it is interesting to note that fitting the four-nucleon force to exactly reproduce the 4 He binding energy at NLO will inevitably absorb isospin-breaking contributions as well. Clearly a careful overall consideration of the LEC fitting procedure is called for in light of this to avoid possible overfitting of individual parameters. Bayesian methods stand ready as a powerful tool to address this [42,43].

Apart from such more technical issues, it is an exciting question how far into the nuclear chart the unitarity expansion can reach and what exactly its place is in the tower of nuclear EFTs. The observation that bosonic systems at unitarity exhibit saturation for large numbers of particles [44] and recent calculations of nuclear matter using interactions guided by unitarity [45] provide reason to be optimistic that universality, and in particular discrete scale invariance [46], is able to inform more than just few-nucleon calculations. On the other hand, Refs. [47,48] indicate that few-nucleon systems beyond A > 4 may not be bound in the unitarity limit. To further assess this situation one should investigate whether these states can be found as resonance (or virtual state) poles at unitarity, and if so, if these poles are perturbatively close to the situation in the real world.

In the bosonic sector, the promotion of many-body forces to lower orders than where they would be expected according to NDA is a fascinating consequence of universality [35], but it does impose practical limitations on many-body calculations. Beyond four nucleons the influence of Fermi statistics is expected to become important, which will most likely limit the promotion of A-nucleon forces with A > 4. However, at the same time one should wonder how much this might constrain the usefulness of universality in general, and at which point the Fermi momentum becomes a relevant scale for the description of nuclei. A calculation of, for example, n- α scattering within the unitarity expansion will be an important next test of the framework and help assess its exact place in the tower of nuclear EFTs.

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A Partial-wave decomposition of charge operators

A generic one-body charge operator between two-body states $|u;lm\rangle$ can be written as

$$\langle u; lm|\hat{\rho}(\mathbf{q})|u'; l'm'\rangle$$

$$= \int d^{3}p \int d^{3}p' \langle u; lm|\mathbf{p}\rangle\langle \mathbf{p}|\hat{\rho}(\mathbf{q})|\mathbf{p}'\rangle\langle \mathbf{p}'|u'; l'm'\rangle$$

$$= \int d^{3}p \int d^{3}p' Y_{lm}(\hat{p})$$

$$\times \frac{\delta(u-p)}{u^{2}} \delta^{(3)}(\mathbf{p}-\mathbf{p}'-\frac{1}{2}\mathbf{q})Y_{l'm'}^{*}(\hat{p}')\frac{\delta(u'-p')}{u'^{2}}$$

$$= \int dp p^{2} \int d\Omega_{p} Y_{lm}(\hat{p})\frac{\delta(u-p)}{u^{2}}$$

$$\times \frac{\delta(u'-|\mathbf{p}-\frac{1}{2}\mathbf{q}|)}{u'^{2}}Y_{l'm'}^{*}(\widehat{\mathbf{p}}-\frac{1}{2}\mathbf{q}). \tag{61}$$

The dependence on the angles of $\bf q$ can be isolated by using a procedure analogous to the one described in Ref. [30] for the permutation operators appearing in the Faddeev equations. It holds that

$$Y_{l'm'}^{*}(\widehat{\mathbf{p} - \frac{1}{2}\mathbf{q}}) = \sum_{\lambda'_{1} + \lambda'_{2} = l'} \frac{p^{\lambda'_{1}} \left(-\frac{1}{2}q\right)^{\lambda'_{2}}}{\left|\mathbf{p} - \frac{1}{2}\mathbf{q}\right|^{l'}} \times \sqrt{\frac{4\pi (2l' + 1)!}{(2\lambda'_{1} + 1)!(2\lambda'_{2} + 1)!}} \mathcal{Y}_{\lambda'_{1}\lambda'_{2}}^{l'm'*}(\hat{p}, \hat{q}),$$
(62)

and furthermore

$$\frac{\delta(u' - |\mathbf{p} - \frac{1}{2}\mathbf{q}|)}{u'^{2}} = 2\pi \sum_{k} \sqrt{\hat{k}} (-1)^{k}
\times \left[\int_{-1}^{1} dx \, P_{k}(x) \frac{\delta(u' - \sqrt{p^{2} - pqx + q^{2}/4})}{u'^{2}} \right]
\times \mathcal{Y}_{kk}^{00}(\hat{p}, \hat{q}),$$
(63)

where $\hat{k} = 2k + 1$.

In these expressions, $\mathcal{Y}_{l_1,l_2}^{LM}$ is used to denote two coupled spherical harmonics. The resulting product can be reduced:

$$\mathcal{Y}_{\lambda'_{1}\lambda'_{2}}^{l'm'*}(\hat{p},\hat{q})\mathcal{Y}_{kk}^{00}(\hat{p},\hat{q})
= \frac{1}{4\pi} \sqrt{\hat{k}\hat{\lambda}'_{1}\hat{\lambda}'_{2}}(-1)^{\lambda'_{1}+\lambda'_{2}+l'}
\times \sum_{f_{1},f_{2}} \begin{cases} f_{2} & f_{1} & l' \\ \lambda'_{1} & \lambda'_{2} & k \end{cases} C_{k0,\lambda'_{1}0}^{f_{1}0} C_{k0,\lambda'_{2}0}^{f_{2}0} \mathcal{Y}_{f_{1}f_{2}}^{l'm'*}(\hat{p},\hat{q}).$$
(64)

At this point, it is possible to perform the integral over \hat{p} in Eq. (61), yielding:

$$\int d\Omega_{p} Y_{lm}(\hat{p}) \mathcal{Y}_{f_{1}f_{2}}^{l'm'*}(\hat{p},\hat{q}) = \sum_{m_{2}} C_{lm,f_{2}m_{2}}^{l'm'} Y_{f_{2}m_{2}}^{*}(\hat{q}) \, \delta_{f_{1}l}.$$
(65)

Since $\hat{\rho}$ is a scalar operator that cannot connect $l' \neq l$, only $f_2 = m_2 = 0$ can contribute and $Y^*_{f_2m_2}(\hat{q})$ reduced to a factor $1/\sqrt{4\pi}$. This leads to a cascade of further simplifications:

$$C_{lm,00}^{l'm'} = \delta_{ll'}\delta_{mm'} \tag{66a}$$

$$\begin{cases}
0 & l & l' \\
\lambda_1' & \lambda_2' & k
\end{cases} = (-1)^{l' + \lambda_1' + \lambda_2'} \delta_{ll'} \delta_{\lambda_2' k} / \sqrt{\hat{l}\hat{k}}$$
(66b)

$$C_{k0,k0}^{00} = (-1)^k / \sqrt{\hat{k}}.$$
 (66c)

Putting everything together leads to Eq. (46) in the main text.

The next step is embedding the current into the three-nucleon system. To that end one decouples the states $|s\rangle$ defined in Eq. (24) to isolate the \mathbf{u}_2 part:

$$|u_{1}u_{2}; \left(l_{2}\left((l_{1}s_{1})j_{1}\frac{1}{2}\right)s_{2}\right)JM\rangle$$

$$= \sum_{m_{2},\sigma_{2}} \sum_{\mu_{1},\sigma} C_{l_{2}m_{2},s_{2}\sigma_{2}}^{j_{2}\mu_{2}} C_{j_{1}\mu_{1},\frac{1}{2}\sigma}^{s_{2}\sigma_{2}}$$

$$\times |u_{1}; (l_{1}s_{1})j_{1}\mu_{1}\rangle |u_{2}; l_{2}m_{2}\rangle |\frac{1}{2}\sigma\rangle. \tag{67}$$

Taking matrix elements, it is possible to exploit that $\hat{\rho}$ is diagonal, so one simply gets a number of Kronecker deltas from the reduction of the Clebsch-Gordan coefficients. In particular,

$$\langle (l_1 s_1) j_1 \mu_1 | (l_1' s_1') j_1' \mu_1' \rangle = \delta_{j_1 j_1'} \delta_{\mu_1 \mu_1'} \delta_{l_1 l_1'} \delta_{s_1 s_1'}. \tag{68}$$

For the isospin part, one finds

$$\langle \left(t_1 \frac{1}{2}\right) T \mid \left(t_1' \frac{1}{2}\right) T' \rangle = \delta_{t_1 t_1'} \delta_{T T'}, \tag{69}$$

such that overall one arrives at Eq. (48). Analogously, for the four-nucleon system, the first step is decoupling the \mathbf{u}_3 part from the states (34a). Omitting the intermediate quantum numbers for the (123) subsystem as well as the isospin part, one finds

$$|u_1u_2u_3; (j_2(l_3\frac{1}{2})j_3)JM\rangle = \sum_{u_2,u_3} \sum_{m_3\sigma_3} C_{j_2\mu_2,j_3\mu_3}^{JM}$$



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$$\times C_{l_{3}m_{3},\frac{1}{2}\sigma_{3}}^{j_{3}\mu_{3}}|u_{1}u_{2};j_{2}\mu_{2}\rangle|u_{3};l_{3}m_{3}\rangle|\frac{1}{2}\sigma_{3}\rangle,$$
(70)

leading to Eq. (49).

B Perturbative expansion of few-body bound states

This appendix discusses methods to obtain perturbative corrections for few-body wavefunctions. In the main text of this work, these corrections are used to calculate first-order shifts for three- and four-nucleon charge radii within the unitarity expansion, but having access to wavefunctions generally enables a variety of further calculations. For example, second-order corrections to binding energies can be obtained in a way that is numerically much simpler than the procedure used in Ref. [25] to extract such second-order shifts from off-shell T-matrix corrections. An abstract operator notation is used throughout this section, and conventions for suband superscripts differ from the usage in the main text in order to simplify the notation. After a general discussion that does not assume any fixed number of particles, concrete firstorder equations for three- and four-body states are derived in Sect. B.2. While explicit results are given for first-order perturbation theory, it is clear from the following discussion that higher-order equations can be derived analogously in a recursive fashion, much like it is done in Ref. [38] for a calculation of the triton charge radius up to next-to-next-to leading order. The derivations presented here are built on the concept of using inhomogeneous equations to calculate perturbative corrections, introduced in Ref. [38] specifically for three-body vertex functions, and previously in Ref. [27] for scattering calculations.

B.1 Generic discussion

In principle, one can base a perturbative corrections for bound-state wavefunctions on the corresponding expansion of the Lippmann-Schwinger equation

$$T = V + VG_0T, (71)$$

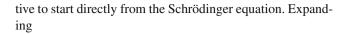
following the discussion in Sect. 1. Using the fact that as the energy approaches a bound-state at E=-B the T matrix has a pole, one has:

$$T(E) = \frac{|\mathcal{B}\rangle\langle\mathcal{B}|}{E+B} + \text{regular terms.}$$
 (72)

In this expression, the vertex function $|\mathcal{B}\rangle$ is related to the wavefunction $|\Psi\rangle$ via

$$|\psi\rangle = G_0(-B)|\mathcal{B}\rangle. \tag{73}$$

In order to develop a formalism that is connected to the bound-state Faddeev formalism, it is however more instruc-



$$V = V_0 + V_1 + \cdots, \tag{74a}$$

$$B = B_0 + B_1 + \cdots, \tag{74b}$$

$$|\Psi\rangle = |\Psi_0\rangle + |\Psi_1\rangle + \cdots. \tag{74c}$$

gives the well-known first-order equation

$$(H_0 + V_0)|\Psi_1\rangle + V_1|\Psi_0\rangle = -B_0|\Psi_1\rangle - B_1|\Psi_0\rangle. \tag{75}$$

In principle, an explicit solution is given by

$$|\Psi_1\rangle = \sum_{\alpha \in \mathcal{S} \setminus \{-B_0\}} \frac{|\Psi_\alpha\rangle \langle \Psi_\alpha | V_1 | \Psi_0 \rangle}{-B_0 - E_\alpha},\tag{76}$$

where $\mathcal S$ denotes the whole spectrum and the sum is generally a sum over discrete states plus an integral over the continuous spectrum (note that for simplicity it is assumed here that there are no degeneracies in the spectrum). Using Eq. (76) however requires a complete diagonalization of the leading-order Hamiltonian, which is not convenient for calculations based on Faddeev- and Faddeev-Yakubovsky equations that by construction only give access to specific individual states. Moreover, for the systems considered in this work almost all contributions in Eq. (76) would come from discretized continuum of scattering states, which is numerically challenging. It is therefore interesting and relevant to look for alternative ways of solving for $|\Psi_1\rangle$.

As a first step, one can rewrite Eq. (75) as

$$[-B_0 - H_0 - V_0] |\Psi_1\rangle = [V_1 + B_1] |\Psi_0\rangle. \tag{77}$$

and recognize from this that Eq. (76) involves the full leading-order Green's function in spectral representation, with the bound-state $|\Psi_0\rangle$ subtracted. This subtraction is crucial because

$$[-B_0 - H_0 - V_0] |\Psi_0\rangle = 0 (78)$$

implies the operator on the right-hand side of Eq. (77) is singular. It is the term $-B_1|\Psi_0\rangle$ on the right-hand side that generates this subtraction. With this insight it is possible to derive methods to deal with the problem. Using the definition of the free Green's function,

$$G_0(z) = (z - H_0)^{-1}, (79)$$

it is possible to further rewrite Eq. (77) as

$$[1 - G_0(z)V_0]|\Psi_1\rangle = G_0(z)[V_1 + B_1]|\Psi_0\rangle. \tag{80}$$

with z fixed at $-B_0$. For any $z \neq -B_0$, the kernel on the left-hand side of Eq. (80) is regular, and it can therefore be solved, after discretization, as a linear system of equations. Knowing from Eq. (76) that the solution is actually well defined in the limit $z \to -B_0$, numerical extrapolation can be used to obtain $|\Psi_1\rangle$.



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An alternative procedure is to consider a potential with the leading-order bound state removed. This can be achieved using the replacement [49–51]

$$V_0 \to V_0 + \lambda |\Psi_0\rangle \langle \Psi_0|$$
 (81)

in Eq. (80), with λ a large positive constant that moves the bound state far away from the low-energy spectrum of interest. This procedure is valid because the orthogonality condition for the first-order correction, $\langle \Psi_1 | \Psi_0 \rangle = 0$, implies that the modified equation is equivalent to original one. Numerically, it gives excellent agreement with the extrapolation method for a two-body test case.

B.2 Faddeev and Faddeev-Yakubovsky decomposition

For a three-body state one can write, in analogy to Eq. (30), $|\Psi_1\rangle=(1+P)|\psi_1\rangle$ and insert this decomposition into Eq. (77). Three-body interactions are neglected here to keep the discussion as transparent as possible. Application of the remaining two-body interaction can be simplified by exploiting the symmetry of the full wavefunction, which implies that

$$V|\Psi\rangle = (1+P)\tilde{V}|\Psi\rangle,\tag{82}$$

where \tilde{V} is the potential acting only on the specific pair of particles used to define the Faddeev component $|\psi\rangle$. This gives:

$$\left[-B_0 - H_0 - (1+P)\tilde{V}_0 \right] (1+P)|\psi_1\rangle
= \left[(1+P)\tilde{V}_1 + B_1 \right] (1+P)|\psi_0\rangle.$$
(83)

Since (1 + P) commutes with the free Hamiltonian H_0 , one can bring an application of (1 + P) to the left of each term in this equation. While (1 + P) can in general have zero eigenvalues, components that are in $\ker(1 + P)$ are clearly irrelevant for the description of the physical bound state and it is therefore possible to proceed with a simplified equation:

$$\begin{aligned}
& \left[G_0 (-B_0)^{-1} - \tilde{V}_0 \right] |\psi_1\rangle \\
&= \tilde{V}_0 P |\psi_1\rangle + \tilde{V}_1 (1+P) |\psi_0\rangle - B_1 |\psi_0\rangle,
\end{aligned} (84)$$

where $G_0(-B_0)^{-1} = -B_0 - H_0$. Acting with $G_0(-B_0)$ on both sides of this one can perform a partial inversion by using the Lippmann-Schwinger equation

$$\left[1 - G_0 \tilde{V}_0\right] G = G_0 \tag{85}$$

for the full two-body Green's function G associated with \tilde{V} . This can then be eliminated in favor of the corresponding t matrix by using

$$G\tilde{V} = G_0 t, \tag{86a}$$

$$G = G_0 + G_0 t G_0, (86b)$$

for $t = \tilde{V} + \tilde{V}G_0t$. Overall one arrives at:

$$[1 - G_0 t_0 P] |\psi_1\rangle$$

$$= (G_0 + G_0 t_0 G_0) \left[\tilde{V}_1 (1+P) + B_1 \right] |\psi_0\rangle$$
(87)

with $G_0 = G_0(-B_0)$. This is an inhomogeneous integral equation which involves exactly the same kernel as the Faddeev equation at leading order, whereas the terms on the left-hand side are straightforward to calculate from known quantities. As a final step one calculates $|\Psi_1\rangle$ from $|\psi_1\rangle$ in exactly the same way $|\Psi_0\rangle$ is calculated from $|\psi_0\rangle$. By keeping three-body forces in the derivation one analogously obtains Eq. (53) given in the main text.

It is useful to note that the same result can alternatively be derived directly from a perturbative expansion of the Faddeev equation

$$|\psi\rangle = G_0 t P |\psi\rangle,\tag{88}$$

setting $|\psi\rangle = |\psi_0\rangle + |\psi_1\rangle + \cdots$. To proceed it is important to carefully regard the energy arguments of the operators, noting that $B = B_0 + B_1 + \cdots$:

$$G_0(-B) = G_0(-B_0) + B_1 G_0(-B_0)^2 + \cdots,$$
(89a)
$$t(-B) = t_0(-B_0) - B_1 \frac{d}{dz} t_0(z) \Big|_{z=-B_0} + t_1(-B_0) + \cdots.$$
(89b)

The term involving the energy derivative in Eq. (89b) looks peculiar, but, assuming the potential to be energy independent, it can be shown by differentiating the Lippmann-Schwinger equation that

$$\frac{\mathrm{d}}{\mathrm{d}z}t_0(z) = -t_0(z)G_0(z)^2t_0(z). \tag{90}$$

Using this and making use of the leading-order Faddeev equation.

$$|\psi_0\rangle = G_0(-B_0)t_0(-B_0)P|\psi_0\rangle,$$
 (91)

to simplify some terms, one obtains

$$[1 - G_0 t_0 P] |\psi_1\rangle$$
= $B_1 (G_0 + G_0 t_0 G_0) |\psi_0\rangle + G_0 t_1 P |\psi_0\rangle$ (92)

All energy arguments are $-B_0$ and have therefore been omitted as well. Using again Eq. (91) as well as the Lippmann-Schwinger equation for t_1 , Eq. (16) in the main text, it possible to show that

$$G_0 t_1 P |\psi_0\rangle = (G_0 + G_0 t G_0) \tilde{V}_1 (1+P) |\psi_0\rangle,$$
 (93)

so that indeed Eq. (92) is equivalent to Eq. (87)

The previous result is both reassuring and very useful because it implies that in order to derive perturbative corrections for four-body states one can simply start from the Faddeev-Yakubovsky equations, avoiding a tedious detour through the Schrödinger equation.



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A convenient starting point for this calculation is generic matrix form of the Faddeev-Yakubovsky equations, analogous to Eq. (37):

$$\left(\mathbb{1} - \hat{K}\right)|\psi\rangle = 0,\tag{94}$$

with $|\psi\rangle = (|\psi_A\rangle, |\psi_B\rangle)^T$ and

$$\hat{K} = G_0 t \,\hat{P},\tag{95}$$

where three-body forces are again neglected for simplicity.

As before, one expands all quantities in this equation, carefully keeping track of the energy arguments. This gives

$$\left(\mathbb{1} - \hat{K}_0\right)|\psi_1\rangle = \hat{K}_1|\psi_0\rangle,\tag{96}$$

with

$$\hat{K}_0 = G_0 t_0 \hat{P},\tag{97a}$$

$$\hat{K}_1 = B_1(G_0 + G_0 t_0 G_0) + G_0 t_1 \hat{P}. \tag{97b}$$

All energy arguments in Eqs. (97a, b) are fixed now at $-B_0$ and have been omitted. Again the term involving t_1 can be rewritten in terms of \tilde{V}_1 as

$$G_0 t_1 \hat{P} | \psi_0 \rangle = (G_0 + G_0 t_0 G_0) \tilde{V}_1 (1 + \hat{P}) | \psi_0 \rangle. \tag{98}$$

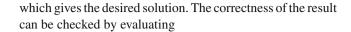
With three-body forces included one obtains the slightly more complicated \hat{K}_1 shown in Eq. (55) in the main text.

The fact that within the Faddeev(-Yakubovsky) formalism one has to calculate B_1 with an increasing number of partial-wave components included to check convergence of the numerical calculation renders the extrapolation method discussed in Sect. B.1 unstable beyond the two-body sector. At the same time, the projection method which modifies the potential to remove the leading-order bound state from the low-energy spectrum becomes impractical because for an A-body state Eq. (81) involves an A-body potential, which is expensive to handle computationally (at leading order) for A > 3. Fortunately, it is possible to employ the projection method directly at the level of Faddeev(-Yakubovsky) components and make the replacement⁶

$$K_0 \to K_0 + \lambda |\psi_0\rangle\langle\psi_0|$$
 (99)

in Eq. (87), where $K_0 = G_0t_0P$, and analogously for Eq. (96). While this ensures that the kernel becomes regular, the solution of the modified equation, denoted by $|\tilde{\Psi}_1\rangle$, will not in general directly provide the correct solution $|\Psi_1\rangle$ (obtained from $|\psi_1\rangle$ by adding the appropriate permutations). In particular, $|\tilde{\Psi}_1\rangle$ may not be orthogonal to $|\Psi_0\rangle$. It is however possible to simply project out the undesired component by setting

$$|\Psi_1\rangle = |\tilde{\Psi}_1\rangle - |\Psi_0\rangle \langle \tilde{\Psi}_1|\Psi_0\rangle, \tag{100}$$



$$\langle \Psi_1 | H_0 + V_0 | \Psi_1 \rangle + \langle \Psi_1 | V_1 | \Psi_0 \rangle = -B_0 \langle \Psi_1 | \Psi_1 \rangle, \tag{101}$$

which follows from Eq. (75) since B_0 is known from the leading-order calculation.

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 $^{^6}$ Alternatively, one can work directly with the singular K_0 and solve the first-order equation as a linear least-squares problems using a singular value decomposition (SVD).

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