



Correction to: Bošković's Method for Determining the Axis and Rate of Solar Rotation by Observing Sunspots in 1777

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The original online version of this article has been corrected to improve grammatical, syntactical and spelling mistakes throughout the text. Figure 6 was also corrected to remove a purple line that was inadvertently included in the axis. More specifically:

On page 24, the paragraph included additional spaces that were removed:

The ecliptic longitude of the ascending node $\Omega = 80^{\circ}07'-04''$ and the solar equator's inclination $i = 7^{\circ}19'-23''$ are arithmetic means of the four values rounded to the whole angular second: $\bar{\Omega} = \Sigma \Omega_i / n = 320^{\circ}28'-14'' / 4 = 80^{\circ}07'-04.2'' = 80.1178333^{\circ} \approx 80^{\circ}07'-04'' = 80.1177778^{\circ} \approx 80.12^{\circ}$ and $\bar{i} = \Sigma i_i / n = 29^{\circ}17'-31'' / 4 = 7^{\circ}19'-22.75'' = 7.322986111^{\circ} \approx 7^{\circ}19'-23'' = 7.323055556^{\circ} \approx 7.32^{\circ}$. The ecliptic longitude of ascending node $\Omega = 80.12^{\circ}$ and solar equator's inclination $i = 7.32^{\circ}$ are like in Table 1 in Wöhl (1978):

Value 1900 was corrected to 1976:

The values $\Delta \Omega_{1900}^{W-PW}$ and $\Delta \Omega_{1900}^{W-PW}$ are slightly different from the present work values Ω_{1900}^{PW} and Ω_{1976}^{PW} and ecliptic longitudes of the ascending node $\Omega_{1900} = \Omega_{1900}^W$ and $\Omega_{1976} = \Omega_{1976}^W$ given by Wöhl (1978).

The page numbers in the following citations were wrongly reported:

- 169 was changed to 168

The original article can be found online at <https://doi.org/10.1007/s11207-023-02214-6>

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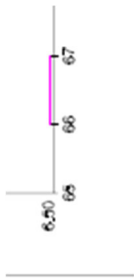
- 168 was changed to 169

- iii) Sidereal period T' , which is determined by the method for the period. It is the arithmetic mean of the six pairs of mean solar time $T.M.$ and the positions of the same sunspot, which Bošković used in *Tab. X.* (Boscovich, 1785b, *Tab. X.*, p. 469).
- iv) The ecliptic longitude of the ascending node $N = \Omega$ and the solar equator inclination towards the ecliptic i , which were determined by the method for Ω and i . It uses three positions of the same sunspot in *Tab. XII.* (Boscovich, 1785b, *Tab. XII.*, p. 468).

The letter I in the heading was changed to lowercase i:

2.2. The Method for *i*

Figure 6 included an unnecessary purple line on the axis that was now removed:



Declarations

Competing Interests The authors declare no competing interests.

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