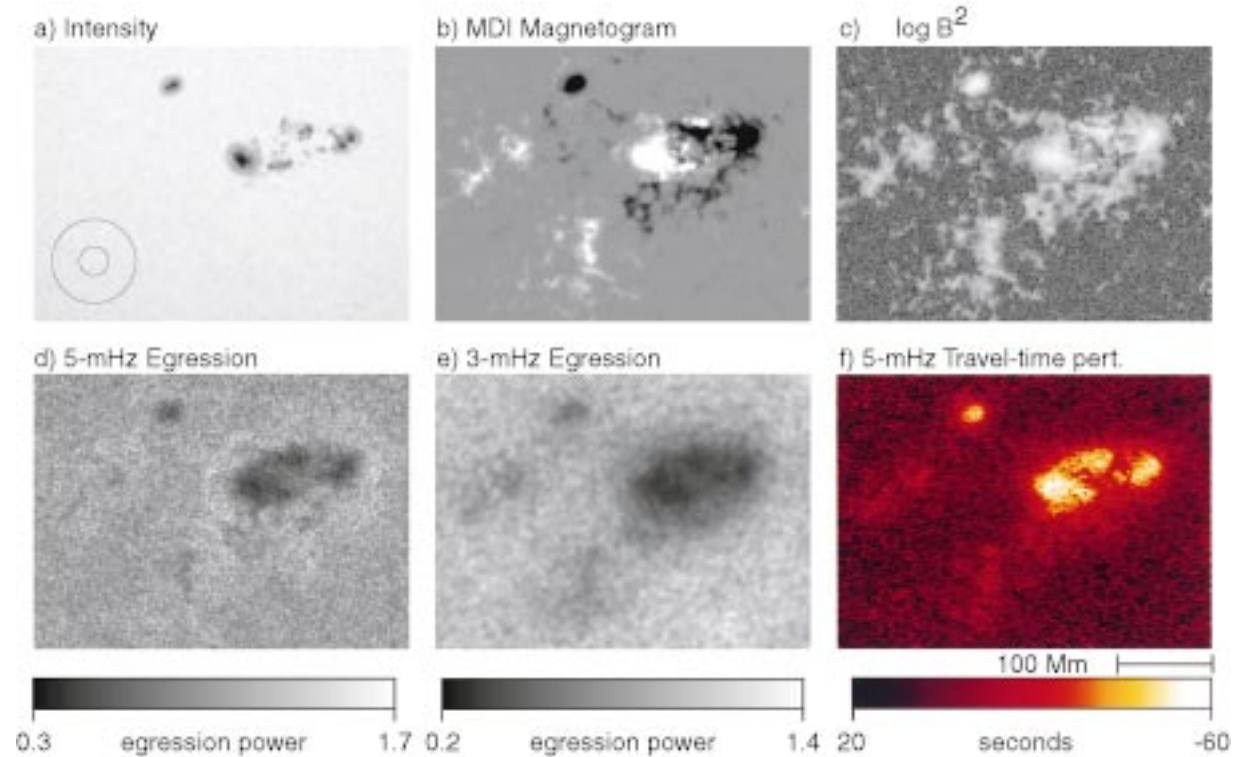**Erratum**

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Due to technical error which occurred during the production of this article, Figures 1, 4, and 5 on pp. 311, 316, and 317, respectively, were reproduced in black and white. Overleaf we present the figures in colour as they should have appeared.

Figure 1. (a) shows an MDI continuum intensity image of AR 8179 taken 1998 March 15.0. (b) is an average of three magnetograms over the subsequent 24-hr period, and (c) illustrates the logarithm of the square of the magnetogram. The areas enclosed by the curves in (b) indicate regions discussed in the text. The greyscale used in (b) saturates below −300 G and above +300 G. The greyscale used in (c) spans a range of magnetic flux density between 0.3 and 1600 G. (d) and (e) show egression power maps of the region, at 5 and 3 mHz, respectively. The greyscales indicate egression power, normalized to unity for the quiet Sun. (f) shows a map of 5 mHz sound travel-time perturbations, with respect to the quiet Sun. The annulus in (a) shows the dimensions of the pupil of both ingression and egression computations.
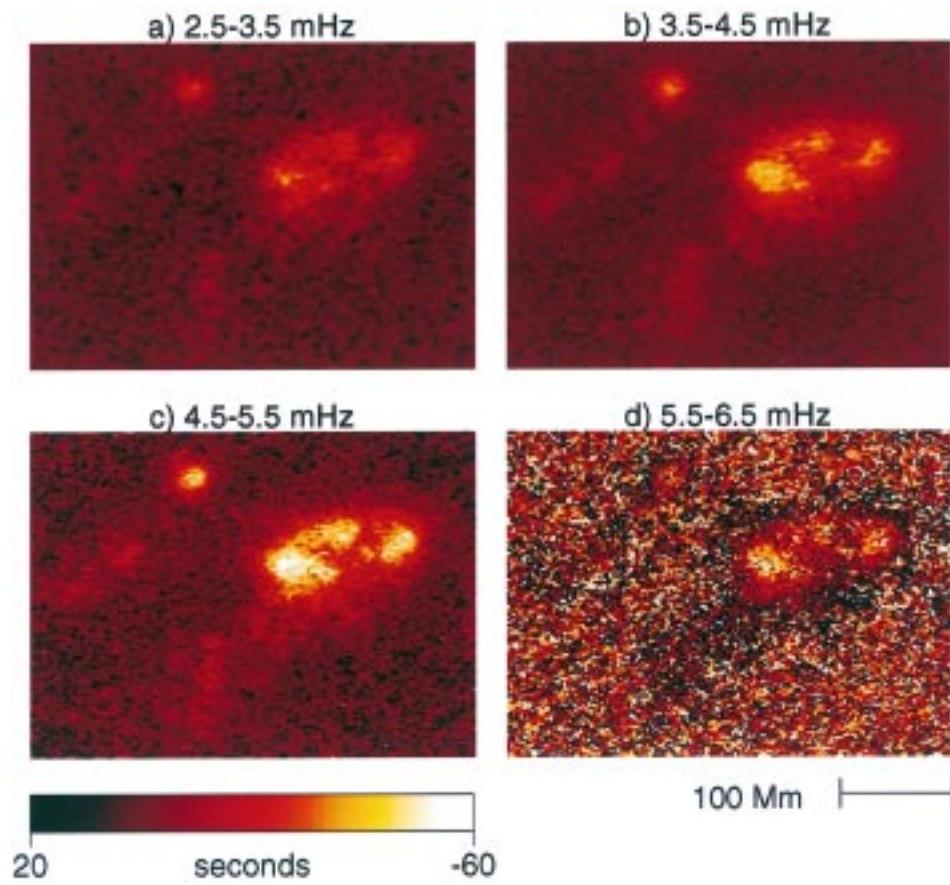
Figure 4. Maps of the sound travel-time perturbations computed over 1 mHz bandpasses centered at frequencies of 3, 4, 5, and 6 mHz.
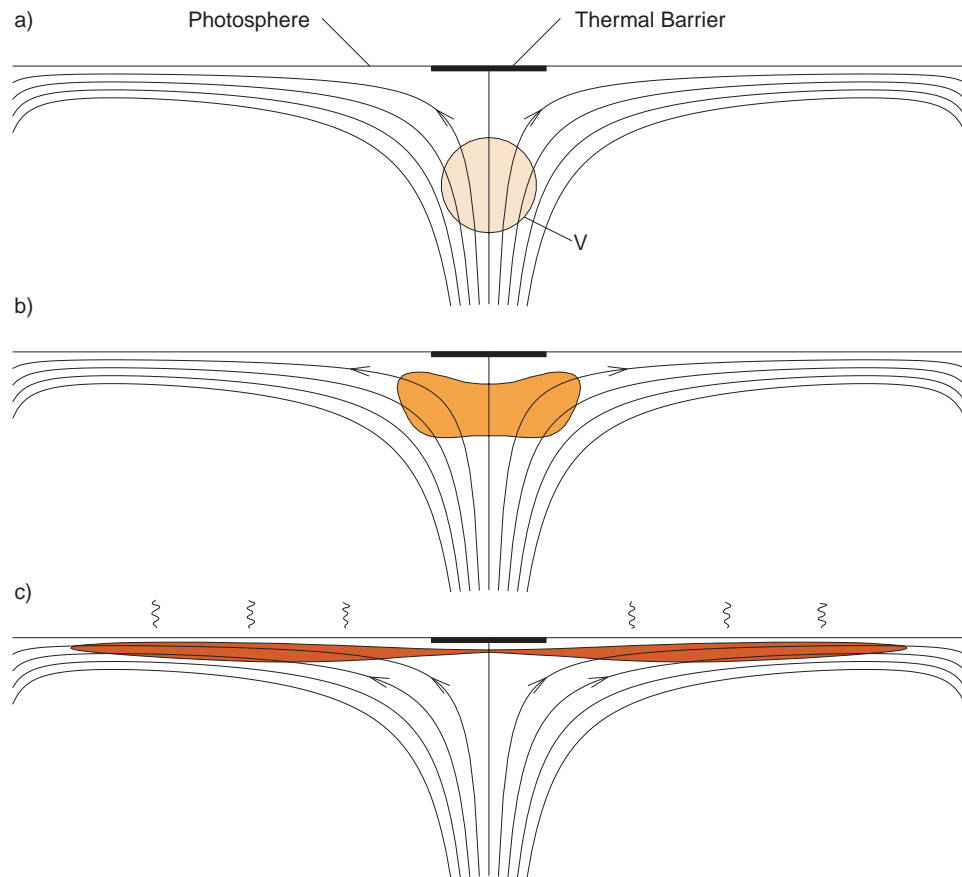
Figure 5. A schematic representation of a convective eddy driven by the local blockage of convective transport by a thermal barrier at the surface (solid rectangle). (a)–(c) show the temporal evolution of a thermal excess indicted initially by the volume *V* in (a).