**Research Article** 

# Investigation on corrosion morphology and products of ancient tin amalgam mirrors by AFM, SEM–EDS and micro-Raman spectroscopies



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#### Abstract

This study detected and investigated the degradation effects and dynamics of the tin amalgam reflective layer of ancient mirrors. The experimental results have been obtained by different spectroscopic techniques: Light Microscopy, Scanning Electron Microscopy with Energy Dispersive X-ray Microanalysis (SEM–EDS), Atomic Force Microscopy (AFM) and µ-Raman Spectroscopy. The latter two were applied for the first time to characterize the amalgam of ancient mirrors. Amalgam alteration products in form of concentric rings with alternate different compositions of tin oxides have been studied on both the amalgam reflective layer inner surface and on its cross-sections. For the first time SEM images of micro-cross-sections of different amalgam degradation products, such as amalgam "drops" and craters, have been obtained. The size of the nano-particles of tin oxides, as amalgam degradation products, was estimated around 3–8 nm by AFM and micro-Raman, as already previously detected through TEM. Micro-Raman showed the presence of both cassiterite and romarkite nano-crystals forming the concentric rings, as only previously hypothesized. This investigation allowed to obtain further experimental evidence and a more advanced comprehension about the structure and formation phases of the amalgam decay products, giving further support to some statements by previous authors and suggesting some assertions different from the ones of previous studies.

Keywords Mirror · Amalgam · Corrosion · Cassiterite · Romarkite · Micro-Raman

## 1 Introduction

#### 1.1 Amalgam mirrors

This study aims to deepen the understanding of dynamics and products of amalgam degradation in ancient amalgam mirrors. It led to new experimental evidence, obtained by multispectroscopic analyses of some samples of Venetian mirrors from the eighteenth century.

The majority of studies on amalgam ancient mirrors were performed in the last decade and concerned the characterization of the tin-mercury amalgam layer, whose interface with glass constitutes the reflective surface of the mirror [1-12]. They indicated that this reflective layer

contains about 25 wt% mercury and 75 wt% tin and that it consists of two different phases: a solid one, distributed as "islands", composed by about 20 wt% mercury and 80 wt% tin, surrounded by a liquid phase composed by about 95–100 wt% mercury and 0–5 wt% tin [1–12]. The amalgam alteration involves mercury evaporation and formation of tin monoxide (SnO, romarkite) and dioxide (SnO<sub>2</sub>, cassiterite). This causes the formation of amalgam drops and tin oxides craters [4–7, 11, 12]. In particular, the results of a recent study [11] suggested a possible mechanism of transformation from amalgam drops to craters.

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## 1.2 Corrosion morphology

Several studies among those cited above concern the alteration processes of the amalgam layer on the back side of ancient mirrors (interface amalgam-air) [1, 4–8, 10–12], while a recent study was based for the first time on the investigation of the amalgam side in contact with the glass (reflective surface) [12]. It highlighted the formation of a superficial morphology of the amalgam alteration products in form of concentric rings, with two alternate different compositions. These rings represent the superficial section of a system of hemispherical concentrically stratified calottes. They seem to have grown from an initial corrosion point on the air-interfaced surface towards the bulk, to reach the opposite reflective surface [12].

Similar patterns were found in some archaeological high tin bronze [13, 14] and glass artifacts [15–17]. These banded structures constitute uncommon cases of corrosion and the composition of the rings and their formation mechanisms, related to corrosion processes of bronze, glass and mirror amalgam, remain still rather unclear.

This particular corrosion morphology, similar to the so-called Liesegang rings [18], could be associated to cyclical chemical processes, maybe related to periodical fluctuations of environmental conditions, such as temperature, humidity and pH, depending on the nature of the involved materials and phenomena. It is nevertheless possible to suppose that these band structures depend simply on some intrinsic cyclical oxide-reductive process. In fact, Fink [19] observed that for archaeological bronzes, buried for hundreds of years in sites where the fluctuations of the above mentioned parameters were negligible, the sequence of corrosion rings cannot be related with environmental periodic variations.

Conversely, Scott [13] considered phenomena similar to a Liesegang process as responsible for the band structure found in ancient bronzes, and hypothesized that the banding phenomena could be related with the presence of tin in the alloy.

A deeper investigation on the mirror amalgam reflective surface could give additional information for understanding the amalgam alteration processes and the causes for the formation of these rings.

# 1.3 Aim of the work

This study aims to obtain further experimental evidence about degradation products and related corrosion morphology of the tin-mercury amalgam of ancient mirrors. The investigation of the amalgam reflective side with different analytical techniques focused on the possibility to improve our understanding of the formation mechanism of the ring structures, at least about the sequence of the phases causing this final effect. The ultimate goal should be not only the full comprehension of these degradation phenomena, but the possibility to create the basis for setting up some techniques of inhibition of these degradation processes.

Light Microscopy (LM) and Scanning Electron Microscopy with Energy Dispersive X-ray Microanalysis (SEM–EDS) have been the preliminary tools to identify and characterize the banded structures. Atomic Force Microscopy (AFM) is proposed to investigate the tridimensional morphology of the corrosion products on the reflective surface of the samples. Both SEM–EDS and  $\mu$ -Raman analyses were aimed to detect further evidence about the composition of the concentric rings.

# 2 Materials and methods

A large number of flakes, detached from the degraded amalgam layers of 5 ancient amalgam mirrors, was studied. The flake detachment was due to the natural decay of the ancient mirrors so that its inner surface in contact with the glass sheet (reflective surface), flat, intact and not contaminated, could be investigated. The flakes analyzed had areas between 0.5 to 1.5 cm<sup>2</sup> and thicknesses of about 50  $\mu$ m.

The reflective surfaces of the flakes were first observed by LM and then studied by different analytical techniques, such as SEM–EDS, AFM and  $\mu$ -Raman Spectroscopy.

The light microscope used was a Leica MZ12.

SEM-EDS investigations were carried out using a Scanning Electron Microscope JEOL JSM 5900 equipped with X-ray microanalysis Oxford ISIS EDS with a Si(Li) detector. The samples were analyzed at low vacuum, with working distance around 10 mm, beam voltage between 10 and 25 kV and current between 0.01 and 10 nA. In the EDS maps the concentration ranges of every analyzed element have been represented in a color scale. Every color scale is independent from the other ones and hence the different color scales are not comparable.

AFM measurements were performed in air with a NOVA microscope equipped with a 10 nm curvature radius cantilever tip used in tapping mode. Lateral resolution was about 60 nm, and the vertical one about 1 nm. AFM scans were obtained on areas of  $30 \times 30$  and  $50 \times 50$  µm on flakes of very degraded amalgam.

Micro-Raman analyses were carried out with a Renishaw InVia Raman Microscope equipped with a green  $Ar^+$  ion laser with wavelength 514 nm and power 50 mW. The

spectra were acquired in the range  $100-1000 \text{ cm}^{-1}$ , with an exposure of 10 s and accumulation number between 10 and 20. The laser power used was between 1 and 10%. The different areas of the samples were investigated by long working distance objectives of different magnifications ( $20 \times$ ,  $50 \times$  and  $100 \times$ ).

SEM–EDS, AFM and  $\mu$ -Raman surface analyses were performed on flake samples of a very degraded amalgam portion, fixed by a carbon double-sided adhesive tape on an aluminum sample holder. Flakes were directly detached from the glass by the adhesive tape laying onto the back side of a degraded amalgam mirror, immediately inspected without further handling.

Cross-sections of the amalgam flakes were prepared by standard metallographic techniques (using transparent epoxy resin, grinding with SiC papers from P180 to P4000 and polishing with diamond suspensions) and then inspected by a LM and SEM–EDS.

## **3** Results and discussion

#### 3.1 Light microscope investigation

The reflective surfaces of different flakes have been observed by LM. They are presented in Figs. 1, 2 and 3. They show some patterns of a very degraded amalgam layer: all samples show the presence of many structures of different size characterized by concentric bright and dark rings.

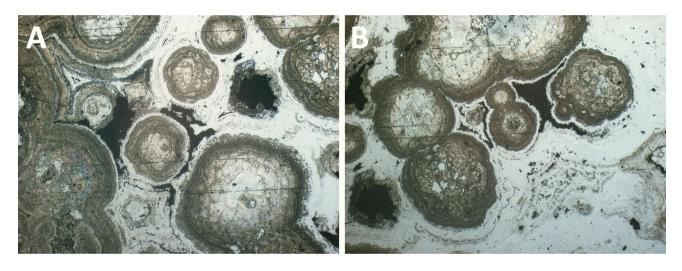


Fig. 1 LM images (×25) of some banded structures on the reflective side of the amalgam layer

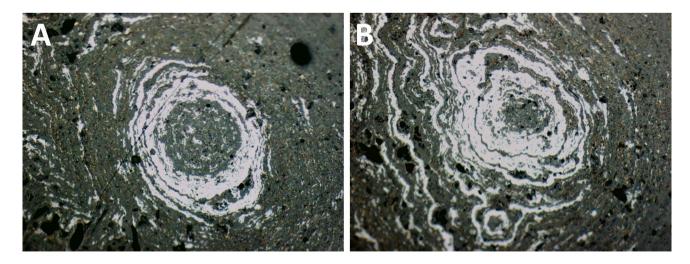


Fig. 2 LM images (×100) of a single banded structure on the reflective side of the amalgam layer

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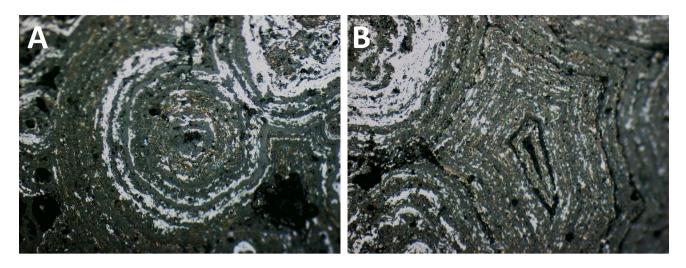


Fig. 3 LM images (× 200) of some differently banded structures on the reflective side of the amalgam layer

#### 3.2 SEM–EDS investigation

Some SEM backscattered electrons images of the amalgam layer in cross sections were recorded in order to get information about morphology and composition of the corrosion products. It is the first time that images of cross sections involving amalgam drops and tin oxide craters have been obtained.

In these images it is possible to recognize the different steps of the corrosion processes, already presented in a previous work on surface images [11].

The Figs. 4 and 5 show cross sections of different areas of a flake: it is worth noting the formation of a layer of tin oxides of about 50 µm thickness together with amalgam drops (Fig. 4) and craters (Fig. 5). Drops are present also on the top of some well pronounced craters, as previously detected [4-7, 10-12]. SEM pictures in Fig. 4a, b showed that they are formed by the two phases of the tin amalgam, in contrast with previous studies, where the drops were characterized as rich in mercury [11] or described as formed almost entirely by mercury [4–7]. In summary, the darker and lighter zones in Figs. 4 and 5 consist of tin oxide and mercury-tin amalgam respectively. Figures 4 and 5 show, moreover, sequential oxidation steps of the amalgam layer: (1) the amalgam drop is surrounded by tin oxides (Fig. 4a), (2) the oxide layer growing around the amalgam drop (Fig. 4b, c), (3) the so called "crater" produced by further growth of the tin oxides around the amalgam drops, with some amalgam remaining on the top of the crater (Fig. 4d, e) and (4) the crater completely formed, without amalgam residues on its top (Fig. 4f).

This sequence, only supposed in a previous study [11] by the SEM surface observation of a large area of the amalgam back side, is now shown also in a cross-section, and

evidenced by the pattern of the subsequent rings along the crater (Fig. 4c, d).

Figures 4 and 5 show that even the craters are formed by subsequent rings, as the calottes found on the reflective surface of the degraded amalgam.

The secondary electron image of the reflective surface of a flake (Fig. 6b) shows that its surface (BE SEM image in Fig. 6a) is almost flat even at the microscopic level. This morphology of the reflective surface is due to the compression of the alteration compounds against the glass flat sheet during the expansion of the corrosion towards the reflective surface.

In the images in Fig. 7 it is possible to observe that the banded structures have different shapes and features: some of these consist of almost circular rings and of a dark central area (Fig. 7a), while in other ones the ring patterns are irregular and weakly pronounced, often with a zone of not corroded amalgam (Fig. 7b), or a hole at the center (Fig. 7c). These differences may be ascribed to the different evolution step of the corrosion process in different areas.

The darker area in the centre of the structure (Fig. 7a) may correspond to a hemispherical calotte center at the surface. As hypothesized in previous studies on the basis of experimental evidence, the corrosion should start at the back free surface of the amalgam layer and grow as a hemispherical calotte into the bulk of the amalgam layer [3, 11, 12]. When the growth of a calotte reaches the reflective surface adherent to the glass sheet, the calotte can expand only two-dimensionally and it is detected as superficial ring structure on the reflective surface.

Otherwise, the presence of the amalgam at the center (Fig. 7b) of a banded structure may indicate that the observed area corresponds, on the reflective surface, to an amalgam drop partially surrounded by tin oxides, like the one shown in Fig. 4b. The image in Fig. 7b could therefore

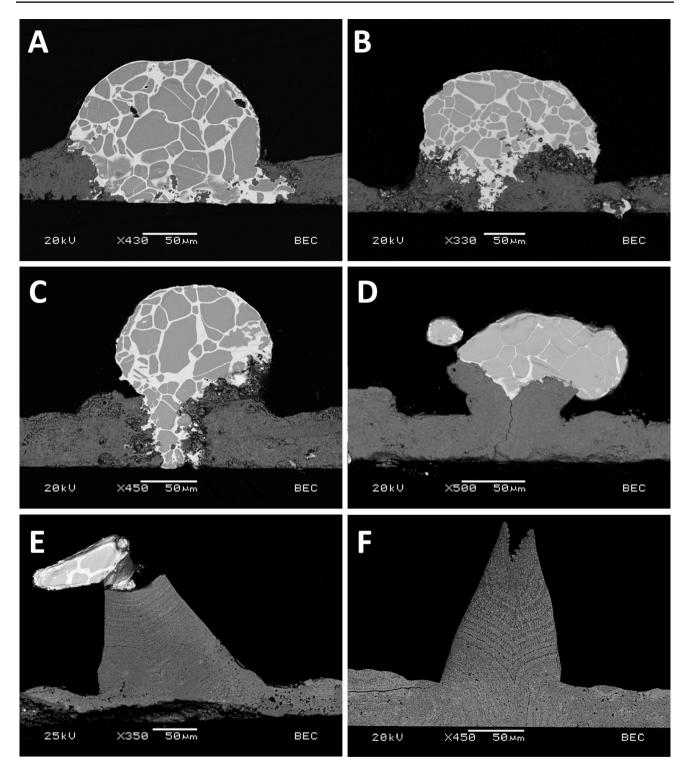
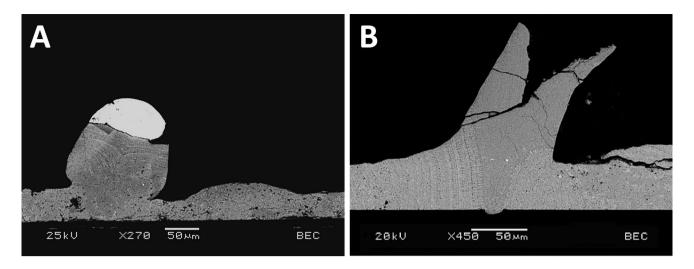


Fig. 4 SEM images in backscattered electrons of a cross section of a degraded amalgam reflective layer flake, showing the tin oxide layer and amalgam drops and craters at different magnifications. The images **a**–**f** show the sequential oxidation of the amalgam layer

correspond to the morphology on the reflective surface of the flake shown by the cross section of Fig. 4b.

The presence of a hole in the centre (Fig. 7), instead of the amalgam as in Fig. 7b, could indicate a subsequent

degradation step of the structure in Fig. 7b, when all the amalgam of a drop is corroded forming the crater, leaving the central area empty.



**Fig. 5** SEM images in backscattered electrons of a cross section of a degraded amalgam reflective layer flake; **a** the tin oxide layer and a crater with an amalgam drop on the top; **b** the tin oxide alone after the detachment of the amalgam drop

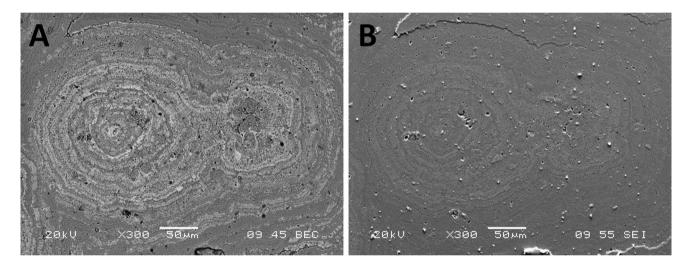
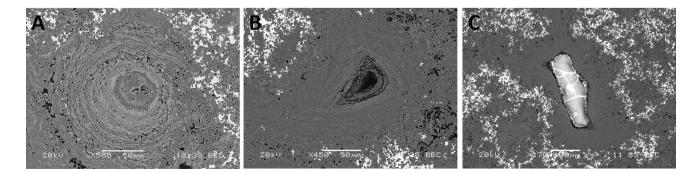


Fig. 6 SEM images of the reflective surface of the degraded amalgam layer with banded structures; **a** backscattered electrons image; **b** secondary electrons image



**Fig. 7** SEM backscattered images of the reflective surface of the degraded amalgam layer with different corrosion patterns; **a** banded structure with circular rings and **a** dark central area;

 ${\bf b}$  banded structure with a zone of uncorroded amalgam at the center;  ${\bf c}$  banded structure with a hole in the center

SN Applied Sciences A Springer Nature journat SEM-EDS maps showed that the dark center and the external rings of the banded structure have different tin contents [12]. Localised EDS measurements indicated that the darker areas are mostly constituted by cassiterite and the grey ones by romarkite.

These results are strengthened by the SEM–EDS X-ray maps of Fig. 8 that show more clearly that the subsequent rings have alternatively two different compositions. The anti-correlation between the Sn and the O maps seems to indicate that the rings are formed by different mixtures of the two tin oxides, romarkite and cassiterite.

#### 3.3 AFM

Atomic Force Microscopy (AFM) was carried out on some areas of the rings in order to better explore the 3D morphology of the banded structures. This technique can be more sensitive and detailed on surface morphology than SE-SEM images, and has here been applied for the first time to the study of amalgam mirrors.

Figure 9b shows an AFM scan image of the square area indicated in Fig. 9a.

The AFM scan shows that the rings have alternatively different width and vertical thickness. In spite of the appearance of different levels of rings in the AFM scan, the measured difference of vertical thickness between adjacent rings is very small (about 30–40 nm). The different

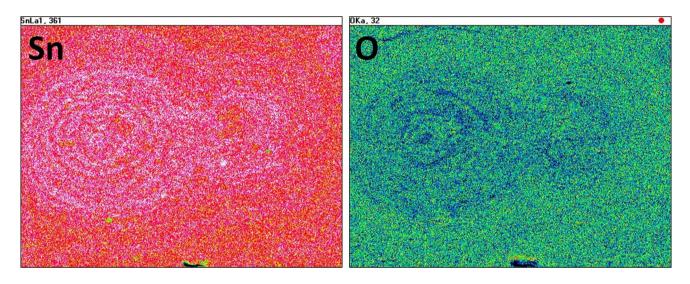


Fig. 8 EDS maps of Sn and O of the area in Fig. 6

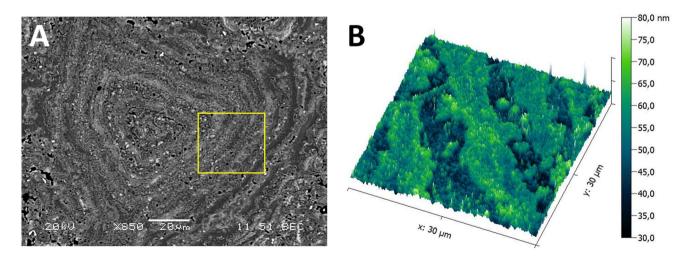


Fig. 9 a SEM image (×850) of a banded structure on the reflective surface of the degraded amalgam layer; b AFM scan of the yellow square area of Fig. 9a

SN Applied Sciences A Springer Nature journal thicknesses of the rings may suggest a process in which two different phases of tin oxidation and precipitation happen subsequently, or the presence of two different oxidation states.

In spite of the high spatial resolution (about 60 nm), AFM scans did not detect the presence of crystal particles, indicating particle sizes smaller than the scan resolution. This is in agreement with Herrera et al. [7], who described the presence of tin oxide nanoparticles in amalgam mirrors. These authors detected and measured by X-ray Diffraction and Transmission Electron Microscopy, SnO<sub>2</sub> nanocrystals of about 4–5 nm average size [7].

AFM measurements with a higher resolution in order to detect the presence of nanocrystals led to artifacts and not clear scans. It must be considered that the surface of the analyzed samples is porous and can contain traces of mercury and dust particles absorbed after the flake detaching.

#### 3.4 Micro-Raman

In order to identify the tin alteration compounds forming these concentric ring structures, a  $\mu$ -Raman investigation was carried out on the flake samples, on both the bright and dark areas of a ring structure.

Figure 10b shows the Raman spectrum of the central bright area of Fig. 10a. The spectrum is characterized by a strong peak at about 210 cm<sup>-1</sup>, related to romarkite (SnO) [20], probably as nanocrystals, as suggested by AFM. Figure 11b shows the Raman spectrum of the central dark area of Fig. 11a. The spectrum shows the presence of the romarkite peak at 210 cm<sup>-1</sup> and of a broad peak at about 633 cm<sup>-1</sup>, distinctive of cassiterite crystals (SnO<sub>2</sub>) [21–23].

A broad band between 400 and 800 cm<sup>-1</sup>, with a maximum at about 575 cm<sup>-1</sup>, is also present in the spectra in Figs. 10b and 11b. This band was found in several studies

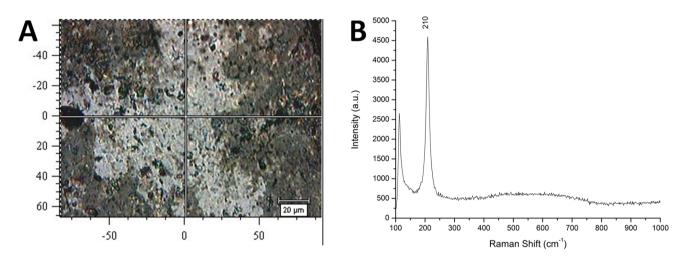


Fig. 10 a Optical image of a bright ring of the banded structure; b Raman spectrum acquired on the bright area at the center of Fig. 10a

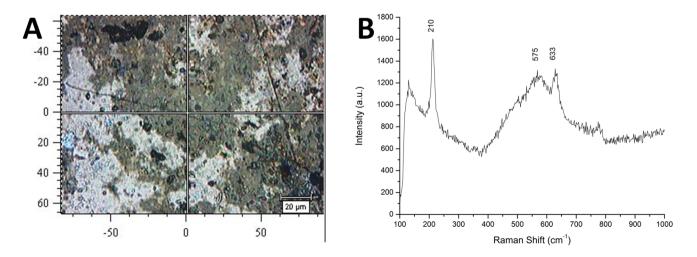


Fig. 11 a Optical image of a dark ring of the banded structure; b Raman spectrum acquired on the dark area at the center of Fig. 11a

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on cassiterite nanoparticles and can be due to the very small size of the particles [21, 24, 25].

According to Dieguez et al. [21], who characterized by TEM cassiterite particles with different sizes obtained by calcinating  $SnO_2$  nanoparticles at different temperatures, it is possible to estimate the linear size of the  $SnO_2$  particles from the study of the Raman spectrum major peaks.

In his work, Dieguez et al. [21] demonstrated that there is a correlation between the cassiterite nanoparticles size and the peaks shift in the Raman spectra (see Table 2 in Ref. 21). According to this study, the maximum at 633 cm<sup>-1</sup> (mode A<sub>1g</sub>) and the maximum of the broad band at around 575 cm<sup>-1</sup> (mode S1) allow to assign to the particles sizes in the range 3-8 nm, taking in account the calculated precision reported in the Ref. 21, where no accuracy determination is considered.

Cassiterite crystals were already found by XRD in previous studies [1, 3, 4] but this technique did not allow to find the presence of cassiterite nanocrystals and to determine the composition of the single ring structures.

The AFM and the micro-Raman analyses of the present study allowed to determine the presence not only of cassiterite nanocrystals as degradation product of amalgam, as suggested by Herrera et al. [7], but also the presence of romarkite nano-crystals in the concentric rings.

Moreover, these results allowed to state that the bright areas shown by SEM images correspond to the white ones detected by LM observation and, vice versa, that the SEM dark looking areas correspond to the dark ones observed by microscopic inspection.

The Raman maps in Fig. 12b, d illustrate the pattern of different corrosion products in the same area as shown in Fig. 12a. Some areas are rich in romarkite (Fig. 12b) and poor in cassiterite (Fig. 12c, d).

By means of this micro-Raman investigation it is thus possible to assume that the bright rings of the concentric structures are formed mostly by romarkite nanocrystals and the dark ones are composed by a mixture of cassiterite and romarkite nanocrystals. These results are in agreement with the localised EDS measurement on the rings by Arizio et al. [12] that suggested that the two areas are differently rich of the two tin oxides romarkite (SnO) and cassiterite  $(SnO_2)$  [12].

#### 3.5 Decay mechanism

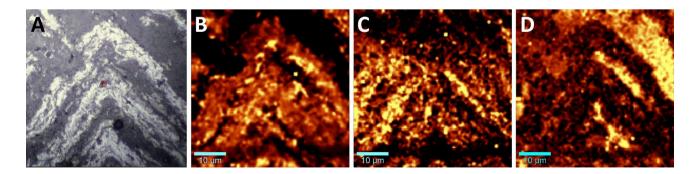
The non-degraded tin amalgam of mirrors consists of a layer about 0.1 mm thick, composed of two phases with different contents of tin and mercury.

Previous studies [1, 10] on amalgam of ancient mirrors demonstrated, by artificial aging, that the size of the crystalline solid phase grows with the amalgam age to reach a maximum size. This seems to be due to the rearrangement of the phases because of the natural evaporation (and loss) of mercury mostly from the liquid phase. In fact, as long as the liquid phase is in excess, mercury evaporation causes the approach of the solid phase crystals to each other. When the liquid phase amount is no more sufficient to surround the solid phase crystals, the solid phase re-arranges, creating crystals of larger sizes, so that a smaller volume of liquid phase is necessary to surround them.

When the solid phase crystals reach their maximum size [1, 10], further mercury evaporation leads to the precipitation of excess tin and oxidation can start from the back side of the mirror in contact with the atmosphere. The tin particles oxidized on the back side are thus the nucleus of the calottes growing in the bulk of the amalgam layer to reach the opposite front of the reflective surface, as found in the previous studies [3, 12].

Instead of conventional tin oxidation mechanisms, that cause the formation of a tin oxide passivation layer, in this case the degradation products are probably porous, allowing the diffusion of oxygen in the deeper layers, forming the oxide calottes. This continuous oxidation can also be promoted by the continuous Hg evaporation from the not still degraded amalgam.

The oxidation process causes the reduction of the amalgam volume of around 50% with respect to the original



**Fig. 12 a** Optical microscope image of some rings; **b** Raman map of the 210 cm<sup>-1</sup> peak ascribable to romarkite; **c** Raman map of the 630 cm<sup>-1</sup> peak ascribable to cassiterite; **d** Raman map of the 580 cm<sup>-1</sup> peak ascribable to cassiterite

SN Applied Sciences A Springer Nature journal one: the thickness of the oxidized layer is usually around 50  $\mu\text{m}.$ 

Because of its high surface tension and the volume reduction due to both mercury evaporation and tin oxidation, the remaining non-degraded amalgam is visible as amalgam drops on the back surface. The growth of the tin oxides around the amalgam drops causes the formation of the craters found in the previous studies, until no amalgam or mercury is present on their top.

The presence of hemispherical layers of tin oxides could indicate that the amalgam progressive corrosion process involves cyclic phases. Presently, no clear experimental evidence is available to clearly state the origin of the presence of alternate layers of different composition. They could depend on the variation of environmental parameters, such as temperature or humidity, or be due to intrinsic features related to composition of amalgam and to oxidation/precipitation processes.

# 4 Conclusions

This study focused on the investigation of degradation of the reflective surface of ancient amalgam mirrors. The degradation processes have been examined through the analytical results obtained by different spectroscopic techniques: SEM–EDS and for the first time AFM and micro-Raman have been used to characterize tin amalgam and degradation products of ancient mirrors.

This combination of techniques allowed to deepen the knowledge on the amalgam reflective layer and to clarify morphologies and processes about its decay products, leading to some considerations differing from those made in previous studies.

For the first time, SEM images of cross-sections of different amalgam degradation products, such as amalgam "drops" and craters, were obtained. They show that the forms previously described as "mercury rich drops" or "mercury spheres" are instead composed by bi-phasic, non-degraded amalgam.

Differently from some previous studies, by micro-Raman analyses was possible to detect both romarkite and cassiterite as degradation compounds of the amalgam layer.

Some new experimental evidence gave further support to previous assertions: AFM and the micro-Raman analyses detected the presence of nanocrystals, as previously suggested. They also allowed to estimate the size of the nanoparticles around 3–8 nm, as already previously detected through TEM.

Moreover, Micro-Raman analyses detected the presence of both cassiterite and romakite nano-crystals forming the concentric rings consisting of amalgam degradation products, as hypothesized only in a previous study on the basis of some SEM–EDS results.

The analysis of the concentric calottes, performed with different analytical techniques on both plane and crosssection led to a deeper understanding of the amalgam degradation products and allowed to explain more clearly structures and processes of the amalgam decay products.

Lastly, this investigation allowed to propose a general macroscopic amalgam decay mechanism defining its subsequent phases. It remains still unclear the origin of the formation of the subsequent calottes with two alternate different compositions, evidenced as Liesegang-like rings. They could depend on cyclic variation of environmental parameters or on intrinsic chemical/physical features of the amalgam system.

## **Compliance with ethical standards**

**Conflict of interest** The authors declare that they have no conflict of interest.

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