# Linearization of elasticity models for incompressible materials 

Edoardo Mainini(0) and Danilo Percivale


#### Abstract

We obtain linear elasticity as $\Gamma$-limit of finite elasticity under incompressibility assumption and Dirichlet boundary conditions. The result is shown for a large class of energy densities for rubber-like materials.


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## 1. Introduction

Organic materials require sophisticated mechanical models. They exhibit nonlinear stress-strain behavior and are often elastic up to large strains. Typical examples are natural rubber as well as artificial elastic polymers with rubber-like properties. These materials resist volume changes, are very compliant in shear and their shear modulus is of orders of magnitude smaller than the shear resistance of most metals. This motivate the modeling of rubber-like materials as being incompressible and hyperelastic [ $26,43,45,52$ ], so that their phenomenological description requires the introduction of an empirical strain energy density [ $9,39,55]$. Similar properties are observed in soft biological tissues: in many biomechanical studies blood vessels are modeled as nonlinear elastic materials that are incompressible under physiological loads [12, 24,27,29].

From a mathematical point of view, we consider a hyperelastic body that occupies a bounded open region $\Omega \subset \mathbb{R}^{3}$ in its reference configuration. In presence of a body force field $\mathbf{g}: \Omega \rightarrow \mathbb{R}^{3}$, the energy of the system is given by the stored elastic energy and the contribution of the external forces

$$
\int_{\Omega} \mathcal{W}^{I}(x, \nabla \mathbf{y}(x)) \mathrm{d} x-\int_{\Omega}(\mathbf{y}(x)-x) \cdot \mathbf{g}(x) \mathrm{d} x .
$$

Here, $\mathbf{y}: \Omega \rightarrow \mathbb{R}^{3}$ denotes the deformation field, $\nabla \mathbf{y}$ denotes the deformation gradient and $\mathcal{W}^{I}$ is the incompressible elastic energy density. $\mathcal{W}^{I}$ is assumed to be frame indifferent and minimized at the identity with $\mathcal{W}^{I}(x, \mathbf{I})=0$, so that without an external load $\mathbf{y}(x)=x$ is a minimizer of the total energy corresponding to the stress-free configuration $\Omega$. In order to take into account that the body is incompressible, $\mathcal{W}^{I}(x, \mathbf{F})=+\infty$ whenever $\operatorname{det} \mathbf{F} \neq 1$.

A common approach in the study of rubber-like materials is to consider a stored energy density $\mathcal{W}$ which is defined in the compressible range, the kinematic constraint $\operatorname{det} \mathbf{F}=1$ being relaxed to a volumetric penalization: a typical expression of $\mathcal{W}$ is given by the usual isochoric-volumetric form

$$
\begin{equation*}
\mathcal{W}(x, \mathbf{F}):=\mathcal{W}_{\text {iso }}\left(x,(\operatorname{det} \mathbf{F})^{-1 / 3} \mathbf{F}\right)+\mathcal{W}_{\text {vol }}(\operatorname{det} \mathbf{F}) \tag{1.1}
\end{equation*}
$$

where $x \in \Omega$ and $\operatorname{det} \mathbf{F}>0$ (extended to $+\infty$ if $\operatorname{det} \mathbf{F} \leq 0$ ). Here, the nonnegative function $\mathcal{W}_{\text {iso }}\left(x, \mathbf{F}_{*}\right)$ is defined for every $\mathbf{F}_{*}$ such that $\operatorname{det} \mathbf{F}_{*}=1$ and satisfies $\mathcal{W}_{\text {iso }}(x, \mathbf{I})=0$. Moreover, $\mathcal{W}_{\text {vol }}(t) \geq 0$ for every $t>0$ and $\mathcal{W}_{\text {vol }}(1)=0$. In fact, we shall first choose a compressible energy density $\mathcal{W}$, for instance in the
form (1.1), requiring frame indifference and other suitable regularity conditions that will be introduced in Sect. 2. Then, we shall define the incompressible energy density $\mathcal{W}^{I}$ by setting $\mathcal{W}^{I}(x, \mathbf{F})=\mathcal{W}(x, \mathbf{F})$ if $\operatorname{det} \mathbf{F}=1$ and $\mathcal{W}^{I}(x, \mathbf{F})=+\infty$ if $\operatorname{det} \mathbf{F} \neq 1$.

The simplest model is the homogeneous Neo-Hookean solid: the energy density is of the form (1.1) with

$$
\begin{equation*}
\mathcal{W}_{\text {iso }}\left(\mathbf{F}_{*}\right):=\mu\left(\left(\operatorname{Tr}\left(\mathbf{F}_{*}^{T} \mathbf{F}_{*}\right)-3\right),\right. \tag{1.2}
\end{equation*}
$$

where $\operatorname{det} \mathbf{F}_{*}=1$ and the shear modulus $\mu$ is determined experimentally: this model fits material behaviors with sufficient accuracy under moderate straining while, at higher strains, it can be replaced by the more general Ogden model, namely

$$
\begin{equation*}
\mathcal{W}_{\mathrm{iso}}\left(\mathbf{F}_{*}\right):=\sum_{p=1}^{N} \frac{\mu_{p}}{\alpha_{p}}\left(\operatorname{Tr}\left(\left(\mathbf{F}_{*}^{T} \mathbf{F}_{*}\right)^{\alpha_{p} / 2}\right)-3\right) \tag{1.3}
\end{equation*}
$$

where $N, \mu_{p}, \alpha_{p}$ are material constants. For particular values of the material constants the Ogden model reduces to either the Neo-Hookean solid ( $N=1, \alpha_{1}=2$ ) or the so called Mooney-Rivlin material $\left(N=2, \alpha_{1}=2, \alpha_{2}=-2\right)$ which is often applied to model incompressible biological tissue, see [41,42]. Another phenomenological material model, motivated for simulating the mechanical behavior of carbonblack filled rubber and for its important applications in the manufacture of automotive tyres, has been introduced by Yeoh, see [56,57]: the isochoric part of the strain energy density is given by

$$
\begin{equation*}
\mathcal{W}_{\text {iso }}\left(\mathbf{F}_{*}\right):=\sum_{k=1}^{3} c_{k}\left(\left(\operatorname{Tr}\left(\mathbf{F}_{*}^{T} \mathbf{F}_{*}\right)-3\right)^{k}\right. \tag{1.4}
\end{equation*}
$$

where $c_{k}, k=1,2,3$ are material constants. For a complete description of the main properties of such energy densities and other models, we refer to the classical monographs such as $[11,26,44]$ or to the reviews in $[3,9,39]$, see also $[7,30,55]$.

Let us now introduce the linearization. If $h>0$ is an adimensional parameter, we scale the body force field by taking $\mathbf{g}:=h \mathbf{f}$ and set $\mathbf{y}(x):=x+h \mathbf{v}(x)$. The resulting total energy is

$$
\mathcal{E}_{h}(\mathbf{v}):=\int_{\Omega} \mathcal{W}^{I}(x, \mathbf{I}+h \nabla \mathbf{v}) \mathrm{d} x-h^{2} \int_{\Omega} \mathbf{f} \cdot \mathbf{v} \mathrm{d} x
$$

and it seems meaningful to ask what is the correct scaling of energies $\mathcal{E}_{h}$ as $h \rightarrow 0^{+}$. Roughly speaking in the spirit of [6] (see also [31,46-49,51]) we will show that, under suitable boundary conditions,

$$
\inf \mathcal{E}_{h}=h^{2} \min \mathcal{E}_{0}+o\left(h^{2}\right)
$$

where

$$
\mathcal{E}_{0}(\mathbf{v}):=\left\{\begin{array}{l}
\frac{1}{2} \int_{\Omega} \mathbb{E}(\mathbf{v}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v}) \mathrm{d} x-\int_{\Omega} \mathbf{f} \cdot \mathbf{v} \mathrm{d} x \text { if } \operatorname{div} \mathbf{v}=0 \text { a.e. in } \Omega \\
+\infty \\
\text { otherwise. }
\end{array}\right.
$$

The quadratic form appearing in the expression of $\mathcal{E}_{0}$ features the infinitesimal strain tensor $\mathbb{E}(\mathbf{v}):=$ $\frac{1}{2}\left(\nabla \mathbf{v}+\nabla \mathbf{v}^{T}\right)$ and can be obtained by a formal Taylor expansion of $\mathcal{W}$ around the identity matrix $\mathbf{I}$, with $D^{2}$ denoting the Hessian of $\mathcal{W}(x, \cdot)$. We stress that purely volumetric perturbations of $\mathcal{W}$ do not affect $\mathcal{W}^{I}$ nor $\mathbb{E}(\mathbf{v}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v})$, due to the divergence-free condition. Moreover, we will prove that if

$$
\mathcal{E}_{h}\left(\mathbf{v}_{h}\right)-\inf \mathcal{E}_{h}=o\left(h^{2}\right),
$$

then

$$
\mathbf{v}_{h} \rightharpoonup \mathbf{v}_{0} \in \operatorname{argmin} \mathcal{E}_{0}
$$

in the weak topology of a suitable Sobolev space. Since the compliance in shear of rubber-like materials and the strong nonlinearity of their stress-strain behavior even at modest strain do not allow to suppose that small strains correspond to small loads, it must be clarified that it would not be reasonable to assume that either $h \mathbf{v}$ or $h \mathbb{E}(\mathbf{v})$ are small in any sense. Anyhow we highlight that linearized models may provide a good approximation that fits experimental data, see for instance [32].

From the viewpoint of the Calculus of Variations, derivation of linearized elasticity from finite elasticity has a long history that started in [15], where $\Gamma$ convergence and convergence of minimizers of the associated Dirichlet boundary value problems are proven in the compressible case (see also [1,2,4,34,35] for more recent results). In this paper we show how these results can be extended to the incompressible case, i.e., assuming the constraint $\operatorname{det} \nabla \mathbf{y}=1$ on admissible deformations fields. It is well-known that such a constraint poses some challenges to the $\Gamma$-convergence analysis (see for instance the derivation of a twodimensional model for elastic plates in [13]). Indeed, some novel approach (that we develop in Lemma 4.1) is required for the construction of recovery sequences, due to the necessity of recovering the linearized incompressibility constraint $\operatorname{div} \mathbf{v}=0$ with a sequence $\mathbf{v}_{h}$ satisfying $\operatorname{det}\left(\mathbf{I}+h \nabla \mathbf{v}_{h}\right)=1$ a.e. in $\Omega$. Moreover, a different strategy is also needed to ensure that the whole sequence $\left(\mathbf{v}_{h}\right)$ and $\mathbf{v}$ satisfy the same Dirichlet condition. To this end the crucial point consists in analyzing vector potentials: we show in Lemma 3.7 that if $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right), \operatorname{div} \mathbf{v}=0$ in $\Omega$ and $\mathbf{v}=0$ on $\Gamma \subset \partial \Omega$, then, under suitable topological assumptions (see conditions (2.1)-(2.2)), there exists $\mathbf{w} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\operatorname{curl} \mathbf{w}=\mathbf{v}$ in $\Omega$ and $\mathbf{w}=0$ on $\Gamma$. Taking advantage of this result, the construction of the recovery sequence relies on a careful approximation of $\mathbf{w}$ thus outflanking the constraint $\operatorname{div} \mathbf{v}=0$.

We finally mention that other recent contributions have developed a variational analysis for the linearization of finite elasticity under incompressibility constraint, including the case of Neumann boundary conditions $[28,37,38]$. Moreover, in [28] the authors have also considered the more general framework of multiwell potentials.

## Plan of the paper

In Sect. 2 we state the main result. Its proof requires the analysis of vector potentials in Sobolev spaces, which is the object of Sect. 3. In Sect. 4, we develop suitable approximation results that are used for the construction of the recovery sequence in the proof of the main theorem, which is instead contained in Sect. 5.

## 2. Main result

In this section we introduce the basic notation and all the assumptions of our theory. Then, we state the main result.

## Assumptions on the reference configuration

Concerning the reference configuration $\Omega \subset \mathbb{R}^{3}$, we assume that
(i) $\Omega$ is a bounded, simply connected open set,
(ii) $\partial \Omega$ is a connected $C^{3}$ manifold
and we let $\mathbf{n} \in C^{2}(\partial \Omega)$ denote its outward unit normal vector. We will prescribe a Dirichlet boundary condition on a subset $\Gamma$ of $\partial \Omega$. Letting $\partial \Gamma$ denote the relative boundary of $\Gamma$ in $\partial \Omega$ and letting $\mathcal{H}^{2}$ denote
the two-dimensional Hausdorff measure, we assume that
(i) $\Gamma$ is a closed subset of $\partial \Omega$ and $\mathcal{H}^{2}(\Gamma)>0$,
(ii) either $\partial \Gamma=\emptyset \quad$ or $\quad \partial \Gamma$ is a $C^{3}$ one-dimensional submanifold of $\partial \Omega$.

Through the proofs we will also suppose, without loss of generality, that $\Gamma$ is connected. Indeed, all the arguments that we shall develop can be extended to the non connected case by considering each connected component. Besides, in case $\partial \Gamma=\emptyset$, it is possible to assume that $\partial \Omega$ is $C^{2,1}$ only (see Remark 4.5 later on).

## Some notation

For a Sobolev vector field $\mathbf{u} \in W^{1, r}\left(\Omega, \mathbb{R}^{3}\right), r \geq 1$, conditions like $\mathbf{u}=0$ on $\Gamma$ are always understood in the sense of traces. Moreover, we shall often use the decomposition in normal and tangential part at the boundary $\mathbf{u}=(\mathbf{u} \cdot \mathbf{n}) \mathbf{n}+\mathbf{n} \wedge(\mathbf{u} \wedge \mathbf{n})$, where $\wedge$ denotes the cross product. Bold letters will be used in general for vector fields.

## Assumptions on the elastic energy density

Let $\mathcal{W}: \Omega \times \mathbb{R}^{3 \times 3} \rightarrow[0,+\infty]$ be $\mathcal{L}^{3} \times \mathcal{B}^{9}$ - measurable. We assume that $\mathcal{W}$ is frame indifferent and minimized at the identity, i.e.

$$
\begin{gather*}
\mathcal{W}(x, \mathbf{R F})=\mathcal{W}(x, \mathbf{F}) \quad \forall \mathbf{R} \in S O(3), \quad \forall \mathbf{F} \in \mathbb{R}^{3 \times 3}, \quad \text { for a.e. } x \in \Omega  \tag{W1}\\
\min \mathcal{W}=\mathcal{W}(x, \mathbf{I})=0 \quad \text { for a.e. } x \in \Omega \tag{W2}
\end{gather*}
$$

Moreover, we assume that $\mathcal{W}(x, \cdot)$ is $C^{2}$ in a neighbor of rotations (with gradient and Hessian denoted by $D$ and $D^{2}$ ), i.e.,
there exists a neighborhood $\mathcal{U}$ of $S O(3)$ s.t., for a.e. $x \in \Omega, \mathcal{W}(x, \cdot) \in C^{2}(\mathcal{U})$, with a modulus of continuity of $D^{2} \mathcal{W}(x, \cdot)$ that does not depend on x .
Moreover, there exists $K>0$ such that $\left|D^{2} \mathcal{W}(x, \mathbf{I})\right| \leq K$ for a.e. $x \in \Omega$.
The coercivity of $\mathcal{W}$ is described by the following property: there exists $C>0$ and $p \in(1,2]$ such that

$$
\begin{equation*}
\mathcal{W}(x, \mathbf{F}) \geq C g_{p}(d(\mathbf{F}, S O(3))) \quad \forall \mathbf{F} \in \mathbb{R}^{3 \times 3}, \quad \text { for a.e. } x \in \Omega \tag{W4}
\end{equation*}
$$

where $g_{p}:[0,+\infty) \rightarrow \mathbb{R}$ is the convex function defined by

$$
g_{p}(t)= \begin{cases}t^{2} & \text { if } 0 \leq t \leq 1  \tag{2.3}\\ \frac{2 t^{p}}{p}-\frac{2}{p}+1 & \text { if } t \geq 1\end{cases}
$$

Concerning the latter assumption, we refer to [2] for a discussion about the growth properties of energy densities of the form (1.1) to (1.3): for certain ranges of the parameters therein and suitable choice of $\mathcal{W}_{\text {vol }}$, they exhibit a quadratic growth for small deformation gradients and a $p$-growth, $1<p \leq 2$, for large deformation gradients, in particular they satisfy all the above assumptions. Indeed, taking the model choice $\mathcal{W}_{\mathrm{vol}}(t)=c\left(t^{2}-1-2 \log t\right), c>0$, it is shown in [2] that the Neo-Hookean energy density (1.1), (1.2) satisfies $(\mathcal{W} 4)$ for some $p \in(1,2)$, and in fact for $p=2$ when restricting to $\operatorname{det} \mathbf{F}=1$ (then the same holds for the Yeoh model (1.1)-(1.4) with positive $c_{i}$ 's); similarly, the general Ogden model (1.1)-(1.3)
has a $p$-growth with $p \in(1,2)$ if $\mu_{i}>0$ and $0<\alpha_{i}<3$ for all $i=1, \ldots, N$ with $\alpha_{i}>6 / 5$ for at least one $i$. We stress that the Ogden model can exhibit a less than quadratic growth even in the incompressibility regime $\operatorname{det} \mathbf{F}=1$, as can be checked for instance by choosing $N=1, \alpha_{1}=3 / 2$ and $\mathbf{F}=\operatorname{diag}\left(\lambda^{-2}, \lambda, \lambda\right)$ therein.

Let us discuss some first consequences of the above assumptions on $\mathcal{W}$. Since $\mathcal{W} \geq 0$, assumptions $(\mathcal{W} 1)$ and $(\mathcal{W} 2)$ yield

$$
\begin{equation*}
\mathcal{W}(x, \mathbf{R})=0, D \mathcal{W}(x, \mathbf{R})=0 \quad \forall \mathbf{R} \in S O(3), \quad \text { for a.e. } x \in \Omega \tag{2.4}
\end{equation*}
$$

and in particular the reference configuration has zero energy and is stress free. Due to frame indifference there exists a function $\mathcal{V}$ such that

$$
\mathcal{W}(x, \mathbf{F})=\mathcal{V}\left(x, \frac{1}{2}\left(\mathbf{F}^{T} \mathbf{F}-\mathbf{I}\right)\right), \quad \forall \mathbf{F} \in \mathbb{R}^{3 \times 3}, \quad \text { for a.e. } x \in \Omega,
$$

so by setting $\mathbf{F}=\mathbf{I}+h \mathbf{B}$, where $h>0$ is an adimensional small parameter, we have

$$
h^{-2} \mathcal{W}(x, \mathbf{I}+h \mathbf{B})=h^{-2} \mathcal{V}\left(x, h \operatorname{sym} \mathbf{B}+h^{2} \mathbf{B}^{T} \mathbf{B}\right)
$$

where symB:= $\frac{1}{2}\left(\mathbf{B}^{T}+\mathbf{B}\right)$, and then we get via Taylor's expansion

$$
\lim _{h \rightarrow 0} h^{-2} \mathcal{W}(x, \mathbf{I}+h \mathbf{B})=\frac{1}{2} \operatorname{sym} \mathbf{B} D^{2} \mathcal{V}(x, \mathbf{0}) \operatorname{sym} \mathbf{B}=\frac{1}{2} \mathbf{B}^{T} D^{2} \mathcal{W}(x, \mathbf{I}) \mathbf{B}
$$

Hence, $(\mathcal{W} 4)$ implies that for a.e. $x \in \Omega$

$$
\begin{equation*}
\frac{1}{2} \mathbf{B}^{T} D^{2} \mathcal{W}(x, \mathbf{I}) \mathbf{B}=\frac{1}{2} \operatorname{sym} \mathbf{B} D^{2} \mathcal{W}(x, \mathbf{I}) \operatorname{sym} \mathbf{B} \geq \frac{C}{4}|\operatorname{sym} \mathbf{B}|^{2} \quad \forall \mathbf{B} \in \mathbb{R}^{3 \times 3} . \tag{2.5}
\end{equation*}
$$

Here and through the paper, for a matrix $\mathbf{B} \in \mathbb{R}^{3 \times 3}$, we denote $|\mathbf{B}|:=\sqrt{\operatorname{Tr}\left(\mathbf{B}^{T} \mathbf{B}\right)}$.

## Incompressibility

We assume that the material is incompressible. This is done by introducing the incompressible elastic energy density $\mathcal{W}^{I}$ as

$$
\mathcal{W}^{I}(x, \mathbf{F}):= \begin{cases}\mathcal{W}(x, \mathbf{F}) & \text { if } \operatorname{det} \mathbf{F}=1 \\ +\infty & \text { otherwise }\end{cases}
$$

## External forces

We introduce a body force field $\mathbf{f} \in L^{\frac{3 p}{4 p-3}}\left(\Omega, \mathbb{R}^{3}\right)$, where $p$ is such that $(\mathcal{W} 4)$ holds. The corresponding contribution to the energy is given by the following functional, defined for $\mathbf{v} \in W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$

$$
\begin{equation*}
\mathcal{L}(\mathbf{v}):=\int_{\Omega} \mathbf{f} \cdot \mathbf{v} \mathrm{d} x . \tag{2.6}
\end{equation*}
$$

By the Sobolev embedding $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right) \hookrightarrow L^{\frac{3 p}{3-p}}\left(\Omega, \mathbb{R}^{3}\right), \mathcal{L}$ is a continuous functional on $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ and $|\mathcal{L}(\mathbf{v})| \leq C_{\mathcal{L}}\|\mathbf{v}\|_{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)}$ holds for a suitable constant $C_{\mathcal{L}}$ that depends on $\Omega$ and $\mathbf{f}$.

## Rescaled energies

The functional representing the scaled total energy is denoted by $\mathcal{F}_{h}^{I}: W^{1, p}\left(\Omega, \mathbb{R}^{3}\right) \rightarrow \mathbb{R} \cup\{+\infty\}$ and defined as follows

$$
\mathcal{F}_{h}^{I}(\mathbf{v}):=\frac{1}{h^{2}} \int_{\Omega} \mathcal{W}^{I}(x, \mathbf{I}+h \nabla \mathbf{v}) \mathrm{d} x-\mathcal{L}(\mathbf{v})
$$

## Linearized functional

In this paper we are interested in the asymptotic behavior as $h \downarrow 0_{+}$of functionals $\mathcal{F}_{h}^{I}$ and to this aim we introduce the limit energy functional $\mathcal{F}^{I}: W^{1, p}\left(\Omega, \mathbb{R}^{3}\right) \rightarrow \mathbb{R} \cup\{+\infty\}$ defined by

$$
\mathcal{F}^{I}(\mathbf{v}):= \begin{cases}\frac{1}{2} \int_{\Omega} \mathbb{E}(\mathbf{v}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v}) \mathrm{d} x-\mathcal{L}(\mathbf{v}) & \text { if } \mathbf{v} \in H_{\mathrm{div}}^{1}\left(\Omega, \mathbb{R}^{3}\right)  \tag{2.7}\\ +\infty & \text { otherwise in } W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)\end{cases}
$$

where $\mathbb{E}(\mathbf{v}):=\operatorname{sym} \nabla \mathbf{v}$ denotes the infinitesimal strain tensor field associated to the displacement field $\mathbf{v}$, and where $H_{\mathrm{div}}^{1}\left(\Omega, \mathbb{R}^{3}\right)$ is the set of $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ vector fields whose divergence vanishes a.e. in $\Omega$.

Since we work with the incompressible energy density $\mathcal{W}^{I}$, we stress that if the function $\mathcal{W}$ is replaced by any other $\widetilde{\mathcal{W}}$ such that assumptions $(\mathcal{W} 1),(\mathcal{W} 2),(\mathcal{W} 3),(\mathcal{W} 4)$ are satisfied and $\widetilde{\mathcal{W}}(x, \mathbf{F})=\mathcal{W}(x, \mathbf{F})$ as soon as $\operatorname{det} \mathbf{F}=1$, then this does not affect functional $\mathcal{F}^{I}$. Indeed, if $\operatorname{Tr} \mathbf{B}=0$ then $\operatorname{det}(\exp (h \mathbf{B}))=$ $\exp (h \operatorname{Tr} \mathbf{B})=1$, so that by Taylor's expansion we have

$$
\begin{aligned}
& \frac{1}{2} \operatorname{sym} \mathbf{B} D^{2} \widetilde{\mathcal{W}}(x, \mathbf{I}) \operatorname{sym} \mathbf{B}=\lim _{h \rightarrow 0} h^{-2} \widetilde{\mathcal{W}}(x, \mathbf{I}+h \mathbf{B})=\lim _{h \rightarrow 0} h^{-2} \widetilde{\mathcal{W}}(x, \mathbf{I}+h \mathbf{B}+o(h)) \\
& \quad=\lim _{h \rightarrow 0} h^{-2} \widetilde{\mathcal{W}}(x, \exp (h \mathbf{B}))=\lim _{h \rightarrow 0} h^{-2} \mathcal{W}^{I}(x, \exp (h \mathbf{B}))=\lim _{h \rightarrow 0} h^{-2} \mathcal{W}(x, \exp (h \mathbf{B})) \\
& \quad=\lim _{h \rightarrow 0} h^{-2} \mathcal{W}(x, \mathbf{I}+h \mathbf{B}+o(h))=\lim _{h \rightarrow 0} h^{-2} \mathcal{W}(x, \mathbf{I}+h \mathbf{B})=\frac{1}{2} \operatorname{sym} \mathbf{B} D^{2} \mathcal{W}(x, \mathbf{I}) \operatorname{sym} \mathbf{B}
\end{aligned}
$$

For instance, if the function $\mathcal{W}$ is in the form (1.1), then $\mathcal{W}_{\text {vol }}$ can be arbitrarily replaced as soon as the assumptions $(\mathcal{W} 1),(\mathcal{W} 2),(\mathcal{W} 3),(\mathcal{W} 4)$ are matched.

## Statement of the main result

In order to prescribe a Dirichlet boundary condition on $\Gamma$ we define $\mathcal{G}^{I}: W^{1, p}\left(\Omega, \mathbb{R}^{3}\right) \rightarrow \mathbb{R} \cup\{+\infty\}$ and $\mathcal{G}_{h}^{I}: W^{1, p}\left(\Omega, \mathbb{R}^{3}\right) \rightarrow \mathbb{R} \cup\{+\infty\}$ as

$$
\begin{aligned}
& \mathcal{G}_{h}^{I}(\mathbf{v}):= \begin{cases}\mathcal{F}_{h}^{I}(\mathbf{v}) & \text { if } \mathbf{v}=0 \text { on } \Gamma \\
+\infty & \text { otherwise }\end{cases} \\
& \mathcal{G}^{I}(\mathbf{v}):= \begin{cases}\mathcal{F}^{I}(\mathbf{v}) & \text { if } \mathbf{v}=0 \text { on } \Gamma \\
+\infty & \text { otherwise. }\end{cases}
\end{aligned}
$$

We are ready for the statement of the main result
Theorem 2.1. Assume (2.1), (2.2), ( $\mathcal{W} 1),(\mathcal{W} 2),(\mathcal{W} 3),(\mathcal{W} 4)$. Then for every vanishing sequence $\left(h_{j}\right)_{j \in \mathbb{N}}$ of strictly positive real numbers we have

$$
\begin{equation*}
\inf _{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)} \mathcal{G}_{h_{j}}^{I} \in \mathbb{R} \tag{2.8}
\end{equation*}
$$

If $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}} \subset W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ is a sequence such that $\mathbf{v}_{j}=0$ on $\Gamma$ for any $j \in \mathbb{N}$ and such that

$$
\begin{equation*}
\lim _{j \rightarrow+\infty}\left(\mathcal{G}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right)-\inf _{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)} \mathcal{G}_{h_{j}}^{I}\right)=0 \tag{2.9}
\end{equation*}
$$

then we have as $j \rightarrow+\infty$

$$
\mathbf{v}_{j} \rightharpoonup \mathbf{v}_{*} \text { weakly in } W^{1, p}\left(\Omega, \mathbb{R}^{3}\right),
$$

where $\mathbf{v}_{*}$ is the unique minimizer of $\mathcal{G}^{I}$ over $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$, and

$$
\mathcal{G}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right) \rightarrow \mathcal{G}^{I}\left(\mathbf{v}_{*}\right), \quad \inf _{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)} \mathcal{G}_{h_{j}}^{I} \rightarrow \min _{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)} \mathcal{G}^{I} .
$$

We remark that under our assumptions, functionals $\mathcal{G}_{h}^{I}$ do not have minimizers in general. In case they have, then it is possible to substitute assumption (2.9) with $\mathbf{v}_{j} \in \operatorname{argmin} \mathcal{G}_{h_{j}}^{I}$ and convergence of minimizers is deduced.

## Nonhomogeneous boundary conditions

It is worth noticing that Theorem 2.1 works only when we assume homogeneous boundary conditions so it is quite natural to ask what happens in a more general case. Unfortunately the proof cannot be extended to the non homogeneous case as well, nevertheless this difficulty can be in some sense circumvented as follows. Fix $\overline{\mathbf{v}} \in W^{1, \infty}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\operatorname{div} \overline{\mathbf{v}}=0$ a.e. in $\Omega$ and define for $h>0$

$$
\begin{gather*}
\overline{\mathcal{G}}^{I}(\mathbf{v}):=\left\{\begin{array}{l}
\mathcal{F}^{I}(\mathbf{v}) \text { if } \mathbf{v}=\overline{\mathbf{v}} \text { on } \Gamma \\
+\infty \quad \text { otherwise, }
\end{array}\right.  \tag{2.10}\\
\widetilde{\mathcal{G}}_{h}^{I}(\mathbf{v}):=\mathcal{G}_{h}^{I}(\mathbf{v})+\int_{\Omega} \mathbb{E}(\overline{\mathbf{v}}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v}) \mathrm{d} x+\overline{\mathcal{G}}^{I}(\overline{\mathbf{v}})
\end{gather*}
$$

and

$$
\begin{equation*}
\widetilde{\mathcal{G}}^{I}(\mathbf{v}):=\mathcal{G}^{I}(\mathbf{v})+\int_{\Omega} \mathbb{E}(\overline{\mathbf{v}}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v}) \mathrm{d} x+\overline{\mathcal{G}}^{I}(\overline{\mathbf{v}}) \tag{2.11}
\end{equation*}
$$

A slight modification of the proof of Theorem 2.1 gives the following
Corollary 2.2. Assume (2.1), (2.2), (W) 1 ), (WW), (WW), (W) 4 ) and let $\overline{\mathbf{v}} \in W^{1, \infty}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\operatorname{div} \overline{\mathbf{v}}=$ 0 in $\Omega$. Then for every vanishing sequence $\left(h_{j}\right)_{j \in \mathbb{N}}$ of strictly positive real numbers we have

$$
\begin{equation*}
\left.\inf _{W^{1}, p}\left(\Omega, \mathbb{R}^{3}\right)\right] \widetilde{\mathcal{G}}_{h_{j}}^{I} \in \mathbb{R} . \tag{2.12}
\end{equation*}
$$

If $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}} \subset W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ is a sequence such that $\mathbf{v}_{j}=0$ on $\Gamma$ for any $j \in \mathbb{N}$ and such that

$$
\begin{equation*}
\lim _{j \rightarrow+\infty}\left(\widetilde{\mathcal{G}}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right)-\inf _{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)} \widetilde{\mathcal{G}}_{h_{j}}^{I}\right)=0 \tag{2.13}
\end{equation*}
$$

then we have as $j \rightarrow+\infty$

$$
\mathbf{v}_{j} \rightharpoonup \mathbf{v}_{0} \text { weakly in } W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)
$$

where $\mathbf{v}_{0}$ is the unique minimizer of $\widetilde{\mathcal{G}}^{I}$ over $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$, and

$$
\widetilde{\mathcal{G}}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right) \rightarrow \widetilde{\mathcal{G}}^{I}\left(\mathbf{v}_{0}\right), \quad \inf _{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)} \widetilde{\mathcal{G}}_{h_{j}}^{I} \rightarrow \min _{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)} \widetilde{\mathcal{G}}^{I}
$$

Moreover, $\mathbf{v}_{0}+\overline{\mathbf{v}}$ is the unique minimizer of $\overline{\mathcal{G}}^{I}$ over $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ and $\widetilde{\mathcal{G}}^{I}\left(\mathbf{v}_{0}\right)=\overline{\mathcal{G}}^{I}\left(\mathbf{v}_{0}+\overline{\mathbf{v}}\right)$.

## 3. Preliminary results on vector potentials

Given a divergence-free deformation field $\mathbf{v}$, it will often be useful to work with a vector potential $\mathbf{w}$, such that curlw $=\mathbf{v}$. The next results gather several properties of vector potentials of deformation fields satisfying suitable Dirichlet boundary conditions. We start by recalling a result that is found for instance in [5].
Lemma 3.1. ([5, Corollary 2.15]) Assume (2.1) and let $\mathbf{w} \in L^{2}\left(\Omega, \mathbb{R}^{3}\right)$ be such that curl $\mathbf{w} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$, $\operatorname{div} \mathbf{w} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$, and $\mathbf{v} \cdot \mathbf{n} \in H^{3 / 2}(\partial \Omega)$. Then $\mathbf{w} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$.
Lemma 3.2. Assume (2.1) and let $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ be such that $\operatorname{div} \mathbf{v}=0$ a.e. in $\Omega$. Then there exists $\mathbf{w} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\mathbf{v}=\operatorname{curl} \mathbf{w}$, div $\mathbf{w}=0$ a.e. in $\Omega$ and $\mathbf{w} \cdot \mathbf{n}=0$ on $\partial \Omega$.

Proof. Since $\partial \Omega$ is connected then by [5, Lemma 3.5] there exists $\mathbf{z} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\mathbf{v}=$ $\operatorname{curl} \mathbf{z}, \operatorname{div} \mathbf{z}=0$ a.e. in $\Omega$. By taking into account that

$$
\mathbf{z} \cdot \mathbf{n} \in H^{1 / 2}(\partial \Omega), \quad \int_{\partial \Omega} \mathbf{z} \cdot \mathbf{n} \mathrm{d} \mathcal{H}^{2}=\int_{\Omega} \operatorname{div} \mathbf{z} \mathrm{d} x=0
$$

there exists $\varphi \in H^{1}(\Omega)$ such that

$$
\left\{\begin{array}{l}
\Delta \varphi=0 \text { in } \Omega \\
\frac{\partial \varphi}{\partial \mathbf{n}}=\mathbf{z} \cdot \mathbf{n} \text { in } \partial \Omega .
\end{array}\right.
$$

By setting $\mathbf{w}:=\mathbf{z}-\nabla \varphi$ we get $\mathbf{w} \cdot \mathbf{n}=0$ on $\partial \Omega, \operatorname{div} \mathbf{w}=0, \operatorname{curl} \mathbf{w}=\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ a.e. in $\Omega$ and hence Lemma 3.1 yields $\mathbf{w} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$.

It is well known that for a function $\zeta \in H_{0}^{1}\left(\Omega, \mathbb{R}^{3}\right)$ there holds

$$
\int_{\Omega}|\nabla \boldsymbol{\zeta}|^{2} \mathrm{~d} x=\int_{\Omega}|\operatorname{curl} \boldsymbol{\zeta}|^{2} \mathrm{~d} x+\int_{\Omega}|\operatorname{div} \boldsymbol{\zeta}|^{2} \mathrm{~d} x .
$$

In presence of boundary values we have the following formula that we borrow from [23].
Lemma 3.3. ([23, Theorem 3.1.1.1]) Assume (2.1) and let $\boldsymbol{\zeta} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$. Then

$$
\begin{aligned}
\int_{\Omega}|\nabla \boldsymbol{\zeta}|^{2} d x= & \int_{\Omega}|\operatorname{curl} \boldsymbol{\zeta}|^{2} d x+\int_{\Omega}|\operatorname{div} \boldsymbol{\zeta}|^{2} d x+2\langle\nabla(\boldsymbol{\zeta} \cdot \mathbf{n}) \wedge \mathbf{n}, \boldsymbol{\zeta} \wedge \mathbf{n}\rangle_{\partial \Omega} \\
& -\int_{\partial \Omega}\left\{\operatorname{div} \mathbf{n}(\boldsymbol{\zeta} \cdot \mathbf{n})^{2}+(\boldsymbol{\zeta} \wedge \mathbf{n})^{T} \nabla \mathbf{n}(\boldsymbol{\zeta} \wedge \mathbf{n})\right\} d \mathcal{H}^{2}
\end{aligned}
$$

where $\langle\cdot, \cdot\rangle_{\partial \Omega}$ denotes the duality between $H^{-1 / 2}(\partial \Omega)$ and $H^{1 / 2}(\partial \Omega)$.
Lemma 3.4. Assume (2.1) and (2.2) Let $\left(\boldsymbol{\zeta}_{h}\right)_{h \in \mathbb{N}} \subset H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ be a sequence such that $\boldsymbol{\zeta}_{h} \cdot \mathbf{n}=0$ on $\Gamma$, $\boldsymbol{\zeta}_{h} \wedge \mathbf{n}=0$ on $\partial \Omega \backslash \Gamma$, $\boldsymbol{\zeta}_{h} \rightharpoonup 0$ weakly in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$, curl $\boldsymbol{\zeta}_{h} \rightarrow 0$ in $L^{2}\left(\Omega, \mathbb{R}^{3}\right)$ and div $\boldsymbol{\zeta}_{h} \rightarrow 0$ in $L^{2}(\Omega)$. Then $\zeta_{h} \rightarrow 0$ strongly in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$.
Proof. Since $\boldsymbol{\zeta}_{h} \cdot \mathbf{n}=0$ on $\Gamma$ and $\boldsymbol{\zeta}_{h} \wedge \mathbf{n}=0$ on $\partial \Omega \backslash \Gamma$ we get

$$
\langle\nabla(\boldsymbol{\zeta} \cdot \mathbf{n}) \wedge \mathbf{n}, \boldsymbol{\zeta} \wedge \mathbf{n}\rangle_{\partial \Omega}=0
$$

for any $h \in \mathbb{N}$ and then Lemma 3.3 yields

$$
\int_{\Omega}\left|\nabla \zeta_{h}\right|^{2} \mathrm{~d} x=A_{h}+B_{h}
$$

where

$$
\begin{array}{r}
A_{h}:=\int_{\Omega}\left|\operatorname{curl} \boldsymbol{\zeta}_{h}\right|^{2} \mathrm{~d} x+\int_{\Omega}\left|\operatorname{div} \boldsymbol{\zeta}_{h}\right|^{2} \mathrm{~d} x, \\
B_{h}:=-\int_{\partial \Omega}\left\{\operatorname{div} \mathbf{n}\left(\boldsymbol{\zeta}_{h} \cdot \mathbf{n}\right)^{2}+\left(\boldsymbol{\zeta}_{h} \wedge \mathbf{n}\right)^{T} \nabla \mathbf{n}\left(\boldsymbol{\zeta}_{h} \wedge \mathbf{n}\right)\right\} \mathrm{d} \mathcal{H}^{2} .
\end{array}
$$

By taking into account that $\boldsymbol{\zeta}_{h} \rightharpoonup 0$ weakly in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$, we get $\boldsymbol{\zeta}_{h} \rightarrow 0$ in $L^{2}\left(\partial \Omega, \mathbb{R}^{3}\right)$, hence $B_{h} \rightarrow 0$. Moreover, since both curl $\zeta_{h}$ and div $\zeta_{h}$ go to zero in $L^{2}(\Omega)$, we get also $A_{h} \rightarrow 0$. We conclude that $\left\|\nabla \zeta_{h}\right\|_{L^{2}(\Omega)} \rightarrow 0$ and the result follows from the Poincaré inequality $\left\|\boldsymbol{\zeta}_{h}\right\|_{L^{2}(\Omega)}^{2} \leq C\left(\left\|\nabla \boldsymbol{\zeta}_{h}\right\|_{L^{2}(\Omega)}^{2}+\right.$ $\left.\left\|\boldsymbol{\zeta}_{h}\right\|_{L^{2}(\partial \Omega)}^{2}\right)$.

Always assuming (2.1) and (2.2), we set

$$
\begin{aligned}
& X(\Gamma):=\left\{\boldsymbol{\zeta} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right): \boldsymbol{\zeta} \cdot \mathbf{n}=0 \text { on } \Gamma, \boldsymbol{\zeta} \wedge \mathbf{n}=0 \text { on } \partial \Omega \backslash \Gamma\right\}, \\
& X_{0}(\Gamma):=\{\boldsymbol{\zeta} \in X(\Gamma): \operatorname{div} \boldsymbol{\zeta}=0 \text { in } \Omega\} .
\end{aligned}
$$

Lemma 3.2 ensures that there exist vector fields $\mathbf{w} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ such that div $\mathbf{w}=0$ a.e. in $\Omega, \mathbf{w} \cdot \mathbf{n}=0$ on $\partial \Omega$ and curlw $\cdot \mathbf{n}=0$ on $\Gamma$. Given $\mathbf{w}$ with such properties, for $\boldsymbol{\zeta} \in X(\Gamma)$ we define the functional

$$
\Phi_{\mathbf{w}}(\boldsymbol{\zeta}):=\frac{1}{2} \int_{\Omega}\left(\left.|\operatorname{cur}| \boldsymbol{\zeta}\right|^{2}+|\operatorname{div} \boldsymbol{\zeta}|^{2}\right) \mathrm{d} x-\int_{\partial \Omega}(\mathbf{w} \wedge \boldsymbol{\zeta}) \cdot \mathbf{n} \mathrm{d} \mathcal{H}^{2} .
$$

We prove the following
Lemma 3.5. Assume (2.1) and (2.2) Let $\mathbf{w} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ be such that $\operatorname{div} \mathbf{w}=0$ a.e. in $\Omega$, $\mathbf{w} \cdot \mathbf{n}=0$ on $\partial \Omega$ and curlw $\cdot \mathbf{n}=0$ on $\Gamma$. Then the functional $\Phi_{\mathbf{w}}$ has minimizers both on $X(\Gamma)$ and $X_{0}(\Gamma)$. Moreover,

$$
\begin{equation*}
\min _{X(\Gamma)} \Phi_{\mathbf{w}}=\min _{X_{0}(\Gamma)} \Phi_{\mathbf{w}} . \tag{3.1}
\end{equation*}
$$

Proof. Let $\left(\boldsymbol{\zeta}_{h}\right) \subset X(\Gamma)$ be a sequence such that $\Phi_{\mathbf{w}}\left(\boldsymbol{\zeta}_{h}\right) \rightarrow \inf _{X(\Gamma)} \Phi_{\mathbf{w}}$. We shall first prove that such a sequence is bounded in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$. Indeed, assume by contradiction that, up to subsequences, $\lambda_{h}:=$ $\left\|\boldsymbol{\zeta}_{h}\right\|_{H^{1}\left(\Omega, \mathbb{R}^{3}\right)} \rightarrow+\infty$ and set $\boldsymbol{\xi}_{h}:=\lambda_{h}^{-1} \boldsymbol{\zeta}_{h}$. Then $\boldsymbol{\xi}_{h} \in X(\Gamma),\left\|\boldsymbol{\xi}_{h}\right\|_{H^{1}\left(\Omega, \mathbb{R}^{3}\right)}=1$ and for $h$ large enough

$$
1 \geq \Phi_{\mathbf{w}}\left(\boldsymbol{\zeta}_{h}\right)=\Phi_{\mathbf{w}}\left(\lambda_{h} \boldsymbol{\xi}_{h}\right)=\frac{\lambda_{h}^{2}}{2} \int_{\Omega}\left(\left|\operatorname{curl} \boldsymbol{\xi}_{h}\right|^{2}+\left|\operatorname{div} \boldsymbol{\xi}_{h}\right|^{2}\right) \mathrm{d} x-\lambda_{h} \int_{\partial \Omega}\left(\mathbf{w} \wedge \boldsymbol{\xi}_{h}\right) \cdot \mathbf{n} \mathrm{d} \mathcal{H}^{2} .
$$

Hence, by recalling that $\left\|\boldsymbol{\xi}_{h}\right\|=1$, there exists $\boldsymbol{\xi}^{*} \in X(\Gamma)$ such that, up to subsequences, $\boldsymbol{\xi}_{h} \rightharpoonup \boldsymbol{\xi}^{*}$ weakly in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ and

$$
\int_{\Omega}\left(\left|\operatorname{curl} \boldsymbol{\xi}_{h}\right|^{2}+\left|\operatorname{div} \boldsymbol{\xi}_{h}\right|^{2}\right) \mathrm{d} x \leq \frac{2}{\lambda_{h}} \int_{\partial \Omega}\left(\mathbf{w} \wedge \boldsymbol{\xi}_{h}\right) \cdot \mathbf{n} \mathrm{d} \mathcal{H}^{2}+2 \lambda_{h}^{-2} \rightarrow 0,
$$

that is, $\operatorname{curl} \boldsymbol{\xi}_{h} \rightarrow \operatorname{curl} \boldsymbol{\xi}^{*}=0$ in $L^{2}\left(\Omega, \mathbb{R}^{3}\right)$, div $\boldsymbol{\xi}_{h} \rightarrow \operatorname{div} \boldsymbol{\xi}^{*}=0$ in $L^{2}(\Omega)$ and $\boldsymbol{\xi}^{*} \in X(\Gamma)$. We claim that $\boldsymbol{\xi}^{*}=0$. Indeed let $\boldsymbol{\psi} \in H_{0}^{1}\left(\Omega, \mathbb{R}^{3}\right)$ be such that div $\boldsymbol{\psi}=0$. Since $\Omega$ is simply connected, then by [5, Theorem 3.17] there exists $\boldsymbol{\omega} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\operatorname{curl} \boldsymbol{\omega}=\boldsymbol{\psi}$ a.e. in $\Omega$, div $\boldsymbol{\omega}=0$ a.e. in $\Omega$ and $\boldsymbol{\omega} \wedge \mathbf{n}=0$ on $\partial \Omega$. Therefore

$$
\begin{aligned}
\int_{\Omega} \xi^{*} \cdot \boldsymbol{\psi} \mathrm{~d} x & =\int_{\Omega} \xi^{*} \cdot \operatorname{curl} \boldsymbol{\omega} \mathrm{~d} x=\int_{\Omega} \operatorname{div}\left(\xi^{*} \wedge \boldsymbol{\omega}\right) \mathrm{d} x \\
& =\int_{\partial \Omega}\left(\xi^{*} \wedge \boldsymbol{\omega}\right) \cdot \mathbf{n} \mathrm{d} \mathcal{H}^{2}=\int_{\partial \Omega} \xi^{*} \cdot(\boldsymbol{\omega} \wedge \mathbf{n}) \mathrm{d} \mathcal{H}^{2}=0 .
\end{aligned}
$$

Hence, by invoking for instance [22, Lemma 2.1], there exists $p \in H^{2}(\Omega)$ such that $\nabla p=\boldsymbol{\xi}^{*}$ a.e. in $\Omega$, so that $\Delta p=0$ in $\Omega, \frac{\partial p}{\partial \mathbf{n}}=0$ on $\Gamma$ and $\nabla p \wedge \mathbf{n}=0$ on $\partial \Omega \backslash \Gamma$. In particular, $p$ is constant on $\partial \Omega \backslash \Gamma$ since $\mathbf{n} \wedge(\nabla p \wedge \mathbf{n})$ is its tangential derivative, hence $p$ is constant on the whole $\Omega$ and $\boldsymbol{\xi}^{*}=\nabla p=0$ as claimed. By summarizing: $\boldsymbol{\xi}_{h} \in X(\Gamma), \boldsymbol{\xi}_{h} \rightharpoonup 0$ weakly in $H^{1}\left(\Omega, \mathbb{R}^{3}\right), \operatorname{curl} \boldsymbol{\xi}_{h} \rightarrow 0$ in $L^{2}\left(\Omega ; \mathbb{R}^{3}\right)$, $\operatorname{div} \boldsymbol{\xi}_{h} \rightarrow 0$ in $L^{2}(\Omega)$. Therefore, by Lemma 3.4, $\boldsymbol{\xi}_{h} \rightarrow 0$ strongly in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$, which is a contradiction since we are assuming $\left\|\boldsymbol{\xi}_{h}\right\|_{H^{1}\left(\Omega, \mathbb{R}^{3}\right)}=1$ (see also similar arguments in $[8,10,35,36]$ ).

Since we have shown that $\left(\zeta_{h}\right)$ is a bounded sequence in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ and since $\Phi_{\mathrm{w}}$ is sequentially l.s.c. with respect to the weak convergence in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ we get existence of $\min _{X(\Gamma)} \Phi_{\mathbf{w}}$. It remains to prove that (3.1) holds. Let $\boldsymbol{\zeta}^{*} \in \operatorname{argmin}_{X(\Gamma)} \Phi_{\mathbf{w}}$. Then div $\boldsymbol{\zeta}^{*} \in L^{2}(\Omega)$ and we let $\varphi \in H^{2}(\Omega)$ be solution (see for instance [50, Theorem 3]) to the boundary value problem

$$
\left\{\begin{array}{l}
\Delta \varphi=\operatorname{div} \zeta^{*} \text { in } \Omega \\
\frac{\partial \varphi}{\partial \mathbf{n}}=0 \text { on } \Gamma \\
\varphi=0 \text { on } \partial \Omega \backslash \Gamma
\end{array}\right.
$$

It is readily seen that $\zeta^{*}-\nabla \varphi \in X_{0}(\Gamma)$ and

$$
\Phi_{\mathbf{w}}\left(\zeta^{*}-\nabla \varphi\right)=\Phi_{\mathbf{w}}\left(\zeta^{*}\right)+\int_{\Gamma}(\mathbf{w} \wedge \nabla \varphi) \cdot \mathbf{n} \mathrm{d} \mathcal{H}^{2}=\Phi_{\mathbf{w}}\left(\zeta^{*}\right)+\int_{\partial \Omega}(\mathbf{w} \wedge \nabla \varphi) \cdot \mathbf{n} \mathrm{d} \mathcal{H}^{2}
$$

since, due to $\varphi=0$ on $\partial \Omega \backslash \Gamma$, there holds

$$
\int_{\partial \Omega \backslash \Gamma}(\mathbf{w} \wedge \nabla \varphi) \cdot \mathbf{n} \mathrm{d} \mathcal{H}^{2}=\int_{\partial \Omega \backslash \backslash \Gamma} \mathbf{w} \cdot(\nabla \varphi \wedge \mathbf{n}) \mathrm{d} \mathcal{H}^{2}=0 .
$$

But

$$
\begin{aligned}
\int_{\partial \Omega}(\mathbf{w} \wedge \nabla \varphi) \cdot \mathbf{n} \mathrm{d} \mathcal{H}^{2} & =\int_{\Omega} \operatorname{div}(\mathbf{w} \wedge \nabla \varphi) \mathrm{d} x=-\int_{\Omega} \operatorname{curl} \mathbf{w} \cdot \nabla \varphi \mathrm{d} x \\
& =-\int_{\partial \Omega}(\operatorname{curl} \mathbf{w} \cdot \mathbf{n}) \varphi \mathrm{d} \mathcal{H}^{2}=0
\end{aligned}
$$

since curlw $\mathbf{w}=0$ on $\Gamma$ and $\varphi=0$ on $\partial \Omega \backslash \Gamma$. We conclude that

$$
\Phi_{\mathbf{w}}\left(\zeta^{*}-\nabla \varphi\right)=\Phi_{\mathbf{w}}\left(\zeta^{*}\right)
$$

thus proving the result.
The next result is based on the Euler-Lagrange equation for functional $\Phi_{\mathbf{w}}$. Before its statement, we recall that the reach of a closed set $A \subset \mathbb{R}^{3}$, introduced in [19], is defined by

$$
\begin{equation*}
R(A):=\sup \{r>0: 0<d(x, A)<r \Rightarrow \exists!y \in A \text { s.t. } d(x, y)=d(x, A)\} \tag{3.2}
\end{equation*}
$$

where the distance function is defined on $\mathbb{R}^{3}$ by $d(x, A):=\inf _{y \in A}|x-y|$. It is well-known that $R(A)>0$ whenever $A$ is a $C^{2}$ compact $1 D$ or $2 D$ manifold without boundary, see for instance [53,54].

Lemma 3.6. Assume (2.1) and (2.2). Let $\mathbf{w}$ as in Lemma 3.5 and let $\boldsymbol{\zeta}_{*} \in \operatorname{argmin}_{X_{0}(\Gamma)} \Phi_{\mathbf{w}}$. Then we have

$$
\begin{equation*}
\int_{\Omega} \operatorname{curl} \zeta_{*} \cdot \operatorname{curl} \varphi d x=0 \quad \forall \varphi \in C_{0}^{1}\left(\Omega, \mathbb{R}^{3}\right) \tag{3.3}
\end{equation*}
$$

and

$$
\begin{equation*}
\left\langle\operatorname{curl} \boldsymbol{\zeta}_{*} \wedge \mathbf{n}, \varphi\right\rangle_{\partial \Omega}=-\int_{\Gamma}(\mathbf{w} \wedge \mathbf{n}) \cdot \varphi d \mathcal{H}^{2} \quad \forall \varphi \in C^{1}\left(\bar{\Omega}, \mathbb{R}^{3}\right) \text { s.t. } \varphi \equiv 0 \text { on } \partial \Omega \backslash \Gamma . \tag{3.4}
\end{equation*}
$$

Proof. If $\boldsymbol{\zeta}_{*} \in \operatorname{argmin}_{X_{0}(\Gamma)} \Phi_{\mathbf{w}}$, then by Lemma 3.5 we have $\boldsymbol{\zeta}_{*} \in \operatorname{argmin}_{X(\Gamma)} \Phi_{\mathbf{w}}$. If $\varphi \in C^{1}\left(\bar{\Omega}, \mathbb{R}^{3}\right)$ is such that $\boldsymbol{\varphi} \wedge \mathbf{n}=0$ on $\partial \Omega \backslash \Gamma$ and $\boldsymbol{\varphi} \cdot \mathbf{n}=0$ on $\Gamma$, for any $\varepsilon \in(-1,1)$ we have $\boldsymbol{\zeta}_{*}+\varepsilon \boldsymbol{\varphi} \in X(\Gamma)$. Hence, the following first order condition holds

$$
\begin{equation*}
\int_{\Omega} \operatorname{curl} \boldsymbol{\zeta}_{*} \cdot \operatorname{curl} \varphi \mathrm{~d} x=\int_{\partial \Omega}(\mathbf{w} \wedge \varphi) \cdot \mathbf{n d} \mathcal{H}^{2} \quad \forall \varphi \in C^{1}\left(\bar{\Omega}, \mathbb{R}^{3}\right) \cap X(\Gamma) . \tag{3.5}
\end{equation*}
$$

Choosing in particular test functions $\varphi$ that vanish on $\partial \Omega$, we deduce (3.3).
Let now $0<\delta<R(\partial \Omega)$, let $\Omega_{\delta}:=\{x \in \Omega: d(x, \partial \Omega)<\delta\}$, so that for any $x \in \Omega$ there is a unique projection $\sigma(x)$ of $x$ on $\partial \Omega$, and set $\mathbf{n}(x):=\mathbf{n}(\sigma(x))$ for every $x \in \Omega_{\delta}$. Let $\eta_{\delta} \in C^{1}(\bar{\Omega})$ be a cutoff function such that $\eta_{\delta} \equiv 1$ in $\Omega_{\delta / 2}$ and $\eta_{\delta} \equiv 0$ in $\Omega \backslash \Omega_{\delta}$. Moreover, for every $\varphi \in C^{1}\left(\bar{\Omega}, \mathbb{R}^{3}\right)$ such that $\varphi \equiv 0$ on $\partial \Omega \backslash \Gamma$, we define $\boldsymbol{\varphi}_{\delta}:=\eta_{\delta} \varphi$. We take advantage of the cutoff function to define $\boldsymbol{\varphi}_{\delta} \cdot \mathbf{n}$ on the whole $\Omega$ even if $\mathbf{n}$ is defined only on $\Omega_{\delta}$, as follows

$$
\left(\varphi_{\delta} \cdot \mathbf{n}\right)(x):=\left\{\begin{array}{l}
\varphi_{\delta}(x) \cdot \mathbf{n}(x) \text { if } x \in \Omega_{\delta} \\
0 \text { otherwise in } \Omega
\end{array}\right.
$$

It is readily seen that $\boldsymbol{\theta}_{\delta}:=\boldsymbol{\varphi}_{\delta}-\left(\boldsymbol{\varphi}_{\delta} \cdot \mathbf{n}\right) \mathbf{n} \in X(\Gamma) \cap C^{1}\left(\bar{\Omega}, \mathbb{R}^{3}\right)$ and therefore we can make use of (3.5) and get

$$
\begin{equation*}
\int_{\Omega} \operatorname{curl} \boldsymbol{\zeta}_{*} \cdot \operatorname{curl} \boldsymbol{\theta}_{\delta} \mathrm{d} x=-\int_{\Gamma}(\mathbf{w} \wedge \mathbf{n}) \cdot \boldsymbol{\theta}_{\delta} \mathrm{d} \mathcal{H}^{2}=-\int_{\Gamma}(\mathbf{w} \wedge \mathbf{n}) \cdot \boldsymbol{\varphi} \mathrm{d} \mathcal{H}^{2} \tag{3.6}
\end{equation*}
$$

A density argument shows that

$$
\begin{equation*}
\left\langle\operatorname{curl} \boldsymbol{\zeta}_{*} \wedge \mathbf{n},(\boldsymbol{\varphi} \cdot \mathbf{n}) \mathbf{n}\right\rangle_{\partial \Omega}=0 . \tag{3.7}
\end{equation*}
$$

On the other hand, since (3.3) shows that curl curl $\boldsymbol{\zeta}_{*}=0$ a.e. in $\Omega$, integration by parts entails

$$
\begin{equation*}
\int_{\Omega} \operatorname{curl} \boldsymbol{\zeta}_{*} \cdot \operatorname{curl} \boldsymbol{\theta}_{\delta} \mathrm{d} x=\left\langle\operatorname{curl} \boldsymbol{\zeta}_{*} \wedge \mathbf{n}, \boldsymbol{\varphi}-(\boldsymbol{\varphi} \cdot \mathbf{n}) \mathbf{n}\right\rangle_{\partial \Omega} . \tag{3.8}
\end{equation*}
$$

Combining (3.6), (3.7) and (3.8) yields the result.
For a divergence-free deformation field $\mathbf{v}$ that vanishes on $\Gamma$, taking advantage of the latter results we can construct a vector potential $\widetilde{\mathbf{w}}$ that vanishes on $\Gamma$ as well.

Lemma 3.7. Assume (2.1) and (2.2). Let $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ such that div $\mathbf{v}=0$ a.e. in $\Omega$ and $\mathbf{v}=0$ on $\Gamma$. Then there exists $\widetilde{\mathbf{w}} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\widetilde{\mathbf{w}}=0$ on $\Gamma$ and curl $\widetilde{\mathbf{w}}=\mathbf{v}$ a.e. in $\Omega$.

Proof. Let $\boldsymbol{\zeta}_{*}$ as in Lemma 3.6. Then curlcurl $\boldsymbol{\zeta}_{*}=0$ a.e. in $\Omega$ and by arguing as in Lemma 3.5 with the help [22, Lemma 2.1] there exists $\theta \in H^{1}(\Omega)$ such that $-\nabla \theta=\operatorname{curl} \zeta_{*}$ a.e. in $\Omega$, hence by (3.4) there holds $\nabla \theta \wedge \mathbf{n}=\mathbf{w} \wedge \mathbf{n}$ in the sense of $H^{-1 / 2}\left(\Gamma ; \mathbb{R}^{3}\right)$. By taking into account that $\mathbf{w} \in H^{3 / 2}\left(\partial \Omega, \mathbb{R}^{3}\right)$ and that $\mathbf{n} \wedge(\nabla \theta \wedge \mathbf{n})$ is the (weak) tangential gradient of $\theta$ on $\Gamma$ we get $\theta \in H^{5 / 2}(\Gamma)$, so there exists $\widetilde{\theta} \in H^{5 / 2}(\partial \Omega)$ such that $\widetilde{\theta}=\theta$ on $\Gamma$ (thanks to the regularity of $\partial \Gamma$ ). Let now $\psi_{*}$ be the unique solution
to the biharmonic boundary value problem

$$
\left\{\begin{array}{l}
\Delta^{2} \psi=0 \text { in } \Omega \\
\psi=\widetilde{\theta} \text { on } \partial \Omega \\
\frac{\partial \psi}{\partial \mathbf{n}}=0 \text { on } \partial \Omega
\end{array}\right.
$$

Since $\widetilde{\theta} \in H^{5 / 2}(\partial \Omega)$ then $\psi_{*} \in H^{3}(\Omega)$ so by setting $\widetilde{\mathbf{w}}:=\mathbf{w}-\nabla \psi_{*} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ we get $\widetilde{\mathbf{w}} \wedge \mathbf{n}=$ $\left(\mathbf{w}-\nabla \psi_{*}\right) \wedge \mathbf{n}=\mathbf{w} \wedge \mathbf{n}-\nabla \theta \wedge \mathbf{n}=0$ on $\Gamma$ and $\operatorname{curl} \widetilde{\mathbf{w}}=\operatorname{curl} \mathbf{w}=\mathbf{v}$ a.e. in $\Omega$ and thesis follows by recalling that $\mathbf{w} \cdot \mathbf{n}=0$ on the whole $\partial \Omega$.

Lastly, we prove a property of traces of $H^{2}\left(\Omega ; \mathbb{R}^{3}\right)$ functions which will be extremely useful in the proof of our main result.

Lemma 3.8. Assume (2.1) and (2.2). Let $\mathbf{w} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\mathbf{w}=0$ on $\Gamma$. Then

$$
\begin{equation*}
\frac{\partial}{\partial \mathbf{n}}(\mathbf{w} \wedge \mathbf{n})=0 \quad \text { on } \Gamma \tag{3.9}
\end{equation*}
$$

if and only if

$$
\operatorname{curl} \mathbf{w}=0 \quad \text { on } \Gamma \text {. }
$$

Proof. As in the proof of Lemma 3.6 we define here $\mathbf{n}(x):=\mathbf{n}(\sigma(x))$ for every $x$ in a small neighborhood $\Omega_{\delta}$ of $\partial \Omega$, being $\sigma(x)$ the unique projection of $x$ on $\partial \Omega$. In particular, curl $\mathbf{n}=0$ in $\overline{\Omega_{\delta}}$ so that $\partial_{k} \mathbf{n}=\nabla n_{k}$ on $\overline{\Omega_{\delta}}$ for any $k \in\{1,2,3\}$. Suppose first that curl $\mathbf{w}=0$ on $\Gamma$. Then we have on $\Gamma$

$$
\begin{equation*}
\frac{\partial}{\partial \mathbf{n}}(\mathbf{w} \wedge \mathbf{n})=\sum_{k=1}^{3} n_{k} \partial_{k}(\mathbf{w} \wedge \mathbf{n})=\sum_{k=1}^{3}\left(n_{k} \partial_{k} \mathbf{w} \wedge \mathbf{n}+n_{k} \mathbf{w} \wedge \partial_{k} \mathbf{n}\right)=\sum_{k=1}^{3} n_{k} \nabla \mathbf{w}_{k} \wedge \mathbf{n} \tag{3.10}
\end{equation*}
$$

since $\sum_{k=1}^{3} n_{k} \nabla n_{k}=0$. But $\nabla \mathbf{w}_{k} \wedge \mathbf{n}=0$ on $\Gamma$ for any $k \in\{1,2,3\}$ since $\mathbf{w}=0$ on $\Gamma$, implying that the tangential derivative of $\mathbf{w}$ vanishes on $\Gamma$. Conversely if (3.9) holds, by arguing as in (3.10) with the help of $\nabla \mathbf{w}_{k} \wedge \mathbf{n}=0$ on $\Gamma$ we get

$$
\begin{equation*}
\sum_{k=1}^{3} n_{k}\left(\partial_{k} \mathbf{w}-\nabla \mathbf{w}_{k}\right) \wedge \mathbf{n}=0 \quad \text { on } \Gamma . \tag{3.11}
\end{equation*}
$$

Since by symmetry $\sum_{i=1}^{3} \sum_{k=1}^{3} n_{i} n_{k}\left(\partial_{k} \mathbf{w}_{i}-\partial_{i} \mathbf{w}_{k}\right)=0$, from (3.11) we get

$$
\sum_{k=1}^{3} n_{k}\left(\partial_{k} \mathbf{w}-\nabla \mathbf{w}_{k}\right)=0 \text { on } \Gamma .
$$

Therefore, fixing $k \in\{1,2,3\}$,

$$
\begin{aligned}
\partial_{k} \mathbf{w}-\nabla \mathbf{w}_{k} & =\sum_{i=1}^{3}\left(\partial_{k} \mathbf{w}_{i}-\partial_{i} \mathbf{w}_{k}\right) n_{i} \mathbf{n}+\mathbf{n} \wedge\left(\left(\partial_{k} \mathbf{w}-\nabla \mathbf{w}_{k}\right) \wedge \mathbf{n}\right)=\mathbf{n} \wedge\left(\left(\partial_{k} \mathbf{w}-\nabla \mathbf{w}_{k}\right) \wedge \mathbf{n}\right) \\
& =\mathbf{n} \wedge\left(\partial_{k} \mathbf{w} \wedge \mathbf{n}\right)=\mathbf{n} \wedge\left(\partial_{k}(\mathbf{w} \wedge \mathbf{n})-\mathbf{w} \wedge \partial_{k} \mathbf{n}\right)=0
\end{aligned}
$$

on $\Gamma$, where the latter equality is due to the fact that $\nabla(\mathbf{w} \wedge \mathbf{n})=0$ on $\Gamma$ (since (3.9) holds and since $\mathbf{w} \wedge \mathbf{n}$ vanishes on $\Gamma$ ). The result is proven.

## 4. Approximation results

Several approximation results will be needed for obtaining $\Gamma$-convergence and for giving the proof of the main theorem. The first one is contained in the next lemma, which is a consequence of the well known Reynolds' Transport Theorem. We will prove it in some details since its application in this context seems a novelty, at least to our present knowledge.
Lemma 4.1. Let $\Omega$ be a bounded open subset of $\mathbb{R}^{3}$ and let $\Gamma \subseteq \partial \Omega$. Let $\Omega^{\prime} \subset \mathbb{R}^{3}$ be an open set such that $\bar{\Omega} \subset \Omega^{\prime}$. Let $\mathbf{v} \in C^{1}\left(\overline{\Omega^{\prime}} ; \mathbb{R}^{3}\right)$ be such that $\operatorname{div} \mathbf{v}=0$ in $\Omega^{\prime}$ and $\mathbf{v}=0$ on $\Gamma$. Then, for every sequence $\left(h_{j}\right)_{j \in \mathbb{N}}$ of strictly positive numbers such that $\lim _{j \rightarrow \infty} h_{j}=0$, there exists a sequence $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}} \subset C^{1}\left(\bar{\Omega}, \mathbb{R}^{3}\right)$ such that

$$
\begin{align*}
& \operatorname{det}\left(\mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right)=1 \text { in } \Omega,  \tag{4.1}\\
& \mathbf{v}_{j}=0 \text { on } \Gamma,  \tag{4.2}\\
& \mathbf{v}_{j} \rightarrow \mathbf{v} \text { in } W^{1, p}\left(\Omega, \mathbb{R}^{3}\right) \quad \forall p \in[1,+\infty),  \tag{4.3}\\
& \left\|h_{j} \nabla \mathbf{v}_{j}\right\|_{L^{\infty}\left(\Omega, \mathbb{R}^{3 \times 3}\right)} \rightarrow 0 . \tag{4.4}
\end{align*}
$$

Proof. Let $\Omega_{*}$ be an open set, compactly contained in $\Omega^{\prime}$, such that $\bar{\Omega} \subset \Omega_{*}$. We choose $T \in(0,1)$ small enough, such that $\mathbf{y}(t, x) \in \Omega^{\prime}$ for any $x \in \overline{\Omega_{*}}$ and any $t \in[0, T]$, where $\mathbf{y}(\cdot, x)$ is the unique solution to

$$
\left\{\begin{array}{l}
\frac{\partial \mathbf{y}}{\partial t}(t, x)=\mathbf{v}(\mathbf{y}(t, x)), \quad t \in(0, T]  \tag{4.5}\\
\mathbf{y}(0, x)=x
\end{array}\right.
$$

Classical results show that since $\mathbf{v} \in C^{1}\left(\Omega^{\prime} ; \mathbb{R}^{3}\right)$, then $\mathbf{y} \in C^{1}\left([0, T] \times \overline{\Omega_{*}}\right)$. In particular, given a measurable set $A \subset \Omega$, for any $t \in[0, T]$ we have $A_{t}:=\mathbf{y}(t, A) \subset \Omega^{\prime}$ and (see also [25, Corollary 5.2.8, Corollary 5.2.10])

$$
\frac{\mathrm{d}}{\mathrm{~d} t}\left|A_{t}\right|=\int_{A_{t}} \operatorname{div} \mathbf{v}(x) \mathrm{d} x=0
$$

hence $\left|A_{t}\right|=|A|$ for any $t \in[0, T]$. Therefore, using the change of variables formula

$$
\left|A_{t}\right|=\int_{A} \operatorname{det} \nabla \mathbf{y}(t, x) \mathrm{d} x
$$

we conclude that for any $t \in[0, T]$

$$
\int_{A} \mathrm{~d} x=\int_{A} \operatorname{det} \nabla \mathbf{y}(t, x) \mathrm{d} x .
$$

By the arbitrariness of the measurable set $A \subset \Omega$, for any $t \in[0, T]$ we get

$$
\begin{equation*}
\operatorname{det} \nabla \mathbf{y}(t, x)=1 \quad \text { for every } x \in \Omega \tag{4.6}
\end{equation*}
$$

Assuming wlog that $h_{j}<T$, we define

$$
\mathbf{y}_{j}(x):=\mathbf{y}\left(h_{j}, x\right), \quad \mathbf{v}_{j}(x):=h_{j}^{-1}\left(\mathbf{y}_{j}(x)-x\right), \quad x \in \Omega_{*} .
$$

By taking into account that $\mathbf{v}=0$ on $\Gamma \subset \Omega_{*}$ we get $\mathbf{y}_{j}(x) \equiv x$ on $\Gamma$ so $\mathbf{v}_{j}$ vanishes on $\Gamma$ and (4.6) entails $\operatorname{det}\left(\mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right)=1$ in $\Omega$, thus proving (4.1) and (4.2).

We next prove (4.3). Let $t \in(0, T]$. We notice that from (4.5) we get

$$
\begin{equation*}
\frac{1}{t}(\mathbf{y}(t, x)-x)-\mathbf{v}(x)=\frac{1}{t} \int_{0}^{t}(\mathbf{v}(\mathbf{y}(s, x))-\mathbf{v}(x)) \mathrm{d} s \tag{4.7}
\end{equation*}
$$

and thus

$$
\frac{1}{t}|\mathbf{y}(t, x)-x| \leq|\mathbf{v}(x)|+\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)} \int_{0}^{t} \frac{1}{s}|\mathbf{y}(s, x)-x| \mathrm{d} s
$$

for any $x \in \Omega$, and Gronwall lemma entails

$$
\frac{1}{t}|\mathbf{y}(t, x)-x| \leq|\mathbf{v}(x)| \exp \left\{\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)} t\right\} \leq C_{\mathbf{v}}
$$

where $C_{\mathbf{v}}:=\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)} \exp \left\{\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)}\right\}$. From the definition of $\mathbf{v}_{j}$, from (4.7) and from the latter estimate we obtain

$$
\begin{aligned}
\left|\mathbf{v}_{j}(x)-\mathbf{v}(x)\right| & =\left|\frac{1}{h_{j}}\left(\mathbf{y}\left(h_{j}, x\right)-x\right)-\mathbf{v}(x)\right| \leq\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)} \int_{0}^{h_{j}} \frac{1}{s}|\mathbf{y}(s, x)-x| \mathrm{d} s \\
& \leq C_{\mathbf{v}}\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)} h_{j}
\end{aligned}
$$

for any $x \in \Omega$ and any $j \in \mathbb{N}$. From the latter we get the convergence of $\mathbf{v}_{j}$ to $\mathbf{v}$ in $L^{1} \cap L^{\infty}(\Omega)$ as $j \rightarrow 0$.
We take now the gradient in (4.7), and since the map $\Omega_{*} \ni x \mapsto \mathbf{v}(\mathbf{y}(t, x))$ is Lipschitz continuous for any $t \in(0, T]$ we may take the gradient under integral sign and obtain

$$
\begin{align*}
& \frac{1}{t}(\nabla \mathbf{y}(t, x)-\mathbf{I})-\nabla \mathbf{v}(x)=\frac{1}{t} \int_{0}^{t}(\nabla[\mathbf{v}(\mathbf{y}(s, x))]-\nabla \mathbf{v}(x)) \mathrm{d} s  \tag{4.8}\\
& \quad=\frac{1}{t} \int_{0}^{t}(\nabla \mathbf{v}(\mathbf{y}(s, x)) \nabla \mathbf{y}(s, x)-\nabla \mathbf{v}(x) \nabla \mathbf{y}(s, x)) \mathrm{d} s+\frac{1}{t} \int_{0}^{t}(\nabla \mathbf{v}(x) \nabla \mathbf{y}(s, x)-\nabla \mathbf{v}(x)) \mathrm{d} s
\end{align*}
$$

for every $x \in \Omega$. Form the first equality of (4.8) we get

$$
\begin{align*}
\frac{1}{t}\|\nabla \mathbf{y}(t, \cdot)-\mathbf{I}\|_{L^{\infty}\left(\Omega_{*}\right)} & =\frac{1}{t}\left\|\int_{0}^{t} \nabla \mathbf{v}(\mathbf{y}(s, \cdot)) \nabla \mathbf{y}(s, \cdot) \mathrm{d} s\right\|_{L^{\infty}\left(\Omega_{*}\right)}  \tag{4.9}\\
& \leq\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)} \sup _{t \in[0, T]}\|\nabla \mathbf{y}(t, \cdot)\|_{L^{\infty}\left(\Omega_{*}\right)}=\mathcal{Q}\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)}
\end{align*}
$$

for any $t \in(0, T]$, where $\mathcal{Q}:=\|\nabla \mathbf{y}\|_{C^{0}\left([0, T] \times \overline{\Omega_{*}}\right)}<+\infty$. Still from the first equality of (4.8) we have

$$
\begin{aligned}
\left|\nabla \mathbf{v}_{j}(x)-\nabla \mathbf{v}(x)\right| & =\left|\frac{1}{h_{j}}\left(\nabla \mathbf{y}\left(h_{j}, x\right)-\mathbf{I}\right)-\nabla \mathbf{v}(x)\right| \\
& \leq \int_{0}^{h_{j}} \frac{|\nabla \mathbf{y}(s, x)|}{h_{j}}|\nabla \mathbf{v}(\mathbf{y}(s, x))-\nabla \mathbf{v}(x)| \mathrm{d} s+\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)} \int_{0}^{h_{j}} \frac{|\nabla \mathbf{y}(s, x)-\mathbf{I}|}{s} \mathrm{~d} s,
\end{aligned}
$$

for every $x \in \Omega$ and any $j \in \mathbb{N}$. Hence, (4.9) entails

$$
\begin{equation*}
\left|\nabla \mathbf{v}_{j}(x)-\nabla \mathbf{v}(x)\right| \leq \frac{\mathcal{Q}}{h_{j}} \int_{0}^{h_{j}}|\nabla \mathbf{v}(\mathbf{y}(s, x))-\nabla \mathbf{v}(x)| \mathrm{d} s+\mathcal{Q}\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)} h_{j} \tag{4.10}
\end{equation*}
$$

for every $x \in \Omega$ and any $j \in \mathbb{N}$. We notice that for $p \in[1,+\infty)$, by Jensen inequality there holds

$$
\begin{aligned}
\frac{1}{h_{j}^{p}}\left(\int_{0}^{h_{j}}|\nabla \mathbf{v}(\mathbf{y}(s, x))-\nabla \mathbf{v}(x)| \mathrm{d} s\right)^{p} & \leq \frac{1}{h_{j}} \int_{0}^{h_{j}}|\nabla \mathbf{v}(\mathbf{y}(s, x))-\nabla \mathbf{v}(x)|^{p} \mathrm{~d} s \\
& \leq 2\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)}^{p-1} \frac{1}{h_{j}} \int_{0}^{h_{j}}|\nabla \mathbf{v}(\mathbf{y}(s, x))-\nabla \mathbf{v}(x)| \mathrm{d} s
\end{aligned}
$$

and the above right hand side vanishes for any $x \in \Omega$ as $j \rightarrow+\infty$ since $\nabla \mathbf{v}$ is continuous by assumption, so that we obtain

$$
\begin{equation*}
\lim _{j \rightarrow \infty} \frac{1}{h_{j}^{p}} \int_{\Omega}\left(\int_{0}^{h_{j}}|\nabla \mathbf{v}(\mathbf{y}(s, x))-\nabla \mathbf{v}(x)| \mathrm{d} s\right)^{p} \mathrm{~d} x=0 \tag{4.11}
\end{equation*}
$$

by dominated convergence, using $2\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)}^{p}$ as dominating function on the bounded domain $\Omega$. From (4.10) we find

$$
\int_{\Omega}\left|\nabla \mathbf{v}_{j}(x)-\nabla \mathbf{v}(x)\right|^{p} \mathrm{~d} x \leq \frac{\mathcal{Q}^{p}}{h_{j}^{p}} \int_{\Omega}\left(\int_{0}^{h_{j}}|\nabla \mathbf{v}(\mathbf{y}(s, x))-\nabla \mathbf{v}(x)| \mathrm{d} s\right)^{p} \mathrm{~d} x+|\Omega| \mathcal{Q}^{p}\|\mathbf{v}\|_{W^{1, \infty}\left(\Omega^{\prime}\right)}^{p} h_{j}^{p}
$$

so that the $L^{p}(\Omega)$ convergence of $\nabla \mathbf{v}_{j}$ to $\nabla \mathbf{v}$ follows by taking the limit as $j \rightarrow+\infty$ and by using (4.11). This concludes the proof of (4.3).

Eventually, since $\nabla \mathbf{v}_{j}(x)=\frac{1}{h_{j}}\left(\nabla \mathbf{y}\left(h_{j}, x\right)-\mathbf{I}\right)$, (4.4) directly follows from (4.9).
The next step is an approximation of divergence-free $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ vector fields with divergence-free $C^{1}\left(\bar{\Omega}, \mathbb{R}^{3}\right)$ vector fields, in presence of suitable vanishing conditions on subsets of $\partial \Omega$. It is stated in Lemma 4.4. It requires the introduction of some notation about normal bundles and a couple of preliminary lemmas. From here and through the rest of the paper, assumptions (2.1) and (2.2) are always understood to hold.

Recalling the definition of reach from (3.2), with the convention $R(\emptyset)=+\infty$, let

$$
\begin{equation*}
\mu_{0}:=\frac{1}{2} \min \{R(\partial \Omega), R(\partial \Gamma)\} \tag{4.12}
\end{equation*}
$$

Remark 4.2. (Regularity of the squared distance function) Assume (2.1) and (2.2). Let either $A=\partial \Gamma \neq \emptyset$ or $A=\partial \Omega$. The distance function $d(x, A):=\min _{y \in A}|x-y|$ is differentiable at any point $x \in \mathbb{R}^{3}$ such that $0<d(x, A)<R(A)$, see [18, Theorem 3.3, Chapter 6]. In particular, the squared distance function $d^{2}(\cdot, A)$ inherits the $C^{3}$ regularity of $A$ in the tubular neighbor $U_{0}(A):=\left\{x \in \mathbb{R}^{3}: d(x, A)<\mu_{0}\right\}$, see for instance [33, Proposition 4.6], see also [18, Theorem 6.5, Chapter 6]. We deduce $d(\cdot, A) \in C^{3}\left(U_{0}(A) \backslash A\right)$.

For every $0<\mu<\mu_{0}$, let

$$
\begin{equation*}
S_{\mu}:=\{\sigma+t \mathbf{n}(\sigma): \sigma \in \Gamma,|t| \leq \mu\} . \tag{4.13}
\end{equation*}
$$

We further define, for any $0<\mu<\mu_{0}$ and any $0<\delta<\mu_{0}$,

$$
\begin{equation*}
\Gamma_{\delta}:=\{x \in \partial \Omega: d(x, \Gamma) \leq \delta\}, \quad S_{\mu, \delta}:=\left\{\sigma+\operatorname{tn}(\sigma): \sigma \in \Gamma_{\delta},|t| \leq \mu\right\} . \tag{4.14}
\end{equation*}
$$

In case $\Gamma=\partial \Omega$ (i.e., $\partial \Gamma=\emptyset$ ), we have $\Gamma_{\delta} \equiv \Gamma$ and $S_{\mu, \delta} \equiv S_{\mu}$, for any $0<\delta<\mu_{0}$. We stress that this case in encoded in Lemmas 4.3 and 4.4 below. On the other hand, if $\partial \Gamma \neq \emptyset$, then assumption (2.1) and the regularity properties of the distance function from $\partial \Gamma$ (see Remark 4.2) imply that $\partial \Gamma_{\delta}$ is a $C^{3}$ one-dimensional submanifold of $\partial \Omega$. In particular, $\Gamma_{\delta}$ itself satisfies assumption (2.2).

Let us also introduce the following notation for neighbors of $\Omega$ and $\partial \Omega$

$$
\begin{equation*}
\Omega^{\prime}:=\left\{x \in \mathbb{R}^{3}: d(x, \Omega)<\mu_{0}\right\} \quad \text { and } \quad \Omega^{\prime \prime}:=\left\{x \in \mathbb{R}^{3}: d(x, \partial \Omega)<\mu_{0}\right\} \subset \Omega^{\prime} . \tag{4.15}
\end{equation*}
$$

We notice that $\mathbf{n}$ can be extended to $\Omega^{\prime \prime}$ in the usual way: $\mathbf{n}(x)=\mathbf{n}(\sigma(x))$, where $\sigma(x)$ is the unique projection of $x \in \Omega^{\prime \prime}$ on $\partial \Omega$, and therefore $\mathbf{n}$ is a $C^{2}\left(\Omega^{\prime \prime}, \mathbb{R}^{3}\right)$ vector field, so that there exists $K>1$ such that

$$
\begin{equation*}
|\nabla \mathbf{n}|+\left|\nabla^{2} \mathbf{n}\right| \leq K \text { in } \Omega^{\prime \prime} \tag{4.16}
\end{equation*}
$$

Some auxiliary estimates are given by the next
Lemma 4.3. Assume (2.1) and (2.2). Let $0<\delta<\mu_{0}$, where $\mu_{0}$ is defined by (4.12). Let $\boldsymbol{f} \in H^{2}\left(\Omega^{\prime \prime}, \mathbb{R}^{3}\right)$ be such that $\boldsymbol{f}=0$ on $\Gamma_{\delta}$.

Then there exists $\varepsilon_{0} \in\left(0, \mu_{0}\right)$ such that for any $\varepsilon \in\left(0, \varepsilon_{0}\right)$ and any $\lambda \in(0, \delta)$ there holds

$$
\begin{equation*}
\int_{S_{\varepsilon, \lambda}}|\boldsymbol{f}|^{2} d x \leq 2 \varepsilon^{2} \int_{S_{\varepsilon, \lambda}}|\nabla \boldsymbol{f}|^{2} d x \tag{4.17}
\end{equation*}
$$

and if $\frac{\partial \boldsymbol{f}}{\partial \mathbf{n}}=0$ on $\Gamma_{\delta}$ as well, there holds

$$
\begin{equation*}
\int_{S_{\varepsilon, \lambda}}|\boldsymbol{f}|^{2} d x \leq \frac{\varepsilon^{4}}{2} \int_{S_{\varepsilon, \lambda}}\left|\nabla^{2} \boldsymbol{f}\right|^{2} d x . \tag{4.18}
\end{equation*}
$$

Proof. Let $B$ denote the unit ball in $\mathbb{R}^{2}$ and let $\psi \in C^{3}\left(\bar{B} ; \mathbb{R}^{3}\right)$ be any local chart parametrizing a subset of $\partial \Omega$. Let $B_{\lambda}:=\boldsymbol{\psi}^{-1}\left(\Gamma_{\lambda} \cap \boldsymbol{\psi}(B)\right)$. Let $B \times(-\varepsilon, \varepsilon) \ni(u, t) \mapsto \boldsymbol{\Phi}(u, t):=\boldsymbol{\psi}(u)+t \mathbf{n}(\boldsymbol{\psi}(u))$. Up to covering $\Gamma_{\lambda}$ with local charts, it is enough to show that (4.17) and (4.18) hold, for suitably small $\varepsilon$, with $\boldsymbol{\Phi}\left(B_{\lambda} \times(-\varepsilon, \varepsilon)\right)$ in place of $S_{\varepsilon, \lambda}$. We have $|\operatorname{det} D \boldsymbol{\Phi}(u, 0)|=\left|\partial_{1} \boldsymbol{\psi}(u) \wedge \partial_{2} \boldsymbol{\psi}(u)\right|>0$, where $D$ denotes the gradient in the variables $(u, t)$. $|\operatorname{det} D \boldsymbol{\Phi}(u, 0)|$ is bounded away from zero on $B$ and $|\operatorname{det} D \boldsymbol{\Phi}(u, t)|=\left|\partial_{1} \boldsymbol{\psi}(u) \wedge \partial_{2} \boldsymbol{\psi}(u)\right|+o(1)$ as $t \rightarrow 0$, uniformly with respect to $u \in B$. We notice that $|D \boldsymbol{\Phi}|$ is bounded on $B \times(-\varepsilon, \varepsilon)$. Moreover, it is not difficult to check that for any small enough $\varepsilon$ there holds

$$
\begin{equation*}
\frac{|\operatorname{det} D \boldsymbol{\Phi}(u, t)|}{|\operatorname{det} D \boldsymbol{\Phi}(u, s)|} \leq 2 \quad \text { for any }(u, t, s) \in B \times(-\varepsilon, \varepsilon) \times(-\varepsilon, \varepsilon) \text {. } \tag{4.19}
\end{equation*}
$$

By the properties of Sobolev functions (see for instance [40, Chapter 1]), $\boldsymbol{f} \circ \boldsymbol{\Phi} \in H^{2}(B \times(-\varepsilon, \varepsilon))$, as $\boldsymbol{\Phi}$ is a $C^{2}$ homemorphism whose Jacobian is bounded away from 0 and $+\infty$ on $B \times(-\varepsilon, \varepsilon)$, and $t \mapsto \boldsymbol{f}(\Phi(u, t))$ is absolutely continuous for $\mathcal{L}^{2}$-a.e. $u \in B$. Thus
for $\mathcal{L}^{2}$-a.e $u \in B$ we get

$$
\boldsymbol{f}(\boldsymbol{\Phi}(u, t))=\boldsymbol{f}(\boldsymbol{\Phi}(u, 0))+t \int_{0}^{1} \frac{\mathrm{~d}(\boldsymbol{f} \circ \boldsymbol{\Phi})}{\mathrm{d} s}(u, s t) \mathrm{d} s
$$

so that by Jensen inequality, and since $\boldsymbol{f}=0$ on $\Gamma$ and $\left|\partial_{t} \Phi\right|=1$, we obtain

$$
|\boldsymbol{f}(\boldsymbol{\Phi}(u, t))|^{2} \leq|t| \int_{0 \wedge t}^{0 \vee t}|\nabla \boldsymbol{f}(\boldsymbol{\Phi}(u, s))|^{2} \mathrm{~d} s .
$$

By the latter inequality, by changing variables and by (4.19) we get for any small enough $\varepsilon$

$$
\begin{aligned}
\int_{\Phi\left(B_{\lambda} \times(-\varepsilon, \varepsilon)\right)}|\boldsymbol{f}|^{2} \mathrm{~d} x & =\int_{B_{\lambda}} \int_{-\varepsilon}^{\varepsilon}|\boldsymbol{f}(\boldsymbol{\Phi}(u, t))|^{2}|\operatorname{det} D \boldsymbol{\Phi}(u, t)| \mathrm{d} t \mathrm{~d} u \\
& \leq \int_{B_{\lambda}} \int_{-\varepsilon}^{\varepsilon}|t|\left(\int_{Q_{\wedge t}}^{0 \vee t}|\nabla \boldsymbol{f}(\boldsymbol{\Phi}(u, s))|^{2} \mathrm{~d} s\right)|\operatorname{det} D \boldsymbol{\Phi}(u, t)| \mathrm{d} t \mathrm{~d} u \\
& \leq \int_{B_{\lambda}} \int_{-\varepsilon}^{\varepsilon} 2|t| \int_{0 \wedge t}^{0 \vee t}|\nabla \boldsymbol{f}(\boldsymbol{\Phi}(u, s))|^{2}|\operatorname{det} D \boldsymbol{\Phi}(u, s)| \mathrm{d} s \mathrm{~d} t \mathrm{~d} u \\
& \leq \int_{B_{\lambda}} \int_{-\varepsilon}^{\varepsilon} 2|t| \mathrm{d} t \int_{-\varepsilon}^{\varepsilon}|\nabla \boldsymbol{f}(\boldsymbol{\Phi}(u, s))|^{2}|\operatorname{det} D \boldsymbol{\Phi}(u, s)| \mathrm{d} s \mathrm{~d} u \\
& \leq 2 \varepsilon^{2} \int_{B_{\lambda}} \int_{-\varepsilon}^{\varepsilon}|\nabla \boldsymbol{f}(\boldsymbol{\Phi}(u, \tau))|^{2}|\operatorname{det} D \boldsymbol{\Phi}(u, s)| \mathrm{d} s \mathrm{~d} u=2 \varepsilon^{2} \iint_{\boldsymbol{\Phi}\left(B_{\lambda} \times(-\varepsilon, \varepsilon)\right)}|\nabla \boldsymbol{f}|^{2} \mathrm{~d} x .
\end{aligned}
$$

Similarly, under the further null trace assumption of $\frac{\partial f}{\partial \mathbf{n}}$ on $\Gamma_{\mu, \delta}$ we deduce that

$$
\left.\frac{\mathrm{d}}{\mathrm{~d} t}(\boldsymbol{f}(\Phi(u, t)))\right|_{t=0}=\frac{\partial \boldsymbol{f}}{\partial \boldsymbol{n}}(\psi(u))
$$

vanishes as well and we obtain for $\mathcal{L}^{2}$-a.e $u \in B$, since $\partial_{t}^{2} \Phi=0$,

$$
\begin{aligned}
|\boldsymbol{f}(\Phi(u, t))| & \left.=\left|\boldsymbol{f}(\Phi(u, 0))+t \frac{\mathrm{~d}}{\mathrm{~d} t}(\boldsymbol{f}(\Phi(u, t)))\right|_{t=0}+\frac{1}{2} t^{2} \int_{0}^{1} \frac{d^{2}(\boldsymbol{f} \circ \Phi)}{d s^{2}}(u, s t) \mathrm{d} s \right\rvert\, \\
& \leq \frac{1}{2} t^{2} \int_{0}^{1}\left|\nabla^{2} \boldsymbol{f}(\Phi(u, s t))\right| \mathrm{d} s
\end{aligned}
$$

thus

$$
|\boldsymbol{f}(\boldsymbol{\Phi}(u, t))|^{2} \leq \frac{|t|^{3}}{2} \int_{0 \wedge t}^{0 \vee t}\left|\nabla^{2} \boldsymbol{f}(\boldsymbol{\Phi}(u, s))\right|^{2} \mathrm{~d} s
$$

Arguing as above we get for any small enough $\varepsilon$

$$
\begin{aligned}
\int_{\Phi\left(B_{\lambda} \times(-\varepsilon, \varepsilon)\right)}|\boldsymbol{f}|^{2} \mathrm{~d} x & =\int_{B_{\lambda}} \int_{-\varepsilon}^{\varepsilon}|\boldsymbol{f}(\boldsymbol{\Phi}(u, t))|^{2}|\operatorname{det} D \boldsymbol{\Phi}(u, t)| \mathrm{d} t \mathrm{~d} u \\
& \leq \int_{B_{\lambda}} \int_{-\varepsilon}^{\varepsilon} \frac{|t|^{3}}{2}\left(\int_{D_{\wedge t}}^{0 \vee t}\left|\nabla^{2} \boldsymbol{f}(\Phi(u, s))\right|^{2} \mathrm{~d} s\right)|\operatorname{det} D \boldsymbol{\Phi}(u, t)| \mathrm{d} t \mathrm{~d} u \\
& \leq \int_{B_{\lambda}} \int_{-\varepsilon}^{\varepsilon}|t|^{3} \mathrm{~d} t \int_{-\varepsilon}^{\varepsilon}\left|\nabla^{2} \boldsymbol{f}(\Phi(u, s))\right|^{2}|\operatorname{det} D \boldsymbol{\Phi}(u, s)| \mathrm{d} s \mathrm{~d} u=\frac{\varepsilon^{4}}{2} \int_{\Phi\left(B_{\lambda} \times(-\varepsilon, \varepsilon)\right)}\left|\nabla^{2} \boldsymbol{f}\right|^{2} \mathrm{~d} x
\end{aligned}
$$

as desired.
We are ready for the statement of the approximation result

Lemma 4.4. Assume (2.1) and (2.2). Let $0<\delta<\mu_{0}$, where $\mu_{0}$ is defined by (4.12). Let $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ such that div $\mathbf{v}=0$ a.e. in $\Omega$ and $\mathbf{v}=0$ on $\Gamma_{\delta}$. Then there exists a sequence $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}} \subset C^{1}\left(\overline{\Omega^{\prime}}, \mathbb{R}^{3}\right)$ such that $\operatorname{div} \mathbf{v}_{j}=0$ a.e. in $\Omega^{\prime}, \mathbf{v}_{j}=0$ on $\Gamma$ and $\mathbf{v}_{j} \rightarrow \mathbf{v}$ in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$.

Proof. Let $0<\mu<\mu_{0}$. Recalling (4.13), (4.27) and (4.14), notice that if $\lambda \in(0, \delta / 2), \varepsilon \in(0, \mu / 2)$ then there hold $S_{2 \varepsilon, 2 \lambda} \subset S_{\mu, \delta}$ and $S_{2 \varepsilon, 2 \lambda} \cap \partial \Omega \subset \Gamma_{\delta}$, which will be crucial for the proof: the projection on $\partial \Omega$ of any point in $S_{2 \varepsilon, 2 \lambda}$ lies on $\Gamma_{\delta}$. On the other hand it clear that $S_{2 \varepsilon, 2 \lambda} \subset \Omega^{\prime \prime}$, where $\Omega^{\prime \prime}$ is defined by (4.15), thus $\mathbf{n}$ is well defined on $S_{2 \varepsilon, 2 \lambda}$.

Let $\zeta \in C^{2}(\mathbb{R})$ be defined by

$$
\zeta(\xi):= \begin{cases}\left(3 \xi^{2}-2 \xi^{3}\right)^{2} & \text { if } \xi \in[0,1] \\ 0 & \text { if } \xi<0 \\ 1 & \text { if } \xi>1\end{cases}
$$

We introduce some notation: $S_{\varepsilon, \lambda}^{*}:=S_{2 \varepsilon, \lambda} \backslash S_{\varepsilon, \lambda}, S_{\varepsilon, \lambda}^{* *}:=S_{2 \varepsilon, 2 \lambda} \backslash S_{2 \varepsilon, \lambda}, \tilde{S}_{\varepsilon, \lambda}:=S_{\varepsilon, \lambda}^{*} \cup S_{\varepsilon, \lambda}^{* *}$. Let $\lambda \in(0, \delta / 2)$, let $\varepsilon \in(0, \lambda \wedge(\mu / 2))$ and let

$$
\eta_{\varepsilon, \lambda}(x):=\left\{\begin{array}{l}
0 \text { if } x \in S_{\varepsilon, \lambda} \\
\zeta\left(\frac{d(x, \partial \Omega)-\varepsilon}{\varepsilon}\right) \text { if } x \in S_{\varepsilon, \lambda}^{*} \\
\zeta\left(\frac{d(x, \partial \Omega)-\varepsilon}{\varepsilon}\right)+\zeta\left(\frac{d(\sigma(x), \partial \Gamma)-\lambda}{\lambda}\right)\left(1-\zeta\left(\frac{d(x, \partial \Omega)-\varepsilon}{\varepsilon}\right)\right) \text { if } x \in S_{\varepsilon, \lambda}^{* *} \\
1 \text { otherwise in } \mathbb{R}^{3}
\end{array}\right.
$$

We notice that in the particular case $\partial \Gamma=\emptyset$, we have $S_{\varepsilon, \lambda}=S_{\varepsilon}, S_{\varepsilon, \lambda}^{*}=S_{2 \varepsilon} \backslash S_{\varepsilon}, S_{\varepsilon, \lambda}^{* *}=\emptyset$, and in fact $\eta_{\varepsilon, \lambda}$ does not depend on $\lambda$. Taking advantage of the $C^{3}$ regularity of $d(\cdot, \partial \Omega)$ in $\Omega^{\prime \prime} \cap\left\{x \in \mathbb{R}^{3}: d(x, \partial \Omega)>\right.$ $\varepsilon / 2\}$, of the $C^{2}$ regularity of $\sigma$ in $\Omega^{\prime \prime}$ and of the $C^{2}$ regularity of $x \mapsto d(\sigma(x), \partial \Gamma)$ in $S_{2 \varepsilon, 2 \lambda} \backslash S_{2 \varepsilon, \lambda}$ (see Remark 4.2), it can be easily checked that $\eta_{\varepsilon, \lambda} \in C^{1,1}\left(\mathbb{R}^{3}\right)$. Moreover, we have $|\nabla d(x, \partial \Omega)| \leq 1$ and $\left|\nabla^{2} d(x, \partial \Omega)\right| \leq C_{*} / \varepsilon$ on $S_{\varepsilon, \lambda}^{*}$ (and similarly, $|\nabla(d(\sigma(x), \partial \Gamma))| \leq C_{*}$, and $\left|\nabla^{2}(d(\sigma(x), \partial \Gamma))\right| \leq C_{*} / \lambda$ on $\left.S_{\varepsilon, \lambda}^{* *}\right)$ for some $C_{*}>$ that is independent of $\varepsilon$ and $\lambda$. Taking advantage of these distance estimates, a computation shows that there is a constant $C>0$ (not depending on $\varepsilon$ and $\lambda$ ) such that

$$
\begin{align*}
& \nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}=0 \text { in } S_{\varepsilon, \lambda}, \quad 2\left|\nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}\right| \leq \frac{C}{\varepsilon} \text { in } \tilde{S}_{\varepsilon, \lambda} \\
& \mathbf{n} \wedge\left(\nabla \eta_{\varepsilon, \lambda} \wedge \mathbf{n}\right)=0 \text { in } S_{\varepsilon, \lambda}^{*}, \quad 2\left|\mathbf{n} \wedge\left(\nabla \eta_{\varepsilon, \lambda} \wedge \mathbf{n}\right)\right| \leq \frac{C}{\lambda}<\frac{C}{\varepsilon} \text { in } S_{\varepsilon, \lambda}^{* *},  \tag{4.20}\\
& 2\left|\nabla\left(\nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}\right)\right| \leq \frac{C}{\varepsilon^{2}} \text { in } S_{\varepsilon, \lambda}^{*}, \quad 2\left|\nabla\left(\mathbf{n} \wedge\left(\nabla \eta_{\varepsilon, \lambda} \wedge \mathbf{n}\right)\right)\right| \leq \frac{C}{\lambda \varepsilon} \text { in } S_{\varepsilon, \lambda}^{* *} .
\end{align*}
$$

Thanks to Lemma 3.7, we find $\widetilde{\mathbf{w}} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\widetilde{\mathbf{w}}=0$ on $\Gamma_{\delta}$ and $\operatorname{curl} \widetilde{\mathbf{w}}=\mathbf{v}$ a.e. in $\Omega$ (in particular, $\operatorname{curl} \widetilde{\mathbf{w}}=0$ on $\left.\Gamma_{\delta}\right)$. Let us consider a $H^{2}\left(\mathbb{R}^{3}, \mathbb{R}^{3}\right)$ extension of $\widetilde{\mathbf{w}}$, still denoted by $\widetilde{\mathbf{w}}$, and set

$$
\mathbf{w}_{\varepsilon, \lambda}:=\eta_{\varepsilon, \lambda} \widetilde{\mathbf{w}} .
$$

Therefore, $\mathbf{w}_{\varepsilon, \lambda} \in H^{2}\left(\Omega^{\prime}, \mathbb{R}^{3}\right), \mathbf{w}_{\varepsilon, \lambda}=0$ on $\Gamma_{\delta}$ and

$$
\begin{equation*}
\operatorname{curl} \mathbf{w}_{\varepsilon, \lambda}=\eta_{\varepsilon, \lambda} \operatorname{curl} \widetilde{\mathbf{w}}-\widetilde{\mathbf{w}} \wedge \nabla \eta_{\varepsilon, \lambda} \tag{4.21}
\end{equation*}
$$

so that curl $\mathbf{w}_{\varepsilon, \lambda}=0$ on $\Gamma_{\delta}$ as well. Moreover, for $i=1,2,3$,

$$
\begin{equation*}
\partial_{i} \operatorname{curl} \mathbf{w}_{\varepsilon, \lambda}=\partial_{i} \eta_{\varepsilon, \lambda} \operatorname{curl} \widetilde{\mathbf{w}}+\eta_{\varepsilon, \lambda} \partial_{i} \operatorname{curl} \widetilde{\mathbf{w}}-\partial_{i} \widetilde{\mathbf{w}} \wedge \nabla \eta_{\varepsilon, \lambda}-\widetilde{\mathbf{w}} \wedge \partial_{i} \nabla \eta_{\varepsilon, \lambda} \tag{4.22}
\end{equation*}
$$

and it is readily seen that, as $\varepsilon \rightarrow 0^{+}$,

$$
\begin{equation*}
\eta_{\varepsilon, \lambda} \operatorname{curl} \widetilde{\mathbf{w}} \rightarrow \operatorname{curl} \widetilde{\mathbf{w}} \quad \text { and } \quad \eta_{\varepsilon, \lambda} \partial_{i} \operatorname{curl} \widetilde{\mathbf{w}} \rightarrow \partial_{i} \operatorname{curl} \widetilde{\mathbf{w}} \tag{4.23}
\end{equation*}
$$

in $L^{2}\left(\Omega, \mathbb{R}^{3}\right)$. We claim that

$$
\begin{equation*}
\widetilde{\mathbf{w}} \wedge \nabla \eta_{\varepsilon, \lambda} \rightarrow 0, \quad \partial_{i} \eta_{\varepsilon, \lambda} \operatorname{curl} \widetilde{\mathbf{w}} \rightarrow 0, \quad \partial_{i} \widetilde{\mathbf{w}} \wedge \nabla \eta_{\varepsilon, \lambda} \rightarrow 0, \quad \widetilde{\mathbf{w}} \wedge \partial_{i} \nabla \eta_{\varepsilon, \lambda} \rightarrow 0 \tag{4.24}
\end{equation*}
$$

in $L^{2}\left(\Omega, \mathbb{R}^{3}\right)$ as $\varepsilon \rightarrow 0^{+}$, which will then imply, along with (4.21)-(4.22)-(4.23), that

$$
\begin{equation*}
\operatorname{curl} \mathbf{w}_{\varepsilon, \lambda} \rightarrow \operatorname{curl} \widetilde{\mathbf{w}}=\mathbf{v} \text { in } H^{1}\left(\Omega, \mathbb{R}^{3}\right) \text { as } \varepsilon \rightarrow 0^{+} . \tag{4.25}
\end{equation*}
$$

In order to prove the claim, we separately treat the four terms in (4.24), by taking into account (4.20) and Lemma 3.8, which entails $\widetilde{\mathbf{w}}=\widetilde{\mathbf{w}} \wedge \mathbf{n}=\frac{\partial}{\partial \mathbf{n}}(\widetilde{\mathbf{w}} \wedge \mathbf{n})=0$ on $\Gamma_{\delta}\left(\right.$ thus, $\nabla(\widetilde{\mathbf{w}} \wedge \mathbf{n})=0$ on $\Gamma_{\delta}$ as well). We get, thanks to the usual decomposition $\mathbf{a}=(\mathbf{a} \cdot \mathbf{n}) \mathbf{n}+\mathbf{n} \wedge(\mathbf{a} \wedge \mathbf{n})$, thanks to Lemma 4.3 and to (4.16), and by recalling that $\nabla \eta_{\varepsilon, \lambda}=0$ outside $\tilde{S}_{\varepsilon, \lambda}$ we get,

$$
\begin{aligned}
\int_{\Omega}\left|\widetilde{\mathbf{w}} \wedge \nabla \eta_{\varepsilon, \lambda}\right|^{2} \mathrm{~d} x & \leq \int_{\tilde{S}_{\varepsilon, \lambda}}\left|\widetilde{\mathbf{w}} \wedge \nabla \eta_{\varepsilon, \lambda}\right|^{2} \mathrm{~d} x \\
& \leq 4 \int_{\tilde{S}_{\varepsilon, \lambda}}\left|\nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}\right|^{2}|\widetilde{\mathbf{w}} \wedge \mathbf{n}|^{2} \mathrm{~d} x+4 \int_{S_{\varepsilon, *}^{* *}}|\widetilde{\mathbf{w}}|^{2}\left|\mathbf{n} \wedge\left(\nabla \eta_{\varepsilon, \lambda} \wedge \mathbf{n}\right)\right|^{2} \\
& \leq \frac{C^{2}}{\lambda^{2}} \int_{S_{\varepsilon, \lambda}^{* *}}|\widetilde{\mathbf{w}}|^{2} \mathrm{~d} x+\frac{C^{2}}{\varepsilon^{2}} \int_{S_{2 \varepsilon, 2 \lambda}}|\widetilde{\mathbf{w}}|^{2} \mathrm{~d} x \\
& \leq \frac{C^{2}}{\lambda^{2}} \int_{S_{\varepsilon, \lambda}^{* *}}|\widetilde{\mathbf{w}}|^{2} \mathrm{~d} x+2 C^{2} \int_{S_{2 \varepsilon, 2 \lambda}}|\nabla \widetilde{\mathbf{w}}|^{2} \mathrm{~d} x
\end{aligned}
$$

so that $\widetilde{\mathbf{w}} \wedge \nabla \eta_{\varepsilon, \lambda} \rightarrow 0$ in $L^{2}\left(\Omega, \mathbb{R}^{3}\right)$ as $\varepsilon \rightarrow 0^{+}$. On the other hand, still making use of Lemma 4.3,

$$
\begin{aligned}
\int_{\Omega}\left|\partial_{i} \widetilde{\mathbf{w}} \wedge \nabla \eta_{\varepsilon}\right|^{2} \mathrm{~d} x & \leq \int_{\tilde{S}_{\varepsilon, \lambda}}\left|\partial_{i} \widetilde{\mathbf{w}} \wedge \nabla \eta_{\varepsilon, \lambda}\right|^{2} \mathrm{~d} x \\
& \leq 4 \int_{\tilde{S}_{\varepsilon, \lambda}}\left(\left|\partial_{i} \widetilde{\mathbf{w}}\right|^{2}\left|\mathbf{n} \wedge \nabla \eta_{\varepsilon, \lambda} \wedge \mathbf{n}\right|^{2}+\left|\nabla \eta_{\varepsilon} \cdot \mathbf{n}\right|^{2}\left|\partial_{i} \widetilde{\mathbf{w}} \wedge \mathbf{n}\right|^{2}\right) \mathrm{d} x \\
& \leq \frac{C^{2}}{\lambda^{2}} \int_{\tilde{S}_{\varepsilon, \lambda}}\left|\partial_{i} \widetilde{\mathbf{w}}\right|^{2} \mathrm{~d} x+\frac{C^{2}}{\varepsilon^{2}} \int_{S_{2 \varepsilon, 2 \lambda}}\left|\partial_{i}(\widetilde{\mathbf{w}} \wedge \mathbf{n})\right|^{2} \mathrm{~d} x+\frac{C^{2}}{\varepsilon^{2}} \int_{S_{2 \varepsilon, 2 \lambda}}\left|\widetilde{\mathbf{w}} \wedge \partial_{i} \mathbf{n}\right|^{2} \mathrm{~d} x \\
& \leq \frac{C^{2}}{\lambda^{2}} \int_{\tilde{S}_{\varepsilon, \lambda}}\left|\partial_{i} \widetilde{\mathbf{w}}\right|^{2} \mathrm{~d} x+\frac{C^{2}}{\varepsilon^{2}} \int_{S_{2 \varepsilon, 2 \lambda}}|\nabla(\widetilde{\mathbf{w}} \wedge \mathbf{n})|^{2} \mathrm{~d} x+\frac{K^{2} C^{2}}{\varepsilon^{2}} \int_{S_{2 \varepsilon, 2 \lambda}}|\widetilde{\mathbf{w}}|^{2} \mathrm{~d} x \\
& \leq \frac{C^{2}}{\lambda^{2}} \int_{\tilde{S}_{\varepsilon, \lambda}}\left|\partial_{i} \widetilde{\mathbf{w}}\right|^{2} \mathrm{~d} x+8 C^{2} \int_{S_{2 \varepsilon, 2 \lambda}}\left|\nabla^{2}(\widetilde{\mathbf{w}} \wedge \mathbf{n})\right|^{2} \mathrm{~d} x+8 K^{2} C^{2} \int_{S_{2 \varepsilon, 2 \lambda}}|\nabla \widetilde{\mathbf{w}}|^{2} \mathrm{~d} x \\
& \leq \frac{C^{2}}{\lambda^{2}} \int_{\tilde{S}_{\varepsilon, \lambda}}\left|\partial_{i} \widetilde{\mathbf{w}}\right|^{2} \mathrm{~d} x+24 K^{2} C^{2} \int_{S_{2 \varepsilon, 2 \lambda}}\left(|\widetilde{\mathbf{w}}|^{2}+|\nabla \widetilde{\mathbf{w}}|^{2}+\left|\nabla^{2} \widetilde{\mathbf{w}}\right|^{2}\right) \mathrm{d} x
\end{aligned}
$$

hence $\partial_{i} \widetilde{\mathbf{w}} \wedge \nabla \eta_{\varepsilon, \lambda} \rightarrow 0$ in $L^{2}\left(\Omega, \mathbb{R}^{3}\right)$ as $\varepsilon \rightarrow 0^{+}$. Similarly, taking advantage of the fact that curl $\widetilde{\mathbf{w}}=0$ on $\Gamma_{\mu, \delta}$ and of Lemma 4.3, we get as $\varepsilon \rightarrow 0^{+}$

$$
\int_{\Omega}\left|\partial_{i} \eta_{\varepsilon, \lambda} \operatorname{curl} \widetilde{\mathbf{w}}\right|^{2} \mathrm{~d} x \leq \frac{C^{2}}{\varepsilon^{2}} \int_{\tilde{S}_{\varepsilon, \lambda}}|\operatorname{curl} \widetilde{\mathbf{w}}|^{2} \mathrm{~d} x \leq 8 C^{2} \int_{S_{2 \varepsilon, 2 \lambda}}|\nabla \operatorname{curl} \widetilde{\mathbf{w}}|^{2} \mathrm{~d} x \rightarrow 0,
$$

thus $\partial_{i} \eta_{\varepsilon, \lambda}$ curl $\widetilde{\mathbf{w}} \rightarrow 0$ in $L^{2}\left(\Omega, \mathbb{R}^{3}\right)$ as $\varepsilon \rightarrow 0^{+}$. Moreover, by applying the usual decomposition in normal and tangential part to $\nabla \eta_{\varepsilon, \lambda}$ we get

$$
\begin{aligned}
\widetilde{\mathbf{w}} \wedge \partial_{i}\left(\nabla \eta_{\varepsilon, \lambda}\right)= & \partial_{i}\left(\nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}\right) \widetilde{\mathbf{w}} \wedge \mathbf{n}+\left(\nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}\right) \widetilde{\mathbf{w}} \wedge \partial_{i} \mathbf{n} \\
& +\widetilde{\mathbf{w}} \wedge\left(\partial_{i} \mathbf{n} \wedge\left(\nabla \eta_{\varepsilon, \lambda} \wedge \mathbf{n}\right)\right)+\widetilde{\mathbf{w}} \wedge\left(\mathbf{n} \wedge \partial_{i}\left(\nabla \eta_{\varepsilon, \lambda} \wedge \mathbf{n}\right)\right) \\
= & \partial_{i}\left(\nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}\right) \widetilde{\mathbf{w}} \wedge \mathbf{n}+\left(\nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}\right) \widetilde{\mathbf{w}} \wedge \partial_{i} \mathbf{n}+\widetilde{\mathbf{w}} \wedge\left(\partial_{i}\left(\mathbf{n} \wedge\left(\nabla \eta_{\varepsilon, \lambda} \wedge \mathbf{n}\right)\right)\right),
\end{aligned}
$$

and by taking again advantage of (4.20) we find

$$
\begin{aligned}
& \int_{\Omega}\left|\partial_{i}\left(\nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}\right) \widetilde{\mathbf{w}} \wedge \mathbf{n}\right|^{2} \mathrm{~d} x \leq \int_{\tilde{S}_{\varepsilon, \lambda}}\left|\nabla\left(\nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}\right)\right|^{2}|\widetilde{\mathbf{w}} \wedge \mathbf{n}|^{2} \mathrm{~d} x \leq \frac{C^{2}}{4 \varepsilon^{4}} \int_{S_{2 \varepsilon, 2 \lambda}}|\widetilde{\mathbf{w}} \wedge \mathbf{n}|^{2} \mathrm{~d} x, \\
& \int_{\Omega}\left|\left(\nabla \eta_{\varepsilon, \lambda}\right) \widetilde{\mathbf{w}} \wedge \partial_{i} \mathbf{n}\right|^{2} \mathrm{~d} x \leq K^{2} \int_{\tilde{S}_{\varepsilon, \lambda}}\left|\nabla \eta_{\varepsilon, \lambda} \cdot \mathbf{n}\right|^{2}|\widetilde{\mathbf{w}}|^{2} \mathrm{~d} x \leq \frac{K^{2} C^{2}}{4 \varepsilon^{2}} \int_{S_{2 \varepsilon, 2 \lambda}}|\widetilde{\mathbf{w}}|^{2} \mathrm{~d} x, \\
& \int_{\Omega}\left|\widetilde{\mathbf{w}} \wedge\left(\partial_{i}\left(\mathbf{n} \wedge\left(\nabla \eta_{\varepsilon, \lambda} \wedge \mathbf{n}\right)\right)\right)\right|^{2} \mathrm{~d} x \leq \int_{\tilde{S}_{\varepsilon, \lambda}}\left|\nabla\left(\mathbf{n} \wedge\left(\nabla \eta_{\varepsilon, \lambda} \wedge \mathbf{n}\right)\right)\right|^{2}|\widetilde{\mathbf{w}}|^{2} \mathrm{~d} x \leq \frac{C^{2}}{4 \varepsilon^{2} \lambda^{2}} \int_{S_{2 \varepsilon, 2 \lambda}}|\widetilde{\mathbf{w}}|^{2} d x
\end{aligned}
$$

and we see that all these integrals are reduced to the ones of the previous estimates, so that indeed by using Lemma 4.3 they all vanish as $\varepsilon \rightarrow 0^{+}$, showing that $\widetilde{\mathbf{w}} \wedge \partial_{i}\left(\nabla \eta_{\varepsilon, \lambda}\right) \rightarrow 0$ in $L^{2}\left(\Omega, \mathbb{R}^{3}\right)$ as $\varepsilon \rightarrow 0^{+}$. This proves the claim, so that (4.25) holds true.

We stress that $\widetilde{\mathbf{w}}_{\varepsilon, \lambda}$ and $\operatorname{curl} \widetilde{\mathbf{w}}_{\varepsilon, \lambda}$ vanish in $S_{\varepsilon, \lambda}$, hence in an open neighbor of $\Gamma$ in $\mathbb{R}^{3}$. Let now $\left(\varepsilon_{j}\right)_{j \in \mathbb{N}} \subset(0,+\infty)$ be such that $\varepsilon_{j} \rightarrow 0^{+}$as $j \rightarrow+\infty$, let $\left(\rho_{j}\right)_{j \in \mathbb{N}}$ be a sequence of smooth mollifiers such that the support of $\rho_{j}$ is so small that $\widetilde{\mathbf{w}}_{\varepsilon_{j}, \lambda} * \rho_{j}$ still vanishes on a neighbor of $\Gamma$. Then, we define $\widetilde{\mathbf{w}}_{j, \lambda}:=\widetilde{\mathbf{w}}_{\varepsilon_{j}, \lambda} * \rho_{j}$, and $\mathbf{v}_{j}:=\operatorname{curl} \widetilde{\mathbf{w}}_{j, \lambda}$. It is readily seen that $\mathbf{v}_{j} \in C^{1}\left(\overline{\Omega^{\prime}}, \mathbb{R}^{3}\right)$, that div $\mathbf{v}_{j}=0$ in $\Omega^{\prime}$ and that $\mathbf{v}_{j}=0$ on $\Gamma$. By recalling (4.25) we get also $\mathbf{v}_{j} \rightarrow \mathbf{v}$ in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ thus concluding the proof.
Remark 4.5. If $\Gamma=\partial \Omega$ (i.e., if $\partial \Gamma=\emptyset$ ), then it is enough to assume $C^{2,1}$ regularity of $\partial \Omega$ in Lemma 4.4. Indeed, we still get boundedness of $\nabla^{2} \mathbf{n}$ in $\Omega^{\prime \prime}$, which allows the latter proof to carry over.

The final step of this section is another suitable approximation property of divergence-free $H^{1}$ vector fields, that is required to treat the case $\partial \Gamma \neq \emptyset$. It is stated in Lemma 4.9. It also requires some preliminary lemmas and some further notation.

Suppose that $\partial \Gamma \neq \emptyset$. For every $0<\mu<\mu_{0}$, where $\mu_{0}$ is defined by (4.12), let

$$
\partial_{l} S_{\mu}:=\{\sigma+\operatorname{tn}(\sigma): \sigma \in \partial \Gamma,|t| \leq \mu\} .
$$

Since $\mathbf{n} \in C^{2}(\partial \Omega)$, we see that $\partial_{l} S_{\mu}$ is a compact $C^{2}$ manifold with boundary. It is the lateral boundary of $S_{\mu}$, and we denote by $\nu_{l}$ the corresponding outward unit normal vector to $\partial_{l} S_{\mu}$. Moreover, for every $0<\delta<\mu_{0}$ and $0<\mu<\mu_{0}$, let $\boldsymbol{\phi}_{\mu, \delta}: \partial_{l} S_{\mu} \times[-\delta, \delta] \rightarrow \mathbb{R}^{3}$ be defined by

$$
\begin{equation*}
\phi_{\mu, \delta}(s, \tau):=s+\tau \boldsymbol{\nu}_{l}(s) . \tag{4.26}
\end{equation*}
$$

It is clear that each point of $\phi_{\mu, \delta}\left(\partial_{l} S_{\mu} \times[-\delta, \delta]\right)$ is within the reach of both $\partial \Omega$ and $\partial \Gamma . \phi_{\mu, \delta}$ inherits the $C^{1}$ regularity of $\boldsymbol{\nu}_{l}$. We recall the following simple property.

Lemma 4.6. Assume (2.1) and (2.2), with $\partial \Gamma \neq \emptyset$. Let $\mu<\mu_{0}$. There exists $\delta_{0} \in\left(0, \mu_{0}\right)$ such that the map $\phi_{\mu, \delta}: \partial_{l} S_{\mu} \times\left[-\delta_{0}, \delta_{0}\right] \rightarrow \mathbb{R}^{3}$, defined by (4.26), is one-to-one.
Proof. Assume by contradiction that there is no $\delta_{0}>0$ with the required property. Then there exist sequences $\left(s_{j}\right)_{j \in \mathbb{N}} \subset \partial_{l} S_{\mu},\left(s_{j}^{\prime}\right)_{j \in \mathbb{N}} \subset \partial_{l} S_{\mu}$ and $\left(t_{j}\right)_{j \in \mathbb{N}} \subset(0,1 / n),\left(t_{j}^{\prime}\right)_{j \in \mathbb{N}} \subset(0,1 / n)$ such that $\left(s_{j}, t_{j}\right) \neq$ $\left(s_{j}^{\prime}, t_{j}^{\prime}\right)$ and $\boldsymbol{\phi}_{\mu, \delta}\left(s_{j}, t_{j}\right)=\boldsymbol{\phi}_{\mu, \delta}\left(s_{j}^{\prime}, t_{j}^{\prime}\right)$ for any $j \in \mathbb{N}$ (hence, $s_{j} \neq s_{j}^{\prime}$ for any $j \in \mathbb{N}$ ). Since $\partial_{l} S_{\mu}$ is compact, up to subsequences we have $s_{j} \rightarrow s \in \partial_{l} S_{\mu}, s_{j}^{\prime} \rightarrow s^{\prime} \in \partial_{l} S_{\mu}, t_{j} \rightarrow 0$ and $t_{j}^{\prime} \rightarrow 0$. Therefore, the continuity of $\phi_{\mu, \delta}$ implies $\boldsymbol{\phi}_{\mu, \delta}(s, 0)=\phi_{\mu, \delta}\left(s^{\prime}, 0\right)$, i.e., $s^{\prime}=s$. This means that $(s, 0)$ has no open neighbor in $\mathbb{R}^{3}$ where
$\boldsymbol{\phi}_{\mu, \delta}$ is invertible. Let $B$ denote the unit open ball in $\mathbb{R}^{2}$ : for a given $C^{2}$ local chart $u \ni B \mapsto \boldsymbol{\psi}(u) \in \mathbb{R}^{3}$ describing a neighbor of $s$ on the surface $\partial_{l} S_{\mu}$, such that $\boldsymbol{\psi}(0)=s$, we have $\left|\operatorname{det} D \phi_{\mu, \delta}(\boldsymbol{\psi}(0), 0)\right|=$ $\left|\partial_{1} \boldsymbol{\psi}(0) \wedge \partial_{2} \boldsymbol{\psi}(0)\right| \neq 0$ (by the regularity of the surface), where $D$ denotes the gradient in the variable $(u, t)$. Therefore, the $C^{1}(B \times(-\delta, \delta))$ mapping $\phi_{\mu, \delta} \circ(\boldsymbol{\psi}, \boldsymbol{i})$ has non vanishing Jacobian at the point $(s, 0)=\phi_{\mu, \delta}(\psi(0), 0)$, hence it is invertible in a neighbor of such point, a contradiction.

Still for $\partial \Gamma \neq \emptyset$, for every $0<\mu<\mu_{0}$ and for every $0<\delta<\delta_{0}$, where $\delta_{0}$ is the threshold provided by Lemma 4.6, we define

$$
\begin{equation*}
T_{\mu, \delta}^{ \pm}:=\left\{s \pm \tau \boldsymbol{\nu}_{l}(s): s \in \partial_{l} S_{\mu}, 0 \leq \tau \leq \delta\right\}, \quad T_{\mu, \delta}:=T_{\mu, \delta}^{+} \cup T_{\mu, \delta}^{-} \tag{4.27}
\end{equation*}
$$

We have the following
Lemma 4.7. Assume (2.1) and (2.2), with $\partial \Gamma \neq \emptyset$. Let $\mathbf{w} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ be such that $\mathbf{w}=\frac{\partial}{\partial \mathbf{n}}(\mathbf{w} \wedge \mathbf{n})=0$ on $\Gamma$. Then there exist a vanishing sequence $\left(\lambda_{j}\right)_{j \in \mathbb{N}} \subset\left(0, \mu_{0}\right)$ and a sequence $\left(\mathbf{w}_{j}\right)_{j \in \mathbb{N}} \subset H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\mathbf{w}_{j}=\frac{\partial}{\partial \mathbf{n}}\left(\mathbf{w}_{j} \wedge \mathbf{n}\right)=0$ on $\Gamma_{\lambda_{j}}$ and $\mathbf{w}_{j} \rightarrow \mathbf{w}$ in $H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ as $j \rightarrow+\infty$.

Proof. Let $\delta_{0}$ be the threshold from Lemma 4.6. Let $\mu$ and $\delta$ be such that

$$
0<3 \delta<\mu<\frac{1}{2} \delta_{0} \quad \text { and } \quad 4 \delta<\frac{1}{\operatorname{Lip}\left(\boldsymbol{\nu}_{l}\right)},
$$

where $\operatorname{Lip}\left(\boldsymbol{\nu}_{l}\right):=\sup \left\{\frac{\left|\boldsymbol{\nu}_{l}(z)-\boldsymbol{\nu}_{l}\left(z^{\prime}\right)\right|}{\left|z-z^{\prime}\right|}: z \in \partial_{l} S_{\mu}, z^{\prime} \in \partial_{l} S_{\mu}, z \neq z^{\prime}\right\}<+\infty$, since $\boldsymbol{\nu}_{l} \in C^{1}\left(\partial_{l} S_{\mu}\right)$.
We define for every $y \in \Gamma_{2 \delta}$ and for every $0<\lambda<(\delta / 2) \wedge 1$

$$
\psi_{\lambda}(y):= \begin{cases}y-2 \gamma_{\lambda}\left(d^{2}(y, \partial \Gamma)\right)(\mathbf{t} \wedge \mathbf{n})(s(y)) & \text { if } y \in\left\{x \in \Gamma_{2 \delta}: d(x, \partial \Gamma) \leq 2 \delta\right\} \\ y & \text { otherwise in } \Gamma_{2 \delta}\end{cases}
$$

where $s(y) \in \partial \Gamma$ is the unique nearest point of $\partial \Gamma$ to $y \in\left\{x \in \Gamma_{2 \delta}: d(x, \partial \Gamma) \leq 2 \delta\right\}$ (recalling (4.12) so that $2 \delta<\delta_{0}<\mu_{0}$ implies that $y$ is within the reach of $\partial \Gamma$ ), and $\mathbf{t}$ is the unit tangent vector to $\partial \Gamma$ (positively orienting $\partial \Gamma$ with respect to $\mathbf{n}$, so that $(\mathbf{t} \wedge \mathbf{n})(s(y))$ coincides with the outward unit vector $\boldsymbol{\nu}_{l}(s(y))$ to $\left.\partial_{l} S_{\mu}\right)$. Moreover, here $\gamma_{\lambda} \in C^{2}(\mathbb{R})$ is a decreasing cutoff function such that $\gamma_{\lambda}(\xi)=0$ if $\xi \geq \delta$ and $\gamma_{\lambda}(\xi)=\lambda$ if $\xi \leq \delta / 2$. We stress that $d^{2}(\cdot, \partial \Gamma)$ is a $C^{3}$ function on the set $\left\{x \in \mathbb{R}^{3}: d(x, \partial \Gamma)<\mu_{0}\right\}$, so that since $s(y)=y-\frac{1}{2} \nabla\left(d^{2}(y, \partial \Gamma)\right), s(\cdot)$ is $C^{2}$ on such set (see Remark 4.2).

The following property holds: there exists $\lambda_{0} \in(0,(\delta / 2) \wedge 1)$ such that, for any $\lambda<\lambda_{0}$,

$$
\begin{equation*}
\sigma\left(\psi_{\lambda}(y)\right) \in \Gamma \quad \text { for every } y \in \Gamma_{\lambda} \tag{4.28}
\end{equation*}
$$

where $\sigma(\cdot)$ denotes as usual the unique projection on $\partial \Omega$ (since $2 \gamma_{\lambda} \leq 2 \lambda<\mu_{0}$, then $\psi_{\lambda}(y)$ is within the reach of $\partial \Omega$ ). This crucial property is proved in a separate statement, i.e., in Lemma 4.8 below.

Let us consider an $H^{2}\left(\mathbb{R}^{3}, \mathbb{R}^{3}\right)$ extension of $\mathbf{w}$, still denoted by $\mathbf{w}$.
For any $x \in S_{2 \mu, 2 \delta}$ (thus within the reach of $\partial \Omega$, since $2 \mu<\delta_{0}<\mu_{0}$ ), we introduce the signed distance function

$$
b(x):= \begin{cases}d(x, \partial \Omega) & \text { if } x \notin \Omega \\ -d(x, \partial \Omega) & \text { if } x \in \Omega\end{cases}
$$

and we notice that $b(x)$ is the unique real number such that $x=\sigma(x)+t(x) \mathbf{n}(\sigma(x))$. We notice that $b \in C^{3}\left(\Omega^{\prime \prime}\right)$ : indeed, we have $\mathbf{n}=\nabla b$. See also [17, Theorem 3.1] for regularity results about the signed distance function. We let $g_{\lambda}: S_{2 \mu, 2 \delta} \rightarrow \partial \Omega$ be defined by $g_{\lambda}(x):=\sigma\left(\phi_{\lambda}(\sigma(x))\right)$ and we further define

$$
\begin{aligned}
\mathbf{h}_{\lambda}(x):= & {\left[\mathbf{w}\left(g_{\lambda}(x)+b(x) \mathbf{n}\left(g_{\lambda}(x)\right)\right) \cdot \mathbf{n}\left(g_{\lambda}(x)\right)\right] \mathbf{n}(\sigma(x)) } \\
& +\mathbf{n}(\sigma(x)) \wedge\left[\mathbf{w}\left(g_{\lambda}(x)+b(x) \mathbf{n}\left(g_{\lambda}(x)\right)\right) \wedge \mathbf{n}\left(g_{\lambda}(x)\right)\right] .
\end{aligned}
$$

We note that $\mathbf{n}$ is extended to a $C^{2}\left(\Omega^{\prime \prime}, \mathbb{R}^{3}\right)$ vector field in the usual way, see (4.15), so that $\mathbf{n}(\sigma(x))=$ $\mathbf{n}(x)$. We let $\eta \in C^{2}\left(\mathbb{R}^{3}\right)$ be a cutoff function, such that $0 \leq \eta \leq 1, \eta \equiv 0$ in $S_{\mu, \delta}$, and the support of $1-\eta$ is contained in $\stackrel{\circ}{S}_{2 \mu, 2 \delta}$. Then we define, for every $x \in \mathbb{R}^{3}$,

$$
\mathbf{w}_{\lambda}^{*}(x)=(1-\eta(x)) \mathbf{h}_{\lambda}(x),
$$

so that indeed $\mathbf{w}_{\lambda}^{*}$ is supported in $\stackrel{\circ}{S}_{2 \mu, 2 \delta}$, and

$$
\mathbf{w}_{\lambda}(x):=\mathbf{w}_{\lambda}^{*}(x)+\eta(x) \mathbf{w}(x) .
$$

We claim that $\mathbf{w}_{\lambda} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ for any small enough $\lambda$ and that $\mathbf{w}_{\lambda} \rightarrow \mathbf{w}$ in $H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ as $\lambda \rightarrow 0$. Indeed, recalling the definition of $\gamma_{\lambda}$ and the fact that $d^{2}(\cdot, \partial \Gamma)$ and $s(\cdot)$ are $C^{2}$ functions on $\left\{x \in \mathbb{R}^{3}\right.$ : $\left.d(x, \partial \Gamma)<\mu_{0}\right\}$, it is readily seen that $\psi_{\lambda} \in C^{2}\left(\Gamma_{2 \delta}\right)$ and that $\psi_{\lambda}(y) \rightarrow y$ in $C^{2}\left(\Gamma_{2 \delta}\right)$ as $\lambda \rightarrow 0$. We also observe that $\sigma: \Omega^{\prime \prime} \rightarrow \partial \Omega$ is a $C^{2}$ function by assumption (2.1). Therefore, we obtain $\psi_{\lambda} \circ \sigma \rightarrow \sigma$ in $C^{2}\left(S_{2 \mu, 2 \delta}\right)$ as $\lambda \rightarrow 0$, and similarly $g_{\lambda}=\sigma \circ \psi_{\lambda} \circ \sigma \rightarrow \sigma$ in $C^{2}\left(S_{2 \mu, 2 \delta}\right)$ and $\mathbf{n} \circ g_{\lambda} \rightarrow \mathbf{n} \circ \sigma=\mathbf{n}$ in $C^{2}\left(S_{2 \mu, 2 \delta}\right)$. Since $x=\sigma(x)+b(x) \mathbf{n}(\sigma(x))$, for small enough $\lambda$ we see that the mapping

$$
S_{2 \mu, 2 \delta} \ni x \mapsto q_{\lambda}(x):=g_{\lambda}(x)+b(x) \mathbf{n}\left(g_{\lambda}(x)\right) \in \Omega^{\prime \prime}
$$

is a $C^{2}$ homeomorphism whose Jacobian is bounded away from 0 , and moreover by the previous remarks it converges to the identity as $\lambda \rightarrow 0$ in $C^{2}\left(S_{2 \mu, 2 \delta}\right)$. Since $\mathbf{w} \in H^{2}\left(\mathbb{R}^{3}, \mathbb{R}^{3}\right)$, we obtain by the properties of Sobolev functions (as in Lemma 4.3), that $\mathbf{w} \circ q_{\lambda} \in H^{2}\left(\stackrel{\circ}{S}_{2 \mu, 2 \delta}, \mathbb{R}^{3}\right)$, and moreover it is easy to check that $\mathbf{w} \circ q_{\lambda} \rightarrow \mathbf{w}$ in $H^{2}\left(\stackrel{S}{S}_{2 \mu, 2 \delta}, \mathbb{R}^{3}\right)$ as $\lambda \rightarrow 0$. Taking the product with the smooth cutoff function $1-\eta$ (supported on $\left.S_{2 \mu, 2 \delta}\right)$, we deduce that $(1-\eta)\left(\mathbf{w} \circ q_{\lambda}\right) \rightarrow(1-\eta) \mathbf{w}$ in $H^{2}\left(\mathbb{R}^{3}, \mathbb{R}^{3}\right)$. Moreover, since we also have $\mathbf{n} \circ g_{\lambda} \rightarrow \mathbf{n} \circ \sigma=\mathbf{n}$ in $C^{2}\left(S_{2 \mu, 2 \delta}\right)$, we obtain

$$
\mathbf{w}_{\lambda}^{*}=(1-\eta) \mathbf{h}_{\lambda} \rightarrow(1-\eta)\{[\mathbf{w} \cdot \mathbf{n}] \mathbf{n}+\mathbf{n} \wedge[\mathbf{w} \wedge \mathbf{n}]\}=(1-\eta) \mathbf{w}
$$

in $H^{2}\left(\mathbb{R}^{3}, \mathbb{R}^{3}\right)$ as $\lambda \rightarrow 0$. Thus $\mathbf{w}_{\lambda} \rightarrow \mathbf{w}$ in $H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ as $\lambda \rightarrow 0$. The claim is proved.
We shall now prove that $\mathbf{w}_{\lambda}=\frac{\partial}{\partial \mathbf{n}}\left(\mathbf{w}_{\lambda} \wedge \mathbf{n}\right)=0$ on $\Gamma_{\lambda}$ for any small enough $\lambda$. Indeed, let $\lambda<\lambda_{0}$ be small enough, such that $\mathbf{w}_{\lambda} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$. If $x \in \Gamma_{\lambda}$, then by the property (4.28) we get $g_{\lambda}(x) \in \Gamma$, and since $b(x)=0$ and $\mathbf{w}=0$ on $\Gamma$ we directly obtain $\mathbf{h}_{\lambda}(x)=\mathbf{w}_{\lambda}^{*}(x)=0$ on $\Gamma_{\lambda}$. But $\eta(x)=0$ as well (because $\eta=0$ on $S_{\mu, \delta}$ and $\lambda<\delta / 2$ so that $\left.\Gamma_{\lambda} \subset \Gamma_{\delta} \subset S_{\mu, \delta}\right)$, hence $\mathbf{w}_{\lambda}(x)=0$. On the other hand for every $x \in \Gamma_{\lambda}$, we have $g_{\lambda}(x)=\sigma\left(\psi_{\lambda}(x)\right)$, and we have $\sigma(x+r \mathbf{n}(x))=x, b(x+r \mathbf{n}(x))=r$ when $|r|$ is small enough, therefore

$$
\begin{aligned}
\mathbf{h}_{\lambda}(x+r \mathbf{n}(x)):=[ & \mathbf{w}\left(\sigma\left(\psi_{\lambda}(x)\right)+r \mathbf{n}\left(\sigma\left(\psi_{\lambda}(x)\right)\right) \cdot \mathbf{n}\left(\sigma\left(\psi_{\lambda}(x)\right)\right)\right] \mathbf{n}(x) \\
& +\mathbf{n}(x) \wedge\left[\mathbf{w}\left(\sigma\left(\psi_{\lambda}(x)\right)+r \mathbf{n}\left(\sigma\left(\psi_{\lambda}(x)\right)\right)\right) \wedge \mathbf{n}\left(\sigma\left(\psi_{\lambda}(x)\right)\right)\right]
\end{aligned}
$$

and

$$
\begin{aligned}
& \mathbf{h}_{\lambda}(x+r \mathbf{n}(x)) \wedge \mathbf{n}(x+r \mathbf{n}(x))=\mathbf{h}_{\lambda}(x+r \mathbf{n}(x)) \wedge \mathbf{n}(x) \\
& \quad=\left\{\mathbf{n}(x) \wedge\left[\left(\mathbf{w}\left(\sigma\left(\phi_{\lambda}(x)\right)+r \mathbf{n}\left(\sigma\left(\phi_{\lambda}(x)\right)\right)\right) \wedge \mathbf{n}\left(\sigma\left(\phi_{\lambda}(x)\right)\right)\right]\right\} \wedge \mathbf{n}(x) .\right.
\end{aligned}
$$

As a consequence, by taking into account that $\sigma\left(\psi_{\lambda}(x)\right) \in \Gamma$ and that $\frac{\partial}{\partial \mathbf{n}}(\mathbf{w} \wedge \mathbf{n})=0$ on $\Gamma$ we get

$$
\begin{aligned}
\frac{\partial}{\partial \mathbf{n}}\left(\mathbf{h}_{\lambda} \wedge \mathbf{n}\right)(x) & =\lim _{r \rightarrow 0^{+}} \frac{\mathbf{h}_{\lambda}(x+r \mathbf{n}(x)) \wedge \mathbf{n}(x+r \mathbf{n}(x))-\mathbf{h}_{\lambda}(x) \wedge \mathbf{n}(x)}{r} \\
& =\mathbf{n}(x) \wedge\left(\frac{\partial(\mathbf{w} \wedge \mathbf{n})}{\partial \mathbf{n}}\left(\sigma\left(\psi_{\lambda}(x)\right)\right) \wedge \mathbf{n}(x)\right)=0,
\end{aligned}
$$

and since we have already shown that $\mathbf{h}_{\lambda}$ vanishes on $\Gamma_{\lambda}$, we conclude that $\frac{\partial}{\partial \mathbf{n}}\left(\mathbf{w}_{\lambda}^{*} \wedge \mathbf{n}\right)=0$ on $\Gamma_{\lambda}$. Again, $\eta$ is vanishing on $S_{\mu, \delta}$, hence in an open neighbor of $\Gamma_{\lambda}$. We deduce that $\frac{\partial}{\partial \mathbf{n}}\left(\mathbf{w}_{\lambda} \wedge \mathbf{n}\right)=0$ on $\Gamma_{\lambda}$.

Eventually, by taking a vanishing sequence of small enough positive numbers $\left(\lambda_{j}\right)_{j \in \mathbb{N}}$, we conclude that $\mathbf{w}_{j}:=\mathbf{w}_{\lambda_{j}}$ satisfies all the desired properties.

Lemma 4.8. Assume (2.1) and (2.2), with $\partial \Gamma \neq \emptyset$. Let $\mu, \delta, \lambda, \phi_{\lambda}$ as in the proof of Lemma 4.7. There exists $\lambda_{0} \in(0,(\delta / 2) \wedge 1)$ such that, for any $\lambda<\lambda_{0}$, (4.28) holds.

Proof. If $d(y, \partial \Gamma) \geq \delta$ there is nothing to prove, since in this case either $y=\phi_{\lambda}(y) \in \Gamma$ or $y \notin \Gamma_{\lambda}$ (because $\lambda<\delta / 2)$. Therefore, we assume from now on that $d(y, \partial \Gamma)<\delta$ and we prove the result in four steps.
Step 0 We start by showing the elementary properties $T_{\mu, 3 \delta}^{-} \subset S_{2 \mu}$ and $T_{\mu, 3 \delta}^{+} \cap \dot{S}_{2 \mu}=\emptyset$. Indeed, concerning the first property, we may prove that for any $s \in \partial_{l} S_{\mu}$, a point of the form $s-\alpha \boldsymbol{\nu}_{l}(s)$, with $0 \leq \alpha \leq 3 \delta$, belongs to $S_{2 \mu}$. This is obvious if $\alpha$ is small, since $\boldsymbol{\nu}_{l}$ is normal to $\partial S_{2 \mu}$. Moreover, as $\alpha$ increases without reaching the threshold $\delta_{0}$, the closest point of $\partial_{l} S_{2 \mu}$ is always $s$, by Lemma 4.6 which gives the unique projection property on $\partial_{l} S_{2 \mu}$. This shows that no other point of $\partial_{l} S_{2 \mu}$ can be a reached. And since $\alpha \leq 3 \delta$, we are also far from $\partial S_{2 \mu} \backslash \partial_{l} S_{2 \mu} \subset\left\{x \in \mathbb{R}^{3}: d(x, \partial \Omega)=2 \mu\right\}$, because $d\left(s-\alpha \boldsymbol{\nu}_{l}(s), \partial \Omega\right) \leq d(s, \partial \Omega)+3 \delta=$ $\mu+3 \delta<2 \mu$. The second property is proved in the same way.
Step 1 We check that $y \in T_{\mu, \delta}$. Indeed, we have $d\left(y, \partial_{l} S_{\mu}\right) \leq d(y, \partial \Gamma)=d(y, s(y))<\delta$. We take a point $s_{*} \in \partial_{l} S_{\mu}$ such that $\left|y-s_{*}\right|=d\left(y, \partial_{l} S_{\mu}\right)$, therefore $d\left(s_{*}, s(y)\right) \leq d(y, s(y))+d\left(y, s_{*}(y)\right)<2 \delta<\mu$ so that $s_{*}$ is not on the boundary of $\partial_{l} S_{\mu}$. Since $s_{*}$ is a minimizer of the distance function from the $C^{2}$ manifold $\partial_{l} S_{\mu}$, the corresponding first order minimality conditions immediately imply that $y-s_{*}$ is orthogonal to the tangent plane to $\partial_{l} S_{\mu}$ at $s_{*}$, so that

$$
y=s_{*}+\tau_{*}(y) \boldsymbol{\nu}_{l}\left(s_{*}\right),
$$

where $\left|\tau_{*}(y)\right|=d\left(y, \partial_{l} S_{\mu}\right)<\delta$. Thus, $y \in T_{\mu, \delta}$. We notice that by Lemma 4.6, $s_{*}$ coincides in fact with the unique projection $s_{*}(y)$ of $y$ on $\partial_{l} S_{\mu}$.
Step 2 We prove the result in the case $y \in \Gamma_{\lambda} \backslash \Gamma$. In this case we have $d(y, \partial \Gamma)=d(y, s(y)) \leq \lambda$ and as a consequence $d\left(s(y), s_{*}(y)\right) \leq 2 \lambda$. We also have $\gamma_{\lambda}\left(d^{2}(y, \partial \Gamma)\right)=\lambda$ because $\lambda<(\delta / 2) \wedge 1$, and moreover using Step 1 and $y \notin \Gamma$ we have $y \in T_{\mu, \delta}^{+}$, so that $\tau_{*}(y)>0$. Actually, $y \in T_{2 \lambda, \delta}^{+}$as well, since $d\left(s(y), s_{*}(y)\right) \leq 2 \lambda$.

By taking into account that $(\mathbf{t} \wedge \mathbf{n})(s(y))=\boldsymbol{\nu}_{l}(s(y))$ we get

$$
\begin{align*}
\psi_{\lambda}(y) & =y-2 \gamma_{\lambda}\left(d^{2}(y, \partial \Gamma)\right)(\mathbf{t} \wedge \mathbf{n})(s(y))=y-2 \lambda \boldsymbol{\nu}_{l}(s(y)) \\
& =s_{*}(y)+\tau_{*}(y) \boldsymbol{\nu}_{l}\left(s_{*}(y)\right)-2 \lambda \boldsymbol{\nu}_{l}(s(y))  \tag{4.29}\\
& =s_{*}(y)+\left(\tau_{*}(y)-2 \lambda\right) \boldsymbol{\nu}_{l}\left(s_{*}(y)\right)+2 \lambda\left(\boldsymbol{\nu}_{l}\left(s_{*}(y)\right)-\boldsymbol{\nu}_{l}(s(y))\right) .
\end{align*}
$$

We notice that the point $s_{*}(y)+\left(\tau_{*}(y)-2 \lambda\right) \boldsymbol{\nu}_{l}\left(s_{*}(y)\right)$ belongs to $T_{2 \lambda, \delta}^{-}$, because $\tau_{*}(y)=d\left(y, \partial_{l} S_{\mu}\right) \leq$ $d(y, s(y)) \leq \lambda$ and therefore $-\delta<-2 \lambda \leq \tau_{*}(y)-2 \lambda \leq-\lambda$, and we have in particular

$$
\begin{equation*}
d\left(s_{*}(y)+\left(\tau_{*}(y)-2 \lambda\right) \boldsymbol{\nu}_{l}\left(s_{*}(y)\right), \partial_{l} S_{\mu}\right) \geq \lambda \tag{4.30}
\end{equation*}
$$

But

$$
\begin{equation*}
\left|2 \lambda\left(\boldsymbol{\nu}_{l}\left(s_{*}(y)\right)-\boldsymbol{\nu}_{l}(s(y))\right)\right| \leq 2 \lambda \operatorname{Lip}\left(\boldsymbol{\nu}_{l}\right)\left|s_{*}(y)-s(y)\right| \leq 4 \lambda^{2} \operatorname{Lip}\left(\boldsymbol{\nu}_{l}\right) . \tag{4.31}
\end{equation*}
$$

By (4.29), (4.30) and (4.31), there exists a small enough $\lambda_{0}$ depending on $\operatorname{Lip}\left(\boldsymbol{\nu}_{l}\right)$ such that for any $\lambda<\lambda_{0}$ we have $\psi_{\lambda}(y) \in T_{\mu, \delta}^{-}$. Since $T_{\mu, \delta}^{-} \subset S_{2 \mu}$ by Step 0 , the result is proved.
Step 3 We prove the result in case $y \in \Gamma$. Since $y \in T_{\mu, \delta}$ by Step 1, we have in this case $y \in T_{\mu, \delta}^{-}$, therefore we have $\tau_{*}(y)=-d\left(y, \partial_{l} S_{\mu}\right)$. By $d\left(s_{*}(y), s(y)\right)<2 \delta$, we have in particular, $y \in T_{2 \delta, \delta}^{-}$. With the same computation of Step 2, we obtain an expression which is analogous to (4.29), that is,

$$
\begin{equation*}
\psi_{\lambda}(y)=s_{*}(y)-\left(d\left(y, \partial_{l} S_{\mu}\right)+2 \gamma_{\lambda}\left(d^{2}(y, \partial \Gamma)\right)\right) \boldsymbol{\nu}_{l}\left(s_{*}(y)\right)+2 \gamma_{\lambda}\left(d^{2}(y, \partial \Gamma)\right)\left(\boldsymbol{\nu}_{l}\left(s_{*}(y)\right)-\boldsymbol{\nu}_{l}(s(y))\right) . \tag{4.32}
\end{equation*}
$$

In particular, since $0 \leq d\left(y, \partial_{l} S_{\mu}\right)+2 \gamma_{\lambda}\left(d^{2}(y, \partial \Gamma)\right) \leq \delta+2 \lambda<2 \delta$, we have

$$
\begin{equation*}
s_{*}(y)-\left(d\left(y, \partial_{l} S_{\mu}\right)+2 \gamma_{\lambda}\left(d^{2}(y, \partial \Gamma)\right)\right) \boldsymbol{\nu}_{l}\left(s_{*}(y)\right) \in T_{2 \delta, \delta+2 \lambda}^{-}, \tag{4.33}
\end{equation*}
$$

with

$$
\begin{equation*}
d\left(s_{*}(y)-\left(d\left(y, \partial_{l} S_{\mu}\right)+2 \gamma_{\lambda}\left(d^{2}(y, \partial \Gamma)\right)\right) \boldsymbol{\nu}_{l}\left(s_{*}(y)\right), \partial_{l} S_{\mu}\right) \geq 2 \gamma_{\lambda}\left(d^{2}(y, \partial \Gamma)\right) \tag{4.34}
\end{equation*}
$$

By the assumptions on $\delta$, we have

$$
\begin{equation*}
\left|\boldsymbol{\nu}_{l}\left(s_{*}(y)\right)-\boldsymbol{\nu}_{l}(s(y))\right| \leq \operatorname{Lip}\left(\boldsymbol{\nu}_{l}\right)\left|s_{*}(y)-s(y)\right| \leq 2 \delta \operatorname{Lip}\left(\boldsymbol{\nu}_{l}\right) \leq \frac{1}{2} \tag{4.35}
\end{equation*}
$$

By (4.32), (4.33), (4.34) and (4.35), we conclude that $\psi_{\lambda}(y) \in T_{2 \delta+\lambda, \delta+3 \lambda}^{-}$. Therefore, $y \in T_{\mu, 3 \delta}^{-}$, since $2 \delta+\lambda<\mu$ and $\delta+3 \lambda<3 \delta$. By Step 0 , the result is proved.

Thanks to the results of Sect. 3 and to Lemma 4.7, we deduce the final approximation result for curl vector fields.

Lemma 4.9. Assume (2.1) and (2.2), with $\partial \Gamma \neq \emptyset$. Let $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\operatorname{div} \mathbf{v}=0$ a.e. in $\Omega$ and $\mathbf{v}=0$ on $\Gamma$. Then there exist a vanishing sequence $\left(\lambda_{j}\right)_{j \in \mathbb{N}} \subset\left(0, \mu_{0}\right)$ and a sequence $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}} \subset H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ such that div $\mathbf{v}_{j}=0$ a.e. in $\Omega, \mathbf{v}_{j}=0$ on $\Gamma_{\lambda_{j}}$ and such that $\mathbf{v}_{j} \rightarrow \mathbf{v}$ in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ as $j \rightarrow+\infty$.

Proof. By Lemma 3.7 there exists $\widetilde{\mathbf{w}} \in H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\widetilde{\mathbf{w}}=0$ on $\Gamma$ and $\operatorname{curl} \widetilde{\mathbf{w}}=\mathbf{v}$ a.e. in $\Omega$, so that Lemma 3.8 implies $\frac{\partial}{\partial \mathbf{n}}(\widetilde{\mathbf{w}} \wedge \mathbf{n})=0$ on $\Gamma$. Hence, by Lemma 4.7 there exist a vanishing sequence $\left(\lambda_{j}\right)_{j \in \mathbb{N}} \subset(0,+\infty)$ and a sequence $\left(\mathbf{w}_{j}\right)_{j \in \mathbb{N}} \subset H^{2}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\mathbf{w}_{j}=\frac{\partial}{\partial \mathbf{n}}\left(\mathbf{w}_{j} \wedge \mathbf{n}\right)=0$ on $\Gamma_{\lambda_{j}}$ and $\mathbf{w}_{j} \rightarrow \widetilde{\mathbf{w}}$ in $H^{2}\left(\Omega, \mathbb{R}^{3}\right)$. By Lemma 3.8 we get $\operatorname{curl} \mathbf{w}_{j}=0$ on $\Gamma_{\lambda_{j}}$, hence by setting $\mathbf{v}_{j}:=\operatorname{curl} \mathbf{w}_{j}$ the result follows.

## 5. Proof of the main result

Let us start by recalling the following version of the rigidity inequality by Friesecke, James and Müller [20].

Lemma 5.1. (Geometric Rigidity Inequality [2,21]). Let $g_{p}$ the function defined in (2.3). There exists a constant $C_{p}=C_{p}(\Omega)>0$ such that for every $\mathbf{y} \in W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ there exists a constant $\mathbf{R} \in S O(3)$ such that we have

$$
\begin{equation*}
\int_{\Omega} g_{p}(|\nabla \mathbf{y}-\mathbf{R}|) d x \leq C_{p} \int_{\Omega} g_{p}(d(\nabla \mathbf{y}, S O(3))) d x \tag{5.1}
\end{equation*}
$$

Based on the above result, we deduce compactness of minimizing sequences, which follows in fact from the results in [2].

Lemma 5.2. Assume (2.1), (2.2), (W) , (WW 2 , ( $\mathcal{W} 3),(\mathcal{W} 4)$. Let $\left(h_{j}\right)_{j \in \mathbb{N}}$ be a sequence of positive real numbers and let $\left(\mathbf{v}_{j}\right) \subset W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ be a sequence such that $\mathbf{v}_{j}=0$ on $\Gamma$ for any $j \in \mathbb{N}$. For every $j \in \mathbb{N}$, let $\mathbf{y}_{j}=\boldsymbol{i}+h_{j} \mathbf{v}_{j}$ and let $\mathbf{R}_{j} \in S O(3)$ be a constant rotation satisfying (5.1). Then there exists a constant $C>0$ (only depending on $p, \Omega$ and $\Gamma$ ) such that for any $j \in \mathbb{N}$ there hold

$$
\begin{equation*}
\left|\mathbf{I}-\mathbf{R}_{j}\right|^{2} \leq C \int_{\Omega} \mathcal{W}^{I}\left(x, \mathbf{I}+h_{j} \mathbf{v}_{j}\right) d x \tag{5.2}
\end{equation*}
$$

and

$$
\begin{equation*}
\int_{\Omega}\left|\nabla \mathbf{v}_{j}\right|^{p} d x \leq C\left(1+\int_{\Omega} \mathcal{W}^{I}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) d x\right) \tag{5.3}
\end{equation*}
$$

If we assume in addition that $h_{j} \rightarrow 0$ as $j \rightarrow+\infty$ and that

$$
\begin{equation*}
\lim _{j \rightarrow+\infty}\left(\mathcal{G}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right)-\inf _{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)} \mathcal{G}_{h_{j}}^{I}\right)=0 \tag{5.4}
\end{equation*}
$$

then $\sup _{j \in \mathbb{N}}\left\|\mathbf{v}_{j}\right\|_{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)}<+\infty$.

Proof. We have $\mathcal{W}^{I} \geq \mathcal{W}$, and (5.2) holds true with $\mathcal{W}$ in place of $\mathcal{W}^{I}$ as proven in [2, Lemma 3.3] (by taking advantage of assumption $(\mathcal{W} 4)$ on $\mathcal{W}$ ). Therefore, (5.2) holds.

Using the form of $g_{p}$ it is clear that there exists a constant $c$ (only depending on $p$ ) such that

$$
\int_{\Omega} g_{p}\left(h_{j}\left|\nabla \mathbf{v}_{j}\right|\right) \mathrm{d} x \leq c \int_{\Omega}\left(g_{p}\left(\left|\mathbf{I}+h_{j} \nabla \mathbf{v}_{j}-\mathbf{R}_{j}\right|\right)+\left|\mathbf{I}-\mathbf{R}_{j}\right|^{2}\right) \mathrm{d} x .
$$

Hence, by invoking the rigidity estimate (5.1), there is another constant $K$ (only depending on $p$ and $\Omega$ ) such that

$$
\int_{\Omega} g_{p}\left(h_{j}\left|\nabla \mathbf{v}_{j}\right|\right) \mathrm{d} x \leq K\left(\int_{\Omega} g_{p}\left(d\left(\mathbf{I}+h_{j} \nabla \mathbf{v}_{j}, S O(3)\right)\right) \mathrm{d} x+\left|\mathbf{I}-\mathbf{R}_{j}\right|^{2}\right)
$$

and since $x^{p} \leq 1+2 g_{p}(x)$ holds for $x \geq 0$, by making use of $(\mathcal{W} 4)$ and (5.2) it follows that there is a further constant $C$ (only depending on $\Omega, \Gamma, \mathrm{p}$ ) such that

$$
\int_{\Omega}\left|\nabla \mathbf{v}_{j}\right|^{p} \mathrm{~d} x \leq \int_{\Omega}\left(1+2 g_{p}\left(h_{j}\left|\nabla \mathbf{v}_{j}\right|\right)\right) \mathrm{d} x \leq C\left(1+\int_{\Omega} \mathcal{W}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) \mathrm{d} x\right) .
$$

Since $\mathcal{W} \leq \mathcal{W}^{I}$, (5.3) follows.
Let us prove the last statement. Assuming (5.4) and assuming wlog that $\left\|\mathbf{v}_{j}\right\|_{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)} \geq 1$ for any $j \in \mathbb{N}$, we get for any large enough $j$

$$
\frac{1}{h_{j}^{2}} \int_{\Omega} \mathcal{W}^{I}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) \mathrm{d} x-\mathcal{L}\left(\mathbf{v}_{j}\right)=\mathcal{G}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right) \leq \mathcal{G}_{h_{j}}^{I}(\mathbf{0})+1=1
$$

thus (5.3) implies

$$
\int_{\Omega}\left|\nabla \mathbf{v}_{j}\right|^{p} \mathrm{~d} x \leq C\left(1+\int_{\Omega} \mathcal{W}^{I}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) \mathrm{d} x\right) \leq C+C\left(h_{j}^{2}+C_{\mathcal{L}} h_{j}^{2}\left\|\mathbf{v}_{j}\right\|_{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)}^{p}\right)
$$

Since $h_{j}$ goes to zero, the result follows by Friedrichs inequality.
We next prove $\Gamma$-convergence. The limsup inequality is based on the approximation results from Sect. 4. The liminf inequality builds on previous arguments from [2,15,35].
 vanishing sequence of positive numbers. Then the sequence of functionals $\left(\mathcal{G}_{h_{j}}^{I}\right)_{j \in \mathbb{N}}$ is $\Gamma$-converging to functional $\mathcal{G}^{I}$ with respect to the weak topology of $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$.
Proof. Since the weak topology of $W^{1, p}$ is metrizable then we can characterize the $\Gamma$-limit in terms of weakly converging sequences. In particular, by setting (see [14,16])

$$
\begin{aligned}
& \mathcal{G}_{-}^{I}(\mathbf{v}):=\inf \left\{\liminf _{j \rightarrow \infty} \mathcal{G}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right): \mathbf{v}_{j} \rightharpoonup \mathbf{v} \text { weakly in } W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)\right\}, \\
& \mathcal{G}_{+}^{I}(\mathbf{v}):=\inf \left\{\limsup _{j \rightarrow \infty} \mathcal{G}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right): \mathbf{v}_{j} \rightharpoonup \mathbf{v} \text { weakly in } W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)\right\},
\end{aligned}
$$

since $\mathcal{G}_{+}^{I}(\mathbf{v}) \geq \mathcal{G}_{-}^{I}(\mathbf{v})$, it is enough to prove that $\mathcal{G}_{+}^{I}(\mathbf{v}) \leq \mathcal{G}^{I}(\mathbf{v}) \leq \mathcal{G}_{-}^{I}(\mathbf{v})$ for every $\mathbf{v} \in W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$. We split the proof in two steps.

Step 1 (liminf) We show that $\mathcal{G}^{I}(\mathbf{v}) \leq \mathcal{G}_{-}^{I}(\mathbf{v})$ for every $\mathbf{v} \in W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$.
Let $\mathbf{v} \in W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$, assume without restriction that $\mathcal{G}_{-}^{I}(\mathbf{v})<+\infty$, and let $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}} \subset W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ be a sequence such that $\mathbf{v}_{j} \rightharpoonup \mathbf{v}$ weakly in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ as $j \rightarrow+\infty$ and such that $\sup _{j \in \mathbb{N}} \mathcal{G}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right)<+\infty$.

Then $\mathbf{v}_{j}=0$ on $\Gamma$ for any $j \in \mathbb{N}$, hence $\mathbf{v}=0$ on $\Gamma$ as well, and by setting $\mathbf{B}_{j}:=2 \mathbb{E}\left(\mathbf{v}_{j}\right)+h_{j} \nabla \mathbf{v}_{j}^{T} \nabla \mathbf{v}_{j}$ we get

$$
\begin{aligned}
1 & =\operatorname{det}\left(\mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right)=\operatorname{det}\left(\mathbf{I}+h_{j} \nabla \mathbf{v}_{j}^{T}\right)\left(\mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right)=\operatorname{det}\left(\mathbf{I}+2 h_{j} \mathbb{E}\left(\mathbf{v}_{j}\right)+h_{j}^{2} \nabla \mathbf{v}_{j}^{T} \nabla \mathbf{v}_{j}\right) \\
& =1+h_{j} \operatorname{Tr} \mathbf{B}_{j}-\frac{1}{2} h_{j}^{2}\left(\operatorname{Tr}\left(\mathbf{B}_{j}^{2}\right)-\left(\operatorname{Tr} \mathbf{B}_{j}\right)^{2}\right)+h_{j}^{3} \operatorname{det} \mathbf{B}_{j}
\end{aligned}
$$

a.e. in $\Omega$, that is,

$$
\begin{equation*}
\operatorname{Tr} \mathbf{B}_{j}=2 \operatorname{div} \mathbf{v}_{j}+h_{j}\left|\nabla \mathbf{v}_{j}\right|^{2}=\frac{1}{2} h_{j}\left(\operatorname{Tr}\left(\mathbf{B}_{j}^{2}\right)-\left(\operatorname{Tr} \mathbf{B}_{j}\right)^{2}\right)-h_{j}^{2} \operatorname{det} \mathbf{B}_{j} . \tag{5.5}
\end{equation*}
$$

We next prove, with an argument from [2], that $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ and that

$$
\begin{gather*}
1_{D_{j}} \nabla \mathbf{v}_{j} \rightharpoonup \nabla \mathbf{v} \text { weakly in } L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right),  \tag{5.6}\\
1_{\Omega \backslash D_{j}} \nabla \mathbf{v}_{j} \rightarrow 0 \text { in } L^{\alpha}\left(\Omega, \mathbb{R}^{3 \times 3}\right), \quad \forall \alpha \in[1, p), \tag{5.7}
\end{gather*}
$$

where we have set $D_{j}:=\left\{x \in \Omega: \sqrt{h_{j}}\left|\nabla \mathbf{v}_{j}(x)\right| \leq 1\right\}$. Indeed, since we are assuming that $\sup _{j \in \mathbb{N}} \mathcal{G}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right)$ $<+\infty$, we have

$$
\begin{equation*}
\sup _{j \in \mathbb{N}} \frac{1}{h_{j}^{2}} \int_{\Omega} \mathcal{W}^{I}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) \mathrm{d} x<+\infty \tag{5.8}
\end{equation*}
$$

thanks to the definition of $\mathcal{G}_{h_{j}}^{I}$ and to the boundedness of the sequence $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}}$ in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$. Let $\mathbf{R}_{j} \in S O(3)$ be a constant matrix satisfying (5.1) with respect to $\mathbf{y}_{j}=\boldsymbol{i}+h_{j} \mathbf{v}_{j}$. If $Q_{j}:=\{x \in \Omega$ : $\left.\left|\mathbf{I}+h_{j} \nabla \mathbf{v}_{j}(x)-\mathbf{R}_{j}\right| \leq 3 \sqrt{3}\right\}$, we have $D_{j} \subset Q_{j}$ for any $j$ large enough, and by definition of $g_{p}$ it is clear that there exists a constant $K$ only depending on $p$ such that $g_{p}(x) \geq K x^{2}$ for any $x \in[0,3 \sqrt{3}]$, so that

$$
\begin{aligned}
\int_{D_{j}}\left|\nabla \mathbf{v}_{j}\right|^{2} \mathrm{~d} x & \leq \frac{K}{h_{j}^{2}} \int_{Q_{j}}\left(g_{p}\left(\left|\mathbf{I}+h_{j} \nabla \mathbf{v}_{j}-\mathbf{R}_{j}\right|\right)+\left|\mathbf{I}-\mathbf{R}_{j}\right|^{2}\right) \mathrm{d} x \\
& \leq \frac{K C}{h_{j}^{2}} \int_{\Omega} \mathcal{W}^{I}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) \mathrm{d} x+K|\Omega| \frac{\left|\mathbf{I}-\mathbf{R}_{j}\right|^{2}}{h_{j}^{2}},
\end{aligned}
$$

where we have used ( $\mathcal{W} 4$ ) and (5.1). By taking advantage of (5.2) and of (5.8), we conclude that the sequence $\left(1_{D_{j}} \nabla \mathbf{v}_{j}\right)_{j \in \mathbb{N}}$ is bounded in $L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right)$, so that up to (not relabeled) subsequences, $1_{D_{j}} \nabla \mathbf{v}_{j} \rightharpoonup$ $\mathbf{H}$ weakly in $L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right)$. On the other hand, if $\alpha \in[1, p)$, by Hölder inequality and the definition of $D_{j}$ we have

$$
\left\|1_{\Omega \backslash D_{j}} \nabla \mathbf{v}_{j}\right\|_{L^{\alpha}\left(\Omega, \mathbb{R}^{3 \times 3}\right)} \leq\left.\left\|\nabla \mathbf{v}_{j}\right\|_{L^{p}\left(\Omega, \mathbb{R}^{3 \times 3}\right)}|\Omega| D_{j}\right|^{\frac{p-\alpha}{p \alpha}} \leq\left\|\nabla \mathbf{v}_{j}\right\|_{L^{p}\left(\Omega, \mathbb{R}^{3 \times 3}\right)}\left(\sqrt{h_{j}} \int_{\Omega}\left|\nabla \mathbf{v}_{j}\right| \mathrm{d} x\right)^{\frac{p-\alpha}{p \alpha}}
$$

and (5.7) follows from the fact that the above right hand side is vanishing as $j \rightarrow+\infty$, since the sequence $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}}$ is bounded in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$. The latter property also implies the weak convergence (up to not relabeled subsequences) of $\nabla \mathbf{v}_{j}$ to $\nabla \mathbf{v}$ in $L^{\alpha}\left(\Omega, \mathbb{R}^{3 \times 3}\right)$ : since (5.7) holds and since $1_{D_{j}} \nabla \mathbf{v}_{j}=$ $\left(\nabla \mathbf{v}_{j}-1_{\Omega \backslash D_{j}} \nabla \mathbf{v}_{j}\right)$, we obtain both $\nabla \mathbf{v}=\mathbf{H} \in L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right)$ and (5.6), and Friedrichs inequality yields $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$.

Thanks to the properties (5.6) and (5.7) we get

$$
\sqrt{h_{j}} \nabla \mathbf{v}_{j}=\sqrt{h_{j}}\left(1_{D_{j}} \nabla \mathbf{v}_{j}+1_{\Omega \backslash D_{j}} \nabla \mathbf{v}_{j}\right) \rightarrow 0 \quad \operatorname{in} L^{\alpha}\left(\Omega, \mathbb{R}^{3 \times 3}\right)
$$

as $j \rightarrow+\infty$ for any $\alpha \in[1, p)$, hence (up to not relabeled subsequences) $\sqrt{h_{j}} \nabla \mathbf{v}_{j} \rightarrow 0$ a.e. in $\Omega$. By taking into account that for some constant $c>0$ there hold

$$
\begin{aligned}
& \left|\operatorname{Tr} \mathbf{B}_{j}^{2}\right| \leq c\left(\left|\nabla \mathbf{v}_{j}\right|^{2}+h_{j}^{2}\left|\nabla \mathbf{v}_{j}\right|^{4}+h_{j}\left|\nabla \mathbf{v}_{j}\right|^{3}\right), \\
& \left|\operatorname{Tr} \mathbf{B}_{j}\right|^{2} \leq c\left(\left|\nabla \mathbf{v}_{j}\right|^{2}+h_{j}^{2}\left|\nabla \mathbf{v}_{j}\right|^{4}\right), \\
& \left|\operatorname{det} \mathbf{B}_{j}\right| \leq c\left|\mathbf{B}_{j}\right|^{3} \leq C\left(\left|\nabla \mathbf{v}_{j}\right|^{3}+h_{j}^{3}\left|\nabla \mathbf{v}_{j}\right|^{6}\right),
\end{aligned}
$$

we get

$$
h_{j}\left|\nabla \mathbf{v}_{j}\right|^{2}+\frac{1}{2} h_{j}\left(\operatorname{Tr}\left(\mathbf{B}_{j}^{2}\right)+\left(\operatorname{Tr} \mathbf{B}_{j}\right)^{2}\right)+h_{j}^{2} \operatorname{det} \mathbf{B}_{j} \rightarrow 0
$$

a.e. in $\Omega$ as $j \rightarrow+\infty$. Hence, by (5.5), $\operatorname{div} \mathbf{v}_{j} \rightarrow 0$ a.e. in $\Omega$ and by recalling that $\operatorname{div} \mathbf{v}_{j} \rightharpoonup \operatorname{div} \mathbf{v}$ weakly in $L^{p}(\Omega)$ we have div $\mathbf{v}=0$ a.e. in $\Omega$. Since we have previously shown that $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ and that $\mathbf{v}=0$ on $\Gamma$, we deduce that $\mathcal{G}^{I}(\mathbf{v})$ is finite.

By assumption ( $\mathcal{W} 3), D^{2} \mathcal{W}(x, \cdot) \in C^{2}(\mathcal{U})$ for a.e. $x \in \Omega$ and there is an increasing function $\omega$ : $[0,+\infty) \rightarrow \mathbb{R}$ such that $\lim _{y \rightarrow 0} \omega(y)=0$ and $\left|D^{2} \mathcal{W}(x, \mathbf{I}+\mathbf{F})-D^{2} \mathcal{W}(x, \mathbf{I})\right| \leq \omega(|\mathbf{F}|)$ for any $\mathbf{F} \in \mathcal{U}$ and for a.e. $x \in \Omega$. We notice that for any large enough $j$, we have $\mathbf{I}+h_{j} \nabla \mathbf{v}_{j} \in \mathcal{U}$ for any $x \in D_{j}$. Therefore,

$$
\begin{align*}
& \limsup _{j \rightarrow+\infty} \int_{D_{j}}\left|\frac{1}{h_{j}^{2}} \mathcal{W}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right)-\frac{1}{2} \nabla \mathbf{v}_{j}^{T} D^{2} \mathcal{W}(x, \mathbf{I}) \nabla \mathbf{v}_{j}\right| \mathrm{d} x  \tag{5.9}\\
& \quad \leq \limsup _{j \rightarrow+\infty} \int_{D_{j}} \omega\left(h_{j}\left|\nabla \mathbf{v}_{j}\right|\right)\left|\nabla \mathbf{v}_{j}\right|^{2} \mathrm{~d} x \leq \limsup _{j \rightarrow+\infty} \omega\left(\sqrt{h_{j}}\right) \int_{\Omega} 1_{D_{j}}\left|\nabla \mathbf{v}_{j}\right|^{2} \mathrm{~d} x=0
\end{align*}
$$

where we have also used (2.4) and (5.6).
Finally, by taking advantage of (5.9) and (5.6), since $\mathcal{W}^{I} \geq \mathcal{W}$ and since the map $\mathbf{F} \mapsto \int_{\Omega}$ $\mathbf{F}^{T} D^{2} \mathcal{W}(x, \mathbf{I}) \mathbf{F} \mathrm{d} x$ is lower semicontinuous with respect to the weak $L^{2}\left(\Omega, \mathbb{R}^{3 \times 3}\right)$ convergence, we conclude that

$$
\begin{aligned}
& \liminf _{j \rightarrow+\infty} \int_{\Omega} \frac{1}{h_{j}^{2}} \mathcal{W}^{I}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) \mathrm{d} x \geq \liminf _{j \rightarrow+\infty} \int_{D_{j}} \frac{1}{h_{j}^{2}} \mathcal{W}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) \mathrm{d} x \\
& \quad \geq \liminf _{j \rightarrow+\infty} \int_{D_{j}} \frac{1}{2} \nabla \mathbf{v}_{j}^{T} D^{2} \mathcal{W}(x, \mathbf{I}) \nabla \mathbf{v}_{j} \mathrm{~d} x=\liminf _{j \rightarrow+\infty} \int_{\Omega} \frac{1}{2}\left(1_{D_{j}} \nabla \mathbf{v}_{j}\right)^{T} D^{2} \mathcal{W}(x, \mathbf{I})\left(1_{D_{j}} \nabla \mathbf{v}_{j}\right) \\
& \quad \geq \int_{\Omega} \frac{1}{2} \nabla \mathbf{v}^{T} D^{2} \mathcal{W}(x, \mathbf{I}) \nabla \mathbf{v} \mathrm{d} x=\frac{1}{2} \int_{\Omega} \mathbb{E}(\mathbf{v}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v}) \mathrm{d} x .
\end{aligned}
$$

Since functional $\mathcal{L}$ from (2.6) is continuous with respect to the weak convergence in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$, we get

$$
\begin{aligned}
\liminf _{j \rightarrow+\infty} \mathcal{G}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right) & =\liminf _{j \rightarrow+\infty} \int_{\Omega} \frac{1}{h_{j}^{2}} \mathcal{W}^{I}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) \mathrm{d} x-\mathcal{L}\left(\mathbf{v}_{j}\right) \\
& \left.\geq \frac{1}{2} \int_{\Omega} \mathbb{E}(\mathbf{v}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v})\right) \mathrm{d} x-\mathcal{L}(\mathbf{v})=\mathcal{G}^{I}(\mathbf{v}) .
\end{aligned}
$$

Therefore, $\mathcal{G}_{-}^{I}(\mathbf{v})<+\infty$ only if $\mathbf{v} \in H_{\mathrm{div}}^{1}\left(\Omega, \mathbb{R}^{3}\right)$ with $\mathbf{v}=0$ on $\Gamma$, and $\mathcal{G}^{I}(\mathbf{v}) \leq \mathcal{G}_{-}^{I}(\mathbf{v})$ for every $\mathbf{v} \in W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$.

Step 2 (limsup) We show now that $\mathcal{G}_{+}^{I}(\mathbf{v}) \leq \mathcal{G}^{I}(\mathbf{v})$ for every $\mathbf{v} \in W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$.

It will be enough to prove the inequality for every $\mathbf{v} \in H_{\text {div }}^{1}\left(\Omega, \mathbb{R}^{3}\right)$ such that $\mathbf{v}=0$ on $\Gamma$ (otherwise $\left.\mathcal{G}^{I}(\mathbf{v})=+\infty\right)$. This will be done in three subsequent steps, that make use of Lemmas 4.1, 4.4 and 4.9, respectively.

Assume first that $\mathbf{v}$ is the restriction to $\Omega$ of a function $\mathbf{v} \in C^{1}\left(\Omega^{\prime}, \mathbb{R}^{3}\right)$ such that $\mathbf{v}=0$ on $\Gamma$ and $\operatorname{div} \mathbf{v}=0$ in $\Omega^{\prime}$, being $\Omega^{\prime}$ an open set with $\bar{\Omega} \subset \Omega^{\prime}$. By Lemma 4.1 there exists a sequence $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}} \subset$ $C^{1}\left(\bar{\Omega}, \mathbb{R}^{3}\right)$ such that (4.1), (4.2),(4.3) and (4.4) hold. Hence, $(\mathcal{W} 3),(4.1)$ and (4.4) together with $\mathcal{W}(x, \mathbf{I})=$ $0, D \mathcal{W}(x, \mathbf{I})=0$, see (2.4), imply that

$$
\lim _{j \rightarrow+\infty} h_{j}^{-2} \mathcal{W}^{I}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right)=\lim _{j \rightarrow+\infty} h_{j}^{-2} \mathcal{W}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right)=\frac{1}{2} \mathbb{E}(\mathbf{v}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v})
$$

for a.e. $x \in \Omega$, and that there exists a constant $C^{\prime}>0$ such that for $h_{j}$ small enough there holds $h_{j}^{-2} \mathcal{W}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) \leq C^{\prime}\left|\mathbb{E}\left(\mathbf{v}_{j}\right)\right|^{2}$.

Therefore by (4.3) there exist $q>1$ and a constant $C^{\prime \prime}>0$ such that for any large enough $j$

$$
\int_{\Omega}\left(\frac{1}{h_{j}^{2}} \mathcal{W}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right)\right)^{q} \mathrm{~d} x \leq C^{\prime \prime}
$$

thus

$$
\lim _{j \rightarrow \infty} \int_{\Omega} \frac{1}{h_{j}^{2}} \mathcal{W}^{I}\left(x, \mathbf{I}+h_{j} \nabla \mathbf{v}_{j}\right) \mathrm{d} x-\mathcal{L}\left(\mathbf{v}_{j}\right)=\frac{1}{2} \int_{\Omega} \mathbb{E}(\mathbf{v}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v}) \mathrm{d} x-\mathcal{L}(\mathbf{v})
$$

This shows that $\mathcal{G}_{+}^{I}(\mathbf{v}) \leq \mathcal{G}^{I}(\mathbf{v})$ whenever $\mathbf{v}$ is the restriction to $\Omega$ of a function $\mathbf{v} \in C^{1}\left(\Omega^{\prime}, \mathbb{R}^{3}\right)$ such that $\mathbf{v}=0$ on $\Gamma$ and $\operatorname{div} \mathbf{v}=0$ in $\Omega^{\prime}$, being $\Omega^{\prime}$ an open set with $\bar{\Omega} \subset \Omega^{\prime}$.

Assume now that $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$, div $\mathbf{v}=0$ a.e. in $\Omega$ and $\mathbf{v}=0$ on $\Gamma_{\delta}$ for some $0<\delta<\mu_{0}$, where $\mu_{0}$ is defined by (4.12). Then by Lemma 4.4 there exist an open set $\Omega^{\prime}$ such that $\bar{\Omega} \subset \Omega^{\prime}$ and a sequence $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}} \subset C^{1}\left(\Omega^{\prime}, \mathbb{R}^{3}\right)$ such that $\operatorname{div} \mathbf{v}_{j}=0$ a.e. in $\Omega^{\prime}, \mathbf{v}_{j}=0$ on $\Gamma$ and $\mathbf{v}_{j} \rightarrow \mathbf{v}$ in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$. Therefore,

$$
\mathcal{G}_{+}^{I}\left(\mathbf{v}_{j}\right) \leq \mathcal{G}^{I}\left(\mathbf{v}_{j}\right)
$$

By taking into account that $\mathcal{G}_{+}^{I}$ is weakly lower semicontinuous in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ and that $\mathcal{G}^{I}$ is continuous with respect the strong convergence in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ we get

$$
\mathcal{G}_{+}^{I}(\mathbf{v}) \leq \mathcal{G}^{I}(\mathbf{v})
$$

for every $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ such that div $\mathbf{v}=0$ a.e. in $\Omega$ and $\mathbf{v}=0$ on $\Gamma_{\delta}$ for some $0<\delta<\mu_{0}$. If $\partial \Gamma=\emptyset$, then $\Gamma_{\delta}=\Gamma$ and the proof is concluded. Suppose instead that $\partial \Gamma \neq \emptyset$ and let $\mathbf{v} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$, div $\mathbf{v}=0$ a.e. in $\Omega$ and $\mathbf{v}=0$ on $\Gamma$. By Lemma 4.9 there exist a vanishing sequence $\left(\lambda_{j}\right)_{j \in \mathbb{N}} \subset\left(0, \mu_{0}\right)$ and a sequence $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}} \subset H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ such that div $\mathbf{v}_{j}=0$ a.e. in $\Omega, \mathbf{v}_{j}=0$ on $\Gamma_{\lambda_{j}}$ and $\mathbf{v}_{j} \rightarrow \mathbf{v}$ in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$. Then

$$
\mathcal{G}_{+}^{I}\left(\mathbf{v}_{j}\right) \leq \mathcal{G}^{I}\left(\mathbf{v}_{j}\right)
$$

and by exploiting again the weak lower semicontinuity of $\mathcal{G}_{+}^{I}$ in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ and continuity of $\mathcal{G}^{I}$ in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$, we achieve the result.

The proof of the main result directly follows.
Proof of Theorem 2.1. We prove first that $\mathcal{G}^{I}$ has a unique minimizer. Weak $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ compactness of minimizing sequences follows from (2.5) along with Korn and Poincaré inequalities. Along a sequence that converges weakly in $H^{1}\left(\Omega, \mathbb{R}^{3}\right)$, the elastic part of the energy is lower semicontinuous, functional $\mathcal{L}$ is continuous, and the divergence-free constraint passes to the limit as well as the vanishing constraint on $\Gamma$. This shows existence of minimizers of $\mathcal{G}^{I}$. Let us prove uniqueness of minimizers. Let

$$
\begin{equation*}
\mathcal{V}_{0}(x, \mathbf{B}):=\frac{1}{2} \operatorname{sym} \mathbf{B} D^{2} \mathcal{W}(x, \mathbf{I}) \operatorname{sym} \mathbf{B} \tag{5.10}
\end{equation*}
$$

and let $\mathbf{v}_{*}, \mathbf{v}_{* *}$ be two minimizers of $\mathcal{G}^{I}$ (in particular, $\mathbf{v}_{*}=\mathbf{v}_{* *}=0$ on $\Gamma$ ). Then by first order minimality conditions we have

$$
\begin{align*}
& \int_{\Omega} D \mathcal{V}_{0}\left(x, \mathbb{E}\left(\mathbf{v}_{*}\right)\right) \cdot\left(\mathbb{E}\left(\mathbf{v}_{*}\right)-\mathbb{E}\left(\mathbf{v}_{* *}\right)\right) \mathrm{d} x \\
& \quad=\int_{\Omega} D \mathcal{V}_{0}\left(x, \mathbb{E}\left(\mathbf{v}_{* *}\right)\right) \cdot\left(\mathbb{E}\left(\mathbf{v}_{*}\right)-\mathbb{E}\left(\mathbf{v}_{* *}\right)\right) \mathrm{d} x=\mathcal{L}\left(\mathbf{v}_{*}-\mathbf{v}_{* *}\right) . \tag{5.11}
\end{align*}
$$

Hence, by (5.10) and (5.11)

$$
2 \int_{\Omega} \mathcal{V}_{0}\left(x, \mathbb{E}\left(\mathbf{v}_{*}\right)-\mathbb{E}\left(\mathbf{v}_{* *}\right)\right)=\int_{\Omega} D \mathcal{V}_{0}\left(x, \mathbb{E}\left(\mathbf{v}_{*}\right)-\mathbb{E}\left(\mathbf{v}_{* *}\right)\right) \cdot\left(\mathbb{E}\left(\mathbf{v}_{*}\right)-\mathbb{E}\left(\mathbf{v}_{* *}\right)\right) \mathrm{d} x=0
$$

therefore (2.5) implies $\mathbb{E}\left(\mathbf{v}_{*}\right)-\mathbb{E}\left(\mathbf{v}_{* *}\right)=0$. Since $\mathbf{v}_{*}-\mathbf{v}_{* *}=0$ on $\Gamma$, we deduce $\mathbf{v}_{*}-\mathbf{v}_{* *}=0$ a.e. on $\Omega$ thus proving uniqueness. From now we denote by $\mathbf{v}_{*}$ the unique minimizer of $\overline{\mathcal{G}}^{I}$.

By testing with the trivial displacement field, we see that $\inf \mathcal{G}_{h_{j}}^{I} \leq 0$ for any $j \in \mathbb{N}$. On the other hand, since $\mathcal{L}(\mathbf{v}) \leq C_{\mathcal{L}}\|\mathbf{v}\|_{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)}$, boundedness from below of functional $\mathcal{G}_{h_{j}}^{I}$ easily follows from (5.3) and Friedrichs inequality as soon as $j$ is large enough. This proves (2.8).

The sequence $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}}$ is bounded in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$, thanks to Lemma 5.2. Therefore, let us consider a (not relabeled) subsequence such that $\mathbf{v}_{j} \rightharpoonup \mathbf{v}$ weakly in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$. Let $\widetilde{\mathbf{v}} \in H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ be such that $\widetilde{\mathbf{v}}=0$ on $\Gamma$ and $\operatorname{div} \widetilde{\mathbf{v}}=0$ a.e. in $\Omega$. Let $\left(\widetilde{\mathbf{v}}_{j}\right)_{j \in \mathbb{N}} \subset W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ be a recovery sequence for $\widetilde{\mathbf{v}}$, provided by Lemma 5.3. By taking advantage of (2.9) and of the $\Gamma$-liminf inequality, still provided by Lemma 5.3, we conclude that

$$
\mathcal{G}^{I}(\mathbf{v}) \leq \liminf _{j \rightarrow+\infty} \mathcal{G}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right) \leq \limsup _{j \rightarrow+\infty} \mathcal{G}_{h_{j}}^{I}\left(\widetilde{\mathbf{v}}_{j}\right)=\mathcal{G}^{I}(\widetilde{\mathbf{v}})
$$

By the arbitrariness of $\widetilde{\mathbf{v}}$ we get $\mathbf{v} \in \operatorname{argmin} \mathcal{G}^{I}$ hence $\mathbf{v}=\mathbf{v}_{*}$ and the whole sequence $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}}$ converges to $\mathbf{v}_{*}$ weakly in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$ thus concluding the proof.

The proof of Corollary 2.2 relies on the following preliminary result.
Lemma 5.4. Under the assumptions of Corollary 2.2, let $\overline{\mathbf{v}} \in W^{1, \infty}\left(\Omega, \mathbb{R}^{3}\right)$ be such that $\operatorname{div} \overline{\mathbf{v}}=0$ a.e. in $\Omega$. Then $\overline{\mathcal{G}}^{I}$ from (2.10) has a unique minimizer and if $\mathbf{v}_{*} \in \operatorname{argmin} \overline{\mathcal{G}}^{I}$ then $\mathbf{v}_{*}-\overline{\mathbf{v}}$ is the unique minimizer of $\widetilde{\mathcal{G}}^{I}$ and $\widetilde{\mathcal{G}}^{I}\left(\mathbf{v}_{*}-\overline{\mathbf{v}}\right)=\overline{\mathcal{G}}^{I}\left(\mathbf{v}_{*}\right)$, where $\widetilde{\mathcal{G}}^{I}$ is defined by (2.11).
Proof. Existence of a minimizer of $\overline{\mathcal{G}}^{I}$ and of $\widetilde{\mathcal{G}}^{I}$ again follows from classical results while regarding uniqueness of minimizers of $\overline{\mathcal{G}}^{I}$ and of $\widetilde{\mathcal{G}}^{I}$ we may argue as in the proof of Theorem 2.1 and from now we denote by $\mathbf{v}_{*}$ the unique minimizer of $\overline{\mathcal{G}}^{I}$.

Let $\mathbf{u} \in H_{\text {div }}^{1}\left(\Omega, \mathbb{R}^{3}\right)$ be such that $\mathbf{u}=0$ on $\Gamma$, and set $\mathbf{v}=\mathbf{u}+\overline{\mathbf{v}}$. Then $\mathbf{v}=\overline{\mathbf{v}}$ on $\Gamma, \operatorname{div} \mathbf{v}=0$ a.e. in $\Omega$ and by using (5.10)

$$
\begin{aligned}
\int_{\Omega} \mathcal{V}_{0} & \left(x, \mathbb{E}\left(\mathbf{v}_{*}\right)-\mathbb{E}(\overline{\mathbf{v}})\right)+\int_{\Omega} D \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}})) \cdot\left(\mathbb{E}\left(\mathbf{v}_{*}\right)-\mathbb{E}(\overline{\mathbf{v}})\right) \mathrm{d} x-\mathcal{L}\left(\mathbf{v}_{*}-\overline{\mathbf{v}}\right) \\
= & \int_{\Omega} \mathcal{V}_{0}\left(x, \mathbb{E}\left(\mathbf{v}_{*}\right)\right) \mathrm{d} x+\int_{\Omega} \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}}))-\int_{\Omega} D \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}})) \cdot \mathbb{E}(\overline{\mathbf{v}}) \mathrm{d} x-\mathcal{L}\left(\mathbf{v}_{*}-\overline{\mathbf{v}}\right) \\
\leq & \int_{\Omega} \mathcal{V}_{0}(x, \mathbb{E}(\mathbf{v})) \mathrm{d} x+\int_{\Omega} \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}}))-\int_{\Omega} D \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}})) \cdot \mathbb{E}(\overline{\mathbf{v}}) \mathrm{d} x \\
& +\int_{\Omega} D \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}})) \cdot \mathbb{E}(\mathbf{v}) \mathrm{d} x-\int_{\Omega} D \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}})) \cdot \mathbb{E}(\mathbf{v}) \mathrm{d} x-\mathcal{L}(\mathbf{v}-\overline{\mathbf{v}})
\end{aligned}
$$

$$
=\int_{\Omega} \mathcal{V}_{0}(x, \mathbb{E}(\mathbf{v})-\mathbb{E}(\overline{\mathbf{v}})) \mathrm{d} x+\int_{\Omega} D \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}})) \cdot(\mathbb{E}(\mathbf{v})-\mathbb{E}(\overline{\mathbf{v}})) \mathrm{d} x-\mathcal{L}(\mathbf{v}-\overline{\mathbf{v}}),
$$

that is, $\widetilde{\mathcal{G}}^{I}\left(\mathbf{v}_{*}-\overline{\mathbf{v}}\right) \leq \widetilde{\mathcal{G}}^{I}(\mathbf{u})$, thus proving minimality of $\mathbf{v}_{*}-\overline{\mathbf{v}}$ for $\widetilde{\mathcal{G}}^{I}$ by the arbitrariness of $\mathbf{u}$. Uniqueness of such a minimizer follows by reasoning as in the first part of this proof so we have only to prove that $\widetilde{\mathcal{G}}^{I}\left(\mathbf{v}_{*}-\overline{\mathbf{v}}\right)=\overline{\mathcal{G}}^{I}\left(\mathbf{v}_{*}\right)$. Indeed

$$
\begin{aligned}
\widetilde{\mathcal{G}}^{I}\left(\mathbf{v}_{*}-\overline{\mathbf{v}}\right)= & \int_{\Omega} \mathcal{V}_{0}\left(x, \mathbb{E}\left(\mathbf{v}_{*}\right)\right) \mathrm{d} x+\int_{\Omega} \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}})) \mathrm{d} x-\int_{\Omega} D \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}})) \cdot \mathbb{E}\left(\mathbf{v}_{*}\right) \mathrm{d} x \\
& -\mathcal{L}\left(\mathbf{v}_{*}-\overline{\mathbf{v}}\right)+\int_{\Omega} D \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}})) \cdot\left(\mathbb{E}\left(\mathbf{v}_{*}\right)-\mathbb{E}(\overline{\mathbf{v}})\right) \mathrm{d} x+\overline{\mathcal{G}}^{I}(\overline{\mathbf{v}}) \\
= & \int_{\Omega} \mathcal{V}_{0}\left(x, \mathbb{E}\left(\mathbf{v}_{*}\right)\right) \mathrm{d} x-\int_{\Omega} \mathcal{V}_{0}(x, \mathbb{E}(\overline{\mathbf{v}})) \mathrm{d} x-\mathcal{L}\left(\mathbf{v}_{*}-\overline{\mathbf{v}}\right)+\overline{\mathcal{G}}^{I}(\overline{\mathbf{v}})=\overline{\mathcal{G}}^{I}\left(\mathbf{v}_{*}\right)
\end{aligned}
$$

and the proof is concluded.
Proof of Corollary 2.2. Since the map

$$
W^{1, p}\left(\Omega, \mathbb{R}^{3}\right) \ni \mathbf{v} \mapsto \int_{\Omega} \mathbb{E}(\overline{\mathbf{v}}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v}) \mathrm{d} x
$$

is continuous with respect to the weak topology of $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$, Lemma 5.3 implies the $\Gamma$-convergence of functionals $\widetilde{\mathcal{G}}_{h_{j}}^{I}$ to $\widetilde{\mathcal{G}}^{I}$ with respect to the same topology.

We notice that $\widetilde{\mathcal{G}}_{h_{j}}^{I}(\mathbf{0}) \leq \overline{\mathcal{G}}^{I}(\overline{\mathbf{v}})$ so that $\inf \widetilde{\mathcal{G}}_{h_{j}}^{I}<+\infty$ for any $j \in \mathbb{N}$, where the infimum is taken on $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$. Since $\overline{\mathcal{G}}^{I}(\overline{\mathbf{v}}) \in \mathbb{R}$ and since by assumption $\mathcal{W} 3$ there holds

$$
\begin{equation*}
\int_{\Omega} \mathbb{E}(\overline{\mathbf{v}}) D^{2} \mathcal{W}(x, \mathbf{I}) \mathbb{E}(\mathbf{v}) \mathrm{d} x \leq K|\Omega|^{\frac{p-1}{p}}\|\nabla \overline{\mathbf{v}}\|_{L^{\infty}\left(\Omega, \mathbb{R}^{3 \times 3}\right)}\|\mathbf{v}\|_{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)} \tag{5.12}
\end{equation*}
$$

by the same reasoning of the proof of Theorem 2.1 we deduce boundedness from below of $\widetilde{\mathcal{G}}_{h_{j}}^{I}$ for any large enough $j$ so that (2.12) holds.

Let now $\left(\mathbf{v}_{j}\right)_{j \in \mathbb{N}} \subset H^{1}\left(\Omega, \mathbb{R}^{3}\right)$ be a sequence such that $\mathbf{v}_{j}=0$ on $\Gamma$ and such that (2.13) holds. By the same argument of the proof of Lemma 5.2, this time also taking (5.12) into account, we deduce that $\sup _{j \in \mathbb{N}}\left\|\mathbf{v}_{j}\right\|_{W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)}<+\infty$. Therefore, up to not relabeled subsequences, $\mathbf{v}_{j} \rightharpoonup \mathbf{v}_{0}$ weakly in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$. Thanks to $\Gamma$-convergence, by the same argument of the proof of Theorem 2.1, we conclude that $\widetilde{\mathcal{G}}_{h_{j}}^{I}\left(\mathbf{v}_{j}\right) \rightarrow \widetilde{\mathcal{G}}^{I}\left(\mathbf{v}_{0}\right)$ as $j \rightarrow+\infty$ and that $\mathbf{v}_{0}$ is the unique minimizer of $\widetilde{\mathcal{G}}^{I}$ over $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$. In particular, the whole sequence $\left(\mathbf{v}_{j}\right)$ converges to $\mathbf{v}_{0}$ weakly in $W^{1, p}\left(\Omega, \mathbb{R}^{3}\right)$. By Lemma 5.4, $\mathbf{v}_{0}+\overline{\mathbf{v}} \in$ $\operatorname{argmin} \overline{\mathcal{G}}^{I}$ and $\widetilde{\mathcal{G}}^{I}\left(\mathbf{v}_{0}\right)=\overline{\mathcal{G}}^{I}\left(\mathbf{v}_{0}+\overline{\mathbf{v}}\right)$ so that we have recovered the unique minimizer of $\overline{\mathcal{G}}^{I}$, thus concluding the proof.

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Edoardo Mainini and Danilo Percivale
Dipartimento di Ingegneria Meccanica, Energetica, Gestionale e dei Trasporti
Università degli Studi di Genova
Via all'Opera Pia, 15
16145 Genova
Italy
e-mail: mainini@dime.unige.it
Danilo Percivale
e-mail: percivale@dime.unige.it
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