## On the decay mode $\boldsymbol{\Lambda}_{\mathrm{b}} \rightarrow \mathbf{X}_{\mathbf{s}} \boldsymbol{\gamma}$

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Abstract: We study the inclusive $H_{b} \rightarrow X_{s} \gamma$ decay with $H_{b}$ a beauty baryon, in particular $\Lambda_{b}$, employing an expansion in the heavy quark mass at $\mathcal{O}\left(m_{b}^{-3}\right)$ at leading order in $\alpha_{s}$, keeping the dependence on the hadron spin. For a polarized baryon we compute the distribution $\frac{d^{2} \Gamma}{d y d \cos \theta_{P}}$, with $y=2 E_{\gamma} / m_{b}, E_{\gamma}$ the photon energy and $\theta_{P}$ the angle between the baryon spin vector and the photon momentum in the $H_{b}$ rest-frame. We discuss the correlation between the baryon and photon polarization, and show that effects of physics beyond the Standard Model can modify the photon polarization asymmetry. We also discuss a method to treat the singular terms in the photon energy spectrum obtained by the OPE.

Keywords: Bottom Quarks, Nonperturbative Effects, Rare Decays

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## Contents

1 Introduction ..... 1
$2 b \rightarrow s \gamma$ effective Hamiltonian ..... 3
3 Inclusive decay width ..... 5
4 Photon polarization ..... 8
5 Treatment of the singular terms ..... 10
6 Conclusions ..... 15
A Derivation of the OPE ..... 16
B Factorization formula in the endpoint region and shape functions ..... 18
C Moments of the photon energy spectrum ..... 18

## 1 Introduction

The processes induced at the quark level by the $b \rightarrow s \gamma$ transition are recognized as a powerful testground of the Standard Model (SM) [1-3]. They occur at loop-level in SM and are sensitive to heavy particle exchanges. Upon integration of the heavy quanta, an effective Hamiltonian is obtained in terms of local operators and Wilson coefficients [4, 5]. Physics beyond the Standard Model induces new operators with respect to the SM ones and modifies the Wilson coefficients, hence measurements of various observables tightly constrain operators and coefficients. This allows to probe the SM and select the possible extensions [6-13].

The radiative $b \rightarrow s$ transition has been intensively analyzed in theory. ${ }^{1}$ Strong experimental efforts have also been devoted since the first observation of the $B \rightarrow K^{*}(892) \gamma$ mode [15]. Several measurements of exclusive processes are now available, namely $\mathcal{B}\left(B^{+} \rightarrow\right.$ $\left.K^{*+}(892) \gamma\right)$ and $\mathcal{B}\left(B^{0} \rightarrow K^{* 0}(892) \gamma\right)$, and the rates of $B \rightarrow\left(K_{1}(1400), K_{2}^{*}(1430), K_{3}^{*}(1780)\right) \gamma$, $B \rightarrow K \eta \gamma, B_{s} \rightarrow \phi(1020) \gamma[16,17]$. Time-dependent CP asymmetries for decaying neutral mesons have been investigated [16]. For baryons, the rate and the photon polarization of $\Lambda_{b} \rightarrow \Lambda \gamma$ have been measured [18, 19], and an upper bound has been put to $\mathcal{B}\left(\Xi_{b}^{-} \rightarrow \Xi^{-} \gamma\right)[20]$. For such exclusive processes, the hadronic uncertainties are related to the form factors parametrizing the matrix elements of the local operators at $q^{2}=0$, with $q$ the photon momentum [21].

[^0]Among the inclusive $H_{b} \rightarrow X_{s} \gamma$ modes the prime example is $\bar{B} \rightarrow X_{s} \gamma$. Here we focus on $H_{b}$ a baryon, in particular on $\Lambda_{b} \rightarrow X_{s} \gamma$. The peculiarity of the inclusive modes consists in the possibility, invoking quark-hadron duality, of exploiting a well defined theoretical framework based on controlled expansions in QCD quantities $\alpha_{s}\left(m_{b}\right)$ and $1 / m_{b}$ (the heavy quark expansion, HQE ) to compute decay rates, decay distributions and moments of the distributions. At the leading order in the heavy quark expansion the partonic result is recovered, NLO terms involve nonperturbative corrections. In particular, the combination of Operator Product Expansion (OPE) and Heavy Quark Effective Theory (HQET) [22] allows to express the inclusive decay widths as an expansion in the inverse heavy quark mass. Based on this approach, inclusive semileptonic modes of hadrons comprising a single heavy quark are exploited to access fundamental parameters, the heavy quark masses and elements of the Cabibbo-Kobayashi-Maskawa (CKM) mixing matrix. Input quantities are the hadronic matrix elements of local operators, starting from the kinetic energy and the chromomagnetic operators, which are defined in the following. Differential distributions together with other observables can also be described. In $B \rightarrow X_{s} \gamma$, direct CP asymmetries sensitive to new physics (NP) effects can be studied upon accounting for long-distance contributions [23, 24].

There are several issues in inclusive modes induced by the $b \rightarrow s \gamma$ transition needing to be considered. The actual expansion parameter is the inverse of the energy released in the process. Such energy is $\mathcal{O}\left(m_{b}\right)$ in a portion of the phase-space, but in some regions its inverse is no longer small. Signals about the reliability of the method show up as singularities in differential distributions. This occurs in the calculation of perturbative corrections to the spectra, where Sudakov terms appear. ${ }^{2}$ Singular terms appear at higher orders in the HQE in form of the delta distribution and its derivatives, the argument of which vanishes in the regions corresponding to the endpoints of the differential distributions determined in the partonic kinematics, different from the hadronic kinematics. The gap between the two borders is governed by nonperturbative physics responsible of bound state effects. Such effects can be related to the Fermi motion of the heavy quark in the decaying hadron, and can be accounted for introducing a shape function which encodes information on the distribution of the $b$ quark residual momentum in the hadron [27-29]. The same function enters in the description of different inclusive modes, and it affects the photon energy spectrum in $H_{b} \rightarrow X_{s} \gamma$. In case of $B$, the moments of the shape function have been constrained using measurements $[16,30]$.

Another source of uncertainty in inclusive $b \rightarrow s \gamma$ processes are the resolved photon contributions [31], related to the photon couplings different from the effective weak interaction vertex. The most important operators giving rise to these contributions are $O_{2}$ and $O_{8}$ (in the notation specified in the following section). Such effects appear at $\mathcal{O}\left(1 / m_{b}\right)$ and produce contributions to the total decay width which are not described by the $\mathrm{HQE}[32,33]$. They can be expressed in terms of subleading shape functions [31]. The resolved photon contributions include a nonperturbative term proportional to the matrix element $\mu_{G}^{2}$ of the

[^1]chromomagnetic operator, related to the gluon-photon penguin mechanism, for which the short distance scale is $1 / m_{c}$ rather than $1 / m_{b}$ [34-37]. This is estimated to be small in $\bar{B} \rightarrow X_{s} \gamma$. For $\Lambda_{b}$ the matrix element of the chromomagnetic operator vanishes, therefore the comparison of the inclusive $\Lambda_{b}$ and $B$ radiative decay widths provides a way to shed light on the role of such corrections.

In radiative modes the photon spectrum can be measured above an energy threshold. For $\bar{B}$ the HFLAV collaboration quotes $\mathcal{B}\left(\bar{B} \rightarrow X_{s} \gamma\right)=(3.49 \pm 0.19) \times 10^{-4}$ for $E_{\gamma}>1.6$ GeV , with both charged and neutral mesons included in the average. The SM result is $\mathcal{B}\left(\bar{B} \rightarrow X_{s} \gamma\right)=(3.36 \pm 0.23) \times 10^{-4}$ for the same threshold [5]. A global fit including all data on the photon energy spectrum in $\bar{B} \rightarrow X_{s} \gamma$ has been performed by the SIMBA collaboration [30].

The interest for inclusive $\Lambda_{b} \rightarrow X_{s} \gamma$ relies on the possibility in baryon modes to investigate observables sensitive to the spin of the decaying hadron. This is important for the planned new lepton facilities, since heavy baryons with a b-quark produced from $Z^{0}$ and top-quark decays are expected to have a sizable polarization, as observed at LEP [3840]. The application of HQE to baryons requires new information, namely the operator matrix elements for specified hadron spin. Such matrix elements have been analyzed in [41]. Moreover, the leading and subleading shape functions are different for different hadrons, namely for $\Lambda_{b}$ and $B$, and require dedicated considerations.

In the present study we focus on two issues. The first one is the dependence on the heavy baryon spin in a double differential decay distribution, considering hadronic matrix elements at $\mathcal{O}\left(1 / m_{b}^{3}\right)$, for the leading operator in the SM effective weak Hamiltonian and for a single new physics (NP) operator, studying the correlations between the baryon and photon polarization. The second one is a way to treat the singular terms in the inclusive photon spectrum to reconstruct the $\Lambda_{b}$ leading shape function, a method which can be systematically applied when higher order terms in the heavy quark expansion are computed.

The plan of the paper is the following. Section 2 includes the $b \rightarrow s \gamma$ low-energy Hamiltonian with SM operators and operators obtained in extensions of the Standard Model. In section 3 we describe the application of the HQE to the inclusive $H_{b} \rightarrow X_{s} \gamma$ process with $H_{b}$ a baryon, in particular $\Lambda_{b}$, keeping the dependence on the baryon spin. In section 4 we investigate the correlation between the photon and $\Lambda_{b}$ polarizations. A treatment of the singular terms in the double differential decay rate is discussed in section 5 . Details are collected in the appendices. In the last section we present our conclusions and the perspectives for further progress.

## $2 \quad b \rightarrow s \gamma$ effective Hamiltonian

The low-energy Hamiltonian governing the $\Delta B=-1, \Delta S=1 b \rightarrow s \gamma$ transition can be written as

$$
\begin{equation*}
H_{\mathrm{eff}}^{b \rightarrow s \gamma}=-4 \frac{G_{F}}{\sqrt{2}} V_{t b} V_{t s}^{*} \sum_{i}\left[C_{i}(\mu) O_{i}+C_{i}^{\prime}(\mu) O_{i}^{\prime}\right] \tag{2.1}
\end{equation*}
$$

with $i=1, \ldots 8$ and $i=15, \ldots 20 . G_{F}$ is the Fermi constant and $V_{j k}$ are elements of the CKM matrix. Doubly Cabibbo-suppressed terms proportional to $V_{u b} V_{u s}^{*}$ have been
neglected in (2.1). The effective Hamiltonian comprises the magnetic penguin operators

$$
\begin{align*}
& O_{7}=\frac{e}{16 \pi^{2}}\left[\bar{s} \sigma^{\mu \nu}\left(m_{s} P_{L}+m_{b} P_{R}\right) b\right] F_{\mu \nu},  \tag{2.2}\\
& O_{8}=\frac{g_{s}}{16 \pi^{2}}\left[\bar{s}_{\alpha} \sigma^{\mu \nu}\left(\frac{\lambda^{a}}{2}\right)_{\alpha \beta}\left(m_{s} P_{L}+m_{b} P_{R}\right) b_{\beta}\right] G_{\mu \nu}^{a}, \tag{2.3}
\end{align*}
$$

with $P_{R, L}=\frac{1 \pm \gamma_{5}}{2}$ helicity projectors, $\alpha, \beta$ colour indices, $\lambda^{a}$ the Gell-Mann matrices. $F_{\mu \nu}$ and $G_{\mu \nu}^{a}$ are the electromagnetic and gluonic field strengths, $e$ and $g_{s}$ the electromagnetic and strong coupling constants, $m_{b}$ and $m_{s}$ the $b$ and $s$ quark mass. The Hamiltonian also comprises the current-current operators $O_{1,2}$,

$$
\begin{align*}
& O_{1}=\left(\bar{s}_{\alpha} \gamma^{\mu} P_{L} c_{\beta}\right)\left(\bar{c}_{\beta} \gamma_{\mu} P_{L} b_{\alpha}\right),  \tag{2.4}\\
& O_{2}=\left(\bar{s} \gamma^{\mu} P_{L} c\right)\left(\bar{c} \gamma_{\mu} P_{L} b\right), \tag{2.5}
\end{align*}
$$

and the QCD penguin operators $O_{i=3, \ldots 6}$,

$$
\begin{align*}
& O_{3}=\left(\bar{s} \gamma^{\mu} P_{L} b\right) \sum_{q}\left(\bar{q} \gamma^{\mu} P_{L} q\right), O_{4}=\left(\bar{s}_{\alpha} \gamma^{\mu} P_{L} b_{\beta}\right) \sum_{q}\left(\bar{q}_{\beta} \gamma^{\mu} P_{L} q_{\alpha}\right),  \tag{2.6}\\
& O_{5}=\left(\bar{s} \gamma^{\mu} P_{L} b\right) \sum_{q}\left(\bar{q} \gamma^{\mu} P_{R} q\right), O_{6}=\left(\bar{s}_{\alpha} \gamma^{\mu} P_{L} b_{\beta}\right) \sum_{q}\left(\bar{q}_{\beta} \gamma^{\mu} P_{R} q_{\alpha}\right) . \tag{2.7}
\end{align*}
$$

The sum in (2.6)-(2.7) runs over the flavours $q=u, d, s, c, b$. The remaining operators, absent in SM, are analogous to the QCD penguins but have a scalar or tensor structure [42]:

$$
\begin{array}{ll}
O_{15}^{q}=\left(\bar{s} P_{R} b\right) \sum_{q}\left(\bar{q} P_{R} q\right), & O_{16}^{q}=\left(\bar{s}_{\alpha} P_{R} b_{\beta}\right) \sum_{q}\left(\bar{q}_{\beta} P_{R} q_{\alpha}\right), \\
O_{17}^{q}=\left(\bar{s} P_{R} b\right) \sum_{q}\left(\bar{q} P_{L} q\right), & O_{18}^{q}=\left(\bar{s}_{\alpha} P_{R} b_{\beta}\right) \sum_{q}\left(\bar{q}_{\beta} P_{L} q_{\alpha}\right), \\
O_{19}^{q}=\left(\bar{s} \sigma^{\mu \nu} P_{R} b\right) \sum_{q}\left(\bar{q} \sigma_{\mu \nu} P_{R} q\right), & O_{20}^{q}=\left(\bar{s}_{\alpha} \sigma^{\mu \nu} P_{R} b_{\beta}\right) \sum_{q}\left(\bar{q}_{\beta} \sigma_{\mu \nu} P_{R} q_{\alpha}\right) . \tag{2.8}
\end{array}
$$

The primed operators have opposite chirality with respect to the unprimed ones.
In SM the process $b \rightarrow s \gamma$ is described by photon penguin diagrams, with the photon coupled either to the intermediate fermion or to the $W^{ \pm}$, giving rise to the magnetic operator $O_{7}$. This is the only operator contributing at lowest order in QCD. The renormalization group evolution to the scale $\mu_{b} \simeq \mathcal{O}\left(m_{b}\right)$ also involves the magnetic gluon penguin operator $O_{8}$ and the operators $O_{1, \ldots 6}$. Their mixing into $O_{7}$ generates large logarithms producing a strong enhancement of the rate. The anomalous dimension matrix governing the mixing turns out to be regularization scheme dependent. One can get rid of such a dependence defining an effective coefficient $C_{7}^{\text {eff }}\left(\mu_{b}\right)$ which includes contributions of $O_{1, \ldots 6}$ [43]. In this way $O_{7}$ turns out to be the dominant contribution to $b \rightarrow s \gamma$, with the SM Wilson coefficients known at NNLO in QCD [14, 44]. Extensions of the SM can also induce the operators $O_{15}^{q}-O_{20}^{q}$ and the primed operators in the low-energy Hamiltonian.

In this paper we work at the leading order in $\alpha_{s}$ so that the only operator mediating the $b \rightarrow s \gamma$ transition is $O_{7}$ in SM and possibly $O_{7}^{\prime}$ beyond SM, the effect of $O_{1, \ldots 6}^{(\prime)}$ being
included in the effective coefficients. Therefore, we consider the effective Hamiltonian at the scale $\mu_{b}$ consisting of only two operators,

$$
\begin{equation*}
H_{\mathrm{eff}}^{b \rightarrow s \gamma}=-4 \frac{G_{F}}{\sqrt{2}} V_{t b} V_{t s}^{*}\left\{C_{7}^{\mathrm{eff}} O_{7}+C_{7}^{\prime \text { eff }} O_{7}^{\prime}\right\} . \tag{2.9}
\end{equation*}
$$

We do not consider operators from NLO electroweak corrections.
$H_{\mathrm{eff}}^{b \rightarrow s \gamma}$ can be recast in a way suitable for the heavy quark expansion:

$$
\begin{equation*}
H_{\mathrm{eff}}^{b \rightarrow s \gamma}=-4 \frac{G_{F}}{\sqrt{2}} \lambda_{t} \frac{e}{16 \pi^{2}} \sum_{i=7,7^{\prime}} C_{i}^{\mathrm{eff}} J_{\mu \nu}^{i} F^{\mu \nu} \tag{2.10}
\end{equation*}
$$

where $\lambda_{t}=V_{t b} V_{t s}^{*}, J_{\mu \nu}^{i}=\left[\bar{s} \sigma_{\mu \nu}\left(m_{s}\left(1-P_{i}\right)+m_{b} P_{i}\right) b\right]$ and $P_{i}=P_{R}$ for $i=7, P_{i}=P_{L}$ for $i=7^{\prime}$. In the next section we compute the inclusive width $\Gamma\left(H_{b} \rightarrow X_{s} \gamma\right)$ using the Hamiltonian (2.10), as done for the semileptonic modes in [41].

## 3 Inclusive decay width

To describe the inclusive mode $H_{b}(p, s) \rightarrow X_{s}\left(p_{X}\right) \gamma(q, \epsilon)$ we preliminarly define

$$
\begin{equation*}
\mathcal{F}^{M N}=\mathcal{F}^{\mu \nu \mu^{\prime} \nu^{\prime}}=\sum_{\epsilon} 4 q^{\nu} q^{\nu^{\prime}} \epsilon^{\mu} \epsilon^{* \mu^{\prime}}=-4 q^{\nu} q^{\nu^{\prime}} g^{\mu \mu^{\prime}} \tag{3.1}
\end{equation*}
$$

where $q$ and $\epsilon$ are the photon momentum and polarization four-vector, respectively, using the compact notation $M=\mu \nu, N=\mu^{\prime} \nu^{\prime}$. To obtain the results specifying the photon polarization we also define

$$
\begin{align*}
& \mathcal{F}_{+}^{M N}=4 q^{\nu} q^{\nu^{\prime}} \epsilon_{+}^{\mu} \epsilon_{+}^{* \mu^{\prime}} \\
& \mathcal{F}_{-}^{M N}=4 q^{\nu} q^{\nu^{\prime}} \epsilon_{-}^{\mu} \epsilon_{-}^{* \mu^{\prime}} \tag{3.2}
\end{align*}
$$

where

$$
\begin{equation*}
\epsilon_{ \pm}=\mp \frac{1}{\sqrt{2}}(0,1, \pm i, 0) \tag{3.3}
\end{equation*}
$$

The differential inclusive decay width can be written as

$$
\begin{equation*}
d \Gamma=[d q] \frac{G_{F}^{2}\left|\lambda_{t}\right|^{2}}{8 m_{H_{b}}} \frac{\alpha}{\pi^{2}} \sum_{i, j=7,7^{\prime}} C_{i}^{\mathrm{eff} *} C_{j}^{\mathrm{eff}} W_{M N}^{i j} \mathcal{F}^{M N}, \tag{3.4}
\end{equation*}
$$

with $[d q]=\frac{d^{3} q}{(2 \pi)^{3} 2 q^{0}}$. By the optical theorem, the hadronic tensor $W_{M N}^{i j}$ is related to the discontinuity of the forward scattering amplitude

$$
\begin{equation*}
T_{M N}^{i j}=i \int d^{4} x e^{-i q \cdot x}\left\langle H_{b}(p, s)\right| T\left[J_{M}^{i \dagger}(x) J_{N}^{j}(0)\right]\left|H_{b}(p, s)\right\rangle \tag{3.5}
\end{equation*}
$$

across the cut corresponding to the process $H_{b}(p, s) \rightarrow X_{s}\left(p_{X}\right) \gamma(q, \epsilon)$ :

$$
\begin{equation*}
W_{M N}^{i j}=\frac{1}{\pi} \operatorname{Im} T_{M N}^{i j} \tag{3.6}
\end{equation*}
$$

The range of the invariant mass $p_{X}^{2}$ of the states produced in $B$ and $\Lambda_{b}$ decays (with $\left.p_{X}=p-q\right)$ is $p_{X}^{2} \in\left[m_{K^{*}}^{2}, m_{B}^{2}\right]$ and $p_{X}^{2} \in\left[m_{\Lambda}^{2}, m_{\Lambda_{b}}^{2}\right]$, respectively. For $m_{b} \rightarrow \infty, p_{X}^{2}$ is
almost always large enough to exploit the short distance limit $x \rightarrow 0$ in eq. (3.5), thus allowing a computation of $T^{i j}$ and $W^{i j}$ by an OPE with expansion parameter $\frac{1}{m_{b}}[45,46]$. The first term of the expansion describes the free beauty quark decay, the partonic result. The expansion is valid in the largest part of the phase-space, it fails in the region with small $p_{X}^{2}$, therefore a reliable computation of the decay width and of moments of the decay distributions can be carried out. For spectra, the result obtained by the short distance OPE needs to be smeared: in section 5 we discuss a way to implement the smearing.

The procedure for the derivation of the OPE for eq. (3.5) is summarized in appendix A. Using the definition

$$
\begin{equation*}
\tilde{\mathcal{T}}^{i j}=T_{M N}^{i j} \mathcal{F}^{M N} \tag{3.7}
\end{equation*}
$$

we obtain the expression:

$$
\begin{align*}
\sum_{i, j=7,7^{\prime}} C_{i}^{\text {eff* }} C_{j}^{\text {eff }} \tilde{\mathcal{T}}^{i j}= & {\left[\left(m_{b}^{2}+m_{s}^{2}\right)\left(\left|C_{7}^{\text {eff }}\right|^{2}+\left|C_{7}^{\prime \text { eff }}\right|^{2}\right)+4 m_{b} m_{s} \operatorname{Re}\left[C_{7}^{\text {eff }} C_{7}^{\prime \text { eff } *}\right]\right] \tilde{T} } \\
& +\left(m_{b}^{2}-m_{s}^{2}\right)\left(\left|C_{7}^{\text {eff }}\right|^{2}-\left|C_{7}^{\prime \text { eff }}\right|^{2}\right) \tilde{S} \tag{3.8}
\end{align*}
$$

where

$$
\begin{align*}
& \tilde{T}=16 m_{H}(v \cdot q)^{2} \sum_{n=1}^{4}\left(\frac{m_{b}}{\Delta_{0}}\right)^{n} \tilde{T}_{n}, \\
& \tilde{S}=16 m_{H}(v \cdot q)(q \cdot s) \sum_{n=1}^{4}\left(\frac{m_{b}}{\Delta_{0}}\right)^{n} \tilde{S}_{n}, \tag{3.9}
\end{align*}
$$

with $v$ and $\Delta_{0}$ defined in appendix A . The four terms in (3.9) involve the hadronic parameters $\hat{\mu}_{\pi}^{2}, \hat{\mu}_{G}^{2}, \hat{\rho}_{D}^{3}$ and $\hat{\rho}_{L S}^{3}$ defined in the same appendix, and read:

$$
\begin{align*}
& \tilde{T}_{1}=1+\frac{5}{6 m_{b}^{2}}\left[\hat{\mu}_{\pi}^{2}-\hat{\mu}_{G}^{2}\right]-\frac{2}{3 m_{b}^{3}}\left[\hat{\rho}_{D}^{3}+\hat{\rho}_{L S}^{3}\right],  \tag{3.10}\\
& \tilde{S}_{1}=1+\frac{1}{4 m_{b}^{2}}\left[\hat{\mu}_{\pi}^{2}+\hat{\mu}_{G}^{2}\right]+\frac{1}{6 m_{b}^{3}} \hat{\rho}_{D}^{3},  \tag{3.11}\\
& \tilde{T}_{2}=\frac{7 v \cdot q}{3 m_{b}^{2}} \hat{\mu}_{\pi}^{2}+\frac{1}{3 m_{b}^{2}}\left(4 m_{b}-5 v \cdot q\right) \hat{\mu}_{G}^{2}+\frac{2}{3 m_{b}^{3}}\left[\left(4 m_{b}-3 v \cdot q\right) \hat{\rho}_{D}^{3}+\left(2 m_{b}-3 v \cdot q\right) \hat{\rho}_{L S}^{3}\right],  \tag{3.12}\\
& \tilde{S}_{2}=\frac{7 v \cdot q}{3 m_{b}^{2}} \hat{\mu}_{\pi}^{2}+\frac{1}{m_{b}^{2}}\left(2 m_{b}-v \cdot q\right) \hat{\mu}_{G}^{2}+\frac{2}{3 m_{b}^{3}}\left(4 m_{b}-v \cdot q\right) \hat{\rho}_{D}^{3},  \tag{3.13}\\
& \tilde{T}_{3}=\frac{4}{3 m_{b}^{2}}(v \cdot q)^{2} \hat{\mu}_{\pi}^{2}+\frac{4}{3 m_{b}^{3}}\left(m_{b}-v \cdot q\right)(v \cdot q)\left(2 \hat{\rho}_{D}^{3}+\hat{\rho}_{L S}^{3}\right),  \tag{3.14}\\
& \tilde{S}_{3}=\frac{4}{3 m_{b}^{2}}(v \cdot q)^{2} \hat{\mu}_{\pi}^{2}+\frac{8}{3 m_{b}^{3}}\left(m_{b}-v \cdot q\right)(v \cdot q) \hat{\rho}_{D}^{3},  \tag{3.15}\\
& \tilde{T}_{4}=\frac{8}{3 m_{b}^{3}}\left(m_{b}-v \cdot q\right)(v \cdot q)^{2} \hat{\rho}_{D}^{3},  \tag{3.16}\\
& \tilde{S}_{4}=\frac{8}{3 m_{b}^{3}}\left(m_{b}-v \cdot q\right)(v \cdot q)^{2} \hat{\rho}_{D}^{3} . \tag{3.17}
\end{align*}
$$

Using these results, from eq. (3.4) the distribution $\frac{d^{2} \Gamma}{d y d \cos \theta_{P}}$ in the photon energy $y=$ $2 E_{\gamma} / m_{b}$ and $\cos \theta_{P}$ can be computed, $\theta_{P}$ being the angle between the photon momentum
$\vec{q}$ and the $H_{b}$ spin vector $\vec{s}$ in the $H_{b}$ rest frame. Upon integration, the photon energy spectrum, the $\cos \theta_{P}$ distribution and the decay width are obtained.

In the $H_{b}$ rest frame $v \cdot q=E_{\gamma}=\frac{m_{b}}{2} y$, and the distribution comprises two terms:

$$
\begin{equation*}
\frac{d^{2} \Gamma}{d y d \cos \theta_{P}}=\tilde{\Gamma}_{1}+\cos \theta_{P} \tilde{\Gamma}_{2} \tag{3.18}
\end{equation*}
$$

Integrating (3.18) on $\cos \theta_{P}$ one has $\tilde{\Gamma}_{1}=\frac{1}{2} \frac{d \Gamma}{d y}$ and the photon energy spectrum

$$
\begin{align*}
\frac{1}{\Gamma_{0}} \frac{d \Gamma}{d y}= & {\left[1-\frac{\hat{\mu}_{\pi}^{2}}{2 m_{b}^{2}}-\frac{\hat{\mu}_{G}^{2}}{2 m_{b}^{2}} \frac{3+5 z}{1-z}-\frac{10 \hat{\rho}_{D}^{3}}{3 m_{b}^{3}} \frac{1+z}{1-z}\right] \delta(1-z-y) } \\
& +\left[\frac{\hat{\mu}_{\pi}^{2}}{2 m_{b}^{2}}(1-z)-\frac{\hat{\mu}_{G}^{2}}{6 m_{b}^{2}}(3+5 z)-\frac{4 \hat{\rho}_{D}^{3}}{3 m_{b}^{3}}(1+2 z)+\frac{2 \hat{\rho}_{L S}^{3}}{3 m_{b}^{3}}(1+z)\right] \delta^{\prime}(1-z-y) \\
& +\left[\frac{\hat{\mu}_{\pi}^{2}}{6 m_{b}^{2}}(1-z)^{2}-\frac{\hat{\rho}_{D}^{3}}{3 m_{b}^{3}}(1-z)(1+2 z)+\frac{\hat{\rho}_{L S}^{3}}{6 m_{b}^{3}}\left(1-z^{2}\right)\right] \delta^{\prime \prime}(1-z-y)  \tag{3.19}\\
& -\frac{\hat{\rho}_{D}^{3}}{18 m_{b}^{3}}(1-z)^{2}(1+z) \delta^{\prime \prime \prime}(1-z-y)
\end{align*}
$$

where $z=\frac{m_{s}^{2}}{m_{b}^{2}}$ and

$$
\begin{equation*}
\Gamma_{0}=\frac{\alpha G_{F}^{2}\left|\lambda_{t}\right|^{2}}{32 \pi^{4}} m_{b}^{5}(1-z)^{3}\left[\left|C_{+}^{\mathrm{eff}}\right|^{2}+\left|C_{+}^{\prime \mathrm{eff}}\right|^{2}\right] \tag{3.20}
\end{equation*}
$$

with

$$
C_{+}^{\mathrm{eff}}=C_{7}^{\mathrm{eff}}+\sqrt{z} C_{7}^{\prime \mathrm{eff}}, \quad C_{+}^{\prime \mathrm{eff}}=\sqrt{z} C_{7}^{\mathrm{eff}}+C_{7}^{\prime \mathrm{eff}}
$$

$\tilde{\Gamma}_{2}$ is given by:

$$
\begin{align*}
\frac{2}{\Gamma_{0}} \tilde{\Gamma}_{2}=- & \frac{\left|C_{+}^{\text {eff }}\right|^{2}-\left|C_{+}^{\prime \text { eff }}\right|^{2}}{\left|C_{+}^{\text {eff }}\right|^{2}+\left|C_{+}^{\prime \text { eff }}\right|^{2}} \\
& \times\left\{\left[1-\frac{13 \hat{\mu}_{\pi}^{2}}{12 m_{b}^{2}}-\frac{3 \hat{\mu}_{G}^{2}}{4 m_{b}^{2}} \frac{5+3 z}{1-z}-\frac{\hat{\rho}_{D}^{3}}{6 m_{b}^{3}} \frac{31+9 z}{1-z}\right] \delta(1-z-y)\right. \\
& +\left[\frac{\hat{\mu}_{\pi}^{2}}{2 m_{b}^{2}}(1-z)-\frac{\hat{\mu}_{G}^{2}}{2 m_{b}^{2}}(3+z)-\frac{2 \hat{\rho}_{D}^{3}}{m_{b}^{3}}(1+z)\right] \delta^{\prime}(1-z-y)  \tag{3.21}\\
& +\left[\frac{\hat{\mu}_{\pi}^{2}}{6 m_{b}^{2}}(1-z)^{2}-\frac{\hat{\rho}_{D}^{3}}{3 m_{b}^{3}}(1-z)(1+2 z)\right] \delta^{\prime \prime}(1-z-y) \\
& \left.-\frac{\hat{\rho}_{D}^{3}}{18 m_{b}^{3}}(1-z)^{2}(1+z) \delta^{\prime \prime \prime}(1-z-y)\right\} .
\end{align*}
$$

The $\cos \theta_{P}$ distribution also comprises two terms:

$$
\begin{equation*}
\frac{d \Gamma\left(H_{b} \rightarrow X_{s} \gamma\right)}{d \cos \theta_{P}}=A+B \cos \theta_{P} \tag{3.22}
\end{equation*}
$$

with

$$
\begin{align*}
& A=\frac{1}{2} \Gamma\left(H_{b} \rightarrow X_{s} \gamma\right)  \tag{3.23}\\
& B=-\frac{\Gamma_{0}}{2} \frac{\left|C_{+}^{\mathrm{eff}}\right|^{2}-\left|C_{+}^{\prime \text { eff }}\right|^{2}}{\left|C_{+}^{\mathrm{eff}}\right|^{2}+\left|C_{+}^{\prime \text { eff }}\right|^{2}}\left[1-\frac{13 \hat{\mu}_{\pi}^{2}}{12 m_{b}^{2}}-\frac{3 \hat{\mu}_{G}^{2}}{4 m_{b}^{2}} \frac{5+3 z}{1-z}-\frac{\hat{\rho}_{D}^{3}}{6 m_{b}^{3}} \frac{31+9 z}{1-z}\right] \tag{3.24}
\end{align*}
$$

The inclusive $H_{b} \rightarrow X_{s} \gamma$ decay width is given by

$$
\begin{equation*}
\Gamma\left(H_{b} \rightarrow X_{s} \gamma\right)=\Gamma_{0}\left[1-\frac{\hat{\mu}_{\pi}^{2}}{2 m_{b}^{2}}-\frac{\hat{\mu}_{G}^{2}}{2 m_{b}^{2}} \frac{3+5 z}{1-z}-\frac{10 \hat{\rho}_{D}^{3}}{3 m_{b}^{3}} \frac{1+z}{1-z}\right] \tag{3.25}
\end{equation*}
$$

The SM result is recovered for $C_{7}^{\prime} \rightarrow 0$. At $\mathcal{O}\left(1 / m_{b}^{2}\right)$ and for $m_{s}=0$ eqs. (3.25) and (3.19) agree with the SM expressions obtained in $[47,48]$. At $\mathcal{O}\left(1 / m_{b}^{3}\right)$ they agree with the expressions in [49] substituting

$$
\begin{align*}
\hat{\mu}_{\pi}^{2} & \rightarrow-\left(\lambda_{1}+\frac{\mathcal{T}_{1}+3 \mathcal{T}_{3}}{m_{b}}\right) \\
\hat{\mu}_{G}^{2} & \rightarrow 3\left(\lambda_{2}+\frac{\mathcal{T}_{3}+3 \mathcal{T}_{4}}{m_{b}}\right)-\frac{\rho_{1}+3 \rho_{2}}{m_{b}}  \tag{3.26}\\
\hat{\rho}_{D}^{3} & \rightarrow \rho_{1} \\
\hat{\rho}_{L S}^{3} & \rightarrow 3 \rho_{2}
\end{align*}
$$

For $m_{s}=0$ the $\mathcal{O}\left(1 / m_{b}^{4}\right)$ corrections to the decay width have been computed in [50]. For $\Lambda_{b}$, the distribution (3.22) has been computed at $\mathcal{O}\left(1 / m_{b}^{2}\right)$ [51], we agree with such a result.

## 4 Photon polarization

The photon polarization in the $b \rightarrow s \gamma$ transition can be measured in radiative beauty baryon decays. In SM the photon polarization asymmetry $A_{P}$, measuring the relative abundance of left-handed with respect to right-handed photons, is predicted $A_{P} \simeq-1$. Deviations from this result would hint physics beyond the Standard Model. The photon polarization has been studied for the exclusive [52-57] and inclusive beauty baryon decays [51], the experimental feasibility at LHCb has also been scrutinized [58].

The correlation between the photon polarization and the initial $b$-baryon polarization is particularly relevant, and it can be obtained considering the $A_{P}$ dependence on $\cos \theta_{P}$. We refer to $\Lambda_{b}$, however our results can be applied also to other beauty baryons. Preliminary experimental analyses for $\Sigma_{b}$ are reported in [20].

To present the results for the photon polarizations $\epsilon_{ \pm}$we write the double differential width as

$$
\begin{equation*}
\frac{d^{2} \Gamma_{ \pm}}{d y d \cos \theta_{P}}=\tilde{\Gamma}_{ \pm, 1}+\tilde{\Gamma}_{ \pm, 2} \cos \theta_{P} \tag{4.1}
\end{equation*}
$$

where $\tilde{\Gamma}_{ \pm, 1}=\frac{1}{2} \frac{d \Gamma_{ \pm}}{d y}$. The energy spectrum for definite photon helicities has the expression

$$
\begin{align*}
\frac{d \Gamma_{+}}{d y} & =R_{+, 1} \frac{d \Gamma}{d y}=\frac{\left|C_{+}^{\prime \text { eff }}\right|^{2}}{\left|C_{+}^{\mathrm{eff}}\right|^{2}+\left|C_{+}^{\prime \text { eff }}\right|^{2}} \frac{d \Gamma}{d y}  \tag{4.2}\\
\frac{d \Gamma_{-}}{d y} & =R_{-, 1} \frac{d \Gamma}{d y}=\frac{\left|C_{+}^{\mathrm{eff}}\right|^{2}}{\left|C_{+}^{\mathrm{eff}}\right|^{2}+\left|C_{+}^{\prime \text { eff }}\right|^{2}} \frac{d \Gamma}{d y} \tag{4.3}
\end{align*}
$$

$\tilde{\Gamma}_{ \pm, 2}$ are given by

$$
\begin{align*}
& \tilde{\Gamma}_{+, 2}=R_{+, 2} \tilde{\Gamma}_{2}=-\frac{\left|C_{+}^{\prime \text { eff }}\right|^{2}}{\left|C_{+}^{\text {eff }}\right|^{2}-\left|C_{+}^{\prime \text { eff }}\right|^{2}} \tilde{\Gamma}_{2}  \tag{4.4}\\
& \tilde{\Gamma}_{-, 2}=R_{-, 2} \tilde{\Gamma}_{2}=\frac{\left|C_{+}^{\mathrm{eff}}\right|^{2}}{\left|C_{+}^{\text {eff }}\right|^{2}-\left|C_{+}^{\prime \text { eff }}\right|^{2}} \tilde{\Gamma}_{2} \tag{4.5}
\end{align*}
$$

In SM (for $C_{7}^{\prime \text { eff }} \rightarrow 0$ ) we have:

$$
\begin{align*}
\frac{d \Gamma_{+}^{\mathrm{SM}}}{d y} & =\frac{z}{1+z} \frac{d \Gamma^{\mathrm{SM}}}{d y}  \tag{4.6}\\
\frac{d \Gamma_{-}^{\mathrm{SM}}}{d y} & =\frac{1}{1+z} \frac{d \Gamma^{\mathrm{SM}}}{d y} \tag{4.7}
\end{align*}
$$

and

$$
\begin{align*}
& \tilde{\Gamma}_{+, 2}^{S M}=-\frac{z}{1-z} \tilde{\Gamma}_{2}^{S M}  \tag{4.8}\\
& \tilde{\Gamma}_{-, 2}^{S M}=\frac{1}{1-z} \tilde{\Gamma}_{2}^{S M} \tag{4.9}
\end{align*}
$$

In SM for $m_{s}=0$ only the polarization $\epsilon_{-}$contributes. Indeed, the operator $O_{7}$ produces a right-handed $b$ quark; the massless $s$ quark has fixed helicity and, due to the angular momentum conservation, the $b$ and $s$ quarks have spins aligned to the photon spin, in the opposite direction.

For polarized photons the $\cos \theta_{P}$ distributions have the form

$$
\begin{equation*}
\frac{d \Gamma_{ \pm}\left(H_{b} \rightarrow X_{s} \gamma\right)}{d \cos \theta_{P}}=A_{ \pm}+B_{ \pm} \cos \theta_{P} \tag{4.10}
\end{equation*}
$$

where

$$
\begin{align*}
& A_{ \pm}=\frac{\Gamma_{ \pm}\left(H_{b} \rightarrow X_{s} \gamma\right)}{2}=R_{ \pm, 1} A \\
& B_{ \pm}=R_{ \pm, 2} B \tag{4.11}
\end{align*}
$$

and $R_{ \pm, 1(2)}$ defined in eqs. (4.2)-(4.5). The decay widths to polarized photons are given by

$$
\begin{equation*}
\Gamma_{ \pm}\left(H_{b} \rightarrow X_{s} \gamma\right)=R_{ \pm, 1} \Gamma\left(H_{b} \rightarrow X_{s} \gamma\right) \tag{4.12}
\end{equation*}
$$

The photon polarization asymmetry is defined as

$$
\begin{equation*}
A_{P}\left(\cos \theta_{P}\right)=\frac{\frac{d \Gamma_{+}}{d \cos \theta_{P}}-\frac{d \Gamma_{-}}{d \cos _{-} \theta_{P}}}{\frac{d \Gamma_{+}}{d \cos \theta_{P}}+\frac{d \Gamma_{-}}{d \cos \theta_{P}}} \tag{4.13}
\end{equation*}
$$

In SM the photon polarization asymmetry is $A_{P}\left(\cos \theta_{P}\right) \simeq-1$ for almost all $\cos \theta_{P}$, it increases only for $\cos \theta_{P} \rightarrow 1$, see figure 1 . Physics beyond SM can produce a sizable effect. For a quantitative insight on the possible deviation from SM , we consider ranges for $C_{7}^{\mathrm{NP}}=$ $C_{7}^{\text {eff }}-\left(C_{7}^{\text {eff }}\right)^{\mathrm{SM}}$ and $C_{7}^{\prime \text { eff }}$, assuming that both coefficients are real, exploiting the results of a global fit of the $b \rightarrow s$ transitions [59]. Using $m_{b}=4.62 \mathrm{GeV}, m_{s}=0.150 \mathrm{GeV}, \hat{\mu}_{\pi}^{2}\left(\Lambda_{b}\right)=$ $0.5 \mathrm{GeV}^{2}, \hat{\rho}_{D}^{3}\left(\Lambda_{b}\right)=0.17 \mathrm{GeV}^{3}, \hat{\mu}_{G}^{2}\left(\Lambda_{b}\right)=\hat{\rho}_{L S}^{3}\left(\Lambda_{b}\right)=0$, and varying $C_{7}^{\prime \text { eff }} / C_{7}^{\text {eff }} \in[-0.3,0.3]$ we obtain the asymmetry shown in figure 2 . In the same figure we plot $A_{P}\left(\cos \theta_{P}\right)$ versus $C_{7}^{\prime \text { eff }} / C_{7}^{\text {eff }}$ for selected values of $\cos \theta_{P}$. A deviation of the polarization asymmetry from the SM value can be obtained, with the largest effect for $\cos \theta_{P} \simeq 1$.


Figure 1. Photon polarization asymmetry eq. (4.13) versus $\cos \theta_{P}$ in SM .


Figure 2. Photon polarization asymmetry eq. (4.13) varying $\cos \theta_{P}$ and $C_{7}^{\prime \text { eff }} / C_{7}^{\text {eff }}$ (top panel) and projected for several values of $\cos \theta_{P}$ (bottom panel).

## 5 Treatment of the singular terms

The spectrum obtained by the short distance OPE does not account for the Fermi motion of the $b$ quark due to soft interactions with the light degrees of freedom in the hadron. For decays of a beauty hadron $H_{b}$ to light partons the relevant scales are $m_{b}, \sqrt{m_{b} \Lambda_{\mathrm{QCD}}}$ and $\Lambda_{\mathrm{QCD}}$ [60]. The "shape function" region is the kinematic region where the hadronic invariant mass $p_{X}^{2} \sim \mathcal{O}\left(m_{b} \Lambda_{\mathrm{QCD}}\right)$. Suitable effective field theory must be used in the
various regions: QCD is first matched to the soft collinear effective theory (SCET) [61-64], followed by matching to HQET [60, 65]. In the shape function region the short distance OPE is replaced by a twist expansion, with the infinite set of power corrections resummed into nonperturbative functions. ${ }^{3}$

At leading order in the HQE there is a single shape function defined in HQET. Considering the process $H_{b} \rightarrow X_{s} \gamma$, one defines the spectral function $S_{s}(y)$ [28]

$$
\begin{equation*}
S_{s}(y)=\left\langle\delta\left[1-y-z+\frac{2}{m_{b}}(v-\hat{q}) \cdot i D\right]\right\rangle, \tag{5.1}
\end{equation*}
$$

with $\hat{q}=q / m_{b}$. For a generic operator $\mathcal{O}$ the matrix element in (5.1) is defined as

$$
\begin{equation*}
\langle\mathcal{O}\rangle=\frac{\left\langle H_{b}(v)\right| \bar{h}_{v} \mathcal{O} h_{v}\left|H_{b}(v)\right\rangle}{\left\langle H_{b}(v)\right| \bar{h}_{v} h_{v}\left|H_{b}(v)\right\rangle}, \tag{5.2}
\end{equation*}
$$

with $h_{v}$ the HQET field with velocity $v$. Introducing the vector $n_{\mu}+\delta n_{\mu}=\left.2(v-\hat{q})_{\mu}\right|_{y=1-z}$, with $n^{2}=0, v \cdot n=1$, and $n \cdot \delta n \sim O\left(\Lambda_{\mathrm{QCD}} / m_{b}\right)$ in the shape function region [67], and defining $k_{+}=n \cdot k$ for a vector $k$, we have

$$
\begin{equation*}
S_{s}(y)=\left\langle\delta\left[1-y-z+\frac{i D_{+}}{m_{b}}\right]\right\rangle . \tag{5.3}
\end{equation*}
$$

$S_{s}(y)$ can be expressed as

$$
\begin{equation*}
S_{s}(y)=\int d k_{+} \delta\left(1-y-z+\frac{k_{+}}{m_{b}}\right)\left[f\left(k_{+}\right)+\mathcal{O}\left(m_{b}^{-1}\right)\right] . \tag{5.4}
\end{equation*}
$$

The function

$$
\begin{equation*}
f\left(k_{+}\right)=\left\langle\delta\left(i D_{+}-k_{+}\right)\right\rangle \tag{5.5}
\end{equation*}
$$

is the leading shape function. ${ }^{4}$ The photon energy spectrum is given by the convolution [28]

$$
\begin{equation*}
\frac{d \Gamma}{d y}=\int d k_{+} f\left(k_{+}\right) \frac{d \Gamma^{*}}{d y} . \tag{5.6}
\end{equation*}
$$

In $\frac{d \Gamma^{*}}{d y}$ the $b$ quark mass $m_{b}$ is replaced by $m_{b}^{*}=m_{b}+k_{+}$, an exact substitution at tree level. For $k_{+}$in the range $k_{+} \in\left[-m_{b}, m_{H_{b}}-m_{b}\right]$, replacing $m_{b} \rightarrow m_{b}^{*}$ in the variable $y$, we find (for $m_{s}=0$ to simplify the discussion) that $y \rightarrow \frac{2 E_{\gamma}}{\left(m_{b}+k_{+}\right)}$. Therefore, for $k_{+}^{\max }=m_{H_{b}}-m_{b}$ the maximum photon energy is $E_{\gamma}=\frac{m_{H_{b}}}{2}$, the physical endpoint.

The shape function provides an interpretation of the singular terms in the photon energy spectrum obtained in the previous sections. The distribution in (3.19) can be written as

$$
\begin{equation*}
\frac{1}{\Gamma} \frac{d \Gamma}{d y}=\sum_{n=0}^{3} \frac{M_{n}}{n!} \delta^{(n)}(1-z-y) \tag{5.7}
\end{equation*}
$$

[^2]with $\Gamma$ in eq. (3.25) and $M_{0, \ldots 3}$ given by
\[

$$
\begin{align*}
& M_{0}=1, \\
& M_{1}=\frac{\hat{\mu}_{\pi}^{2}}{2 m_{b}^{2}}(1-z)-\frac{\hat{\mu}_{G}^{2}}{6 m_{b}^{2}}(3+5 z)-\frac{4 \hat{\rho}_{D}^{3}}{3 m_{b}^{3}}(1+2 z)+\frac{2 \hat{\rho}_{L S}^{3}}{3 m_{b}^{3}}(1+z), \\
& M_{2}=\frac{\hat{\mu}_{\pi}^{2}}{3 m_{b}^{2}}(1-z)^{2}-\frac{2 \hat{\rho}_{D}^{3}}{3 m_{b}^{3}}(1-z)(1+2 z)+\frac{\hat{\rho}_{L S}^{3}}{3 m_{b}^{3}}\left(1-z^{2}\right),  \tag{5.8}\\
& M_{3}=-\frac{\hat{\rho}_{D}^{3}}{3 m_{b}^{3}}(1-z)^{2}(1+z) .
\end{align*}
$$
\]

At $\mathcal{O}\left(1 / m_{b}^{2}\right)$ eqs. (5.8) agree with the expressions obtained in [28]. They can be considered as the first few terms of the infinite sum [27-29, 46, 67]

$$
\begin{equation*}
S_{s}(y)=\sum_{n=0}^{\infty} \frac{M_{n}}{n!} \delta^{(n)}(1-z-y) . \tag{5.9}
\end{equation*}
$$

As pointed out in [28], a feature of eqs. (5.8) is that each moment $M_{n}$ has an expansion in powers of $1 / m_{b}$ starting at the same order of the moment,

$$
\begin{equation*}
M_{n}=\sum_{k=n}^{\infty} \frac{M_{n, k}}{m_{b}^{k}} . \tag{5.10}
\end{equation*}
$$

Analogously, the double differential distribution (3.18) can be written as

$$
\begin{equation*}
\frac{d^{2} \Gamma}{d y d \cos \theta_{P}}=A\left[\sum_{n=0}^{3} \frac{M_{n}}{n!} \delta^{(n)}(1-z-y)\right]+B\left[\sum_{n=0}^{3} \frac{M_{n}^{\theta_{P}}}{n!} \delta^{(n)}(1-z-y)\right] \cos \theta_{P} \tag{5.11}
\end{equation*}
$$

with $A, B$ in (3.23), (3.24). The moments $M_{n}$ in eq. (3.19) can be considered as resulting from the expansion of the function (5.9), and the terms

$$
\begin{align*}
& M_{0}^{\theta_{P}}=1 \\
& M_{1}^{\theta_{P}}=\frac{\hat{\mu}_{\pi}^{2}}{2 m_{b}^{2}}(1-z)-\frac{\hat{\mu}_{G}^{2}}{2 m_{b}^{2}}(3+z)-\frac{2 \hat{\rho}_{D}^{3}}{m_{b}^{3}}(1+z) \\
& M_{2}^{\theta_{P}}=\frac{\hat{\mu}_{\pi}^{2}}{3 m_{b}^{2}}(1-z)^{2}-\frac{2 \hat{\rho}_{D}^{3}}{3 m_{b}^{3}}(1-z)(1+2 z)  \tag{5.12}\\
& M_{3}^{\theta_{P}}=-\frac{\hat{\rho}_{D}^{3}}{3 m_{b}^{3}}(1-z)^{2}(1+z)
\end{align*}
$$

as deriving from the function

$$
\begin{equation*}
S_{s}^{\theta_{P}}(y)=\sum_{n=0}^{\infty} \frac{M_{n}^{\theta_{P}}}{n!} \delta^{(n)}(1-z-y) . \tag{5.13}
\end{equation*}
$$

The factors $M_{n}$ in (5.9) are related to the moments of the photon energy spectrum

$$
\begin{equation*}
\left\langle y^{k}\right\rangle=\frac{1}{\Gamma} \int_{0}^{1-z} d y y^{k} \frac{d \Gamma}{d y} \tag{5.14}
\end{equation*}
$$

|  | LO | $O\left(1 / m_{b}^{2}\right)$ | $O\left(1 / m_{b}^{3}\right)$ |
| :---: | :---: | :---: | :---: |
| $\langle y\rangle$ | 0.999 | 1.011 | 1.008 |
| $\left\langle y^{2}\right\rangle$ | 0.998 | 1.029 | 1.023 |
| $\sigma_{y}^{2}$ | 0 | 0.008 | 0.007 |

Table 1. First moments of the photon energy spectrum at LO, $O\left(1 / m_{b}^{2}\right), O\left(1 / m_{b}^{3}\right)$.

Indeed, using (5.7) we have:

$$
\begin{equation*}
\left\langle y^{k}\right\rangle=\sum_{n=0}^{\infty} \frac{M_{n}}{n!} \int_{0}^{1-z} d y y^{k} \delta^{(n)}(1-z-y)=\sum_{j=0}^{k}\binom{k}{j}(1-z)^{k-j} M_{j} \tag{5.15}
\end{equation*}
$$

and

$$
\begin{align*}
\langle y\rangle & =(1-z)+M_{1},  \tag{5.16}\\
\left\langle y^{2}\right\rangle & =(1-z)^{2}+2(1-z) M_{1}+M_{2},  \tag{5.17}\\
\sigma_{y}^{2} & =\left\langle y^{2}\right\rangle-\langle y\rangle^{2}=M_{2}-M_{1}^{2} . \tag{5.18}
\end{align*}
$$

Such results imply that

$$
\begin{align*}
\langle y\rangle & =(1-z)\left[1+\frac{\hat{\mu}_{\pi}^{2}}{2 m_{b}^{2}}-\frac{\hat{\mu}_{G}^{2}}{6 m_{b}^{2}} \frac{3+5 z}{1-z}-\frac{4 \hat{\rho}_{D}^{3}}{3 m_{b}^{3}} \frac{1+2 z}{1-z}+\frac{2 \hat{\rho}_{L S}^{3}}{3 m_{b}^{3}} \frac{1+z}{1-z}\right],  \tag{5.19}\\
\left\langle y^{2}\right\rangle & =(1-z)^{2}\left[1+\frac{4 \hat{\mu}_{\pi}^{2}}{3 m_{b}^{2}}-\frac{\hat{\mu}_{G}^{2}}{3 m_{b}^{2}} \frac{3+5 z}{1-z}-\frac{10 \hat{\rho}_{D}^{3}}{3 m_{b}^{3}} \frac{1+2 z}{1-z}+\frac{5 \hat{\rho}_{L S}^{3}}{3 m_{b}^{3}} \frac{1+z}{1-z}\right],  \tag{5.20}\\
\sigma_{y}^{2} & =(1-z)^{2}\left[\frac{\hat{\mu}_{\pi}^{2}}{3 m_{b}^{2}}-\frac{2 \hat{\rho}_{D}^{3}}{3 m_{b}^{3}} \frac{1+2 z}{1-z}+\frac{\hat{\rho}_{L S}^{3}}{3 m_{b}^{3}} \frac{1+z}{1-z}\right] . \tag{5.21}
\end{align*}
$$

After the substitutions in (3.26), the above expressions agree with those given in [68] for $z=0$.

Table 1 contains numerical results for the first moments, increasing the order in $1 / m_{b}$ and using the parameters in section 4 . The moments of the measured photon energy spectrum can be used to determine the HQET parameters, as for $B$ mesons. Baryons have the advantage that the $\cos \theta_{P}$ distribution (3.22) can also be exploited.

The Fermi motion of the heavy quark in the hadron has the effect of smearing the spectrum. Indeed, at the various orders in the $1 / m_{b}$ expansion, the photon energy spectrum obtained by the local OPE corresponds to a monochromatic line. At the leading order the line is placed at $y=2 E_{\gamma} / m_{b}=1-z$, the next terms correspond to a displacement of this position. The convolution with the shape function $f\left(k_{+}\right)$in (5.6) produces the smearing. The shape function is a nonperturbative quantity, it must be determined by methods as lattice QCD or QCD sum rules starting from the moments, or it must be modelized to reproduce the experimental photon spectrum [26, 69-72]. In the latter case, the uncertainty connected to the functional dependence is usually estimated using different model functions, or varying the model parameters.

Instead of parametrizing the shape function, we proceed considering that, if an infinite number of terms is included, the sum eq. (5.7) gives the spectral function $S_{s}(y)$ in (5.9). In the sum the first term corresponds to a monocromatic line at the zero of the $\delta$-function, with $\langle y\rangle=(1-z)$ and $\sigma_{y}^{2}=0$, the leading order results in eqs. (5.16)-(5.18). We observe that the Dirac delta can be represented as

$$
\begin{equation*}
\delta(1-z-y)=\lim _{\sigma_{y} \rightarrow 0} \frac{1}{\sqrt{2 \pi} \sigma_{y}} e^{-\frac{(b-y)^{2}}{2 \sigma_{y}^{2}}} \tag{5.22}
\end{equation*}
$$

with $b=(1-z)=\langle y\rangle_{L O}$. We can fix $\sigma_{y}^{2}$ at each order in $1 / m_{b}$, starting from $1 / m_{b}^{2}$. For $m_{b} \rightarrow \infty$ the limit $\sigma_{y} \rightarrow 0$ reproduces the partonic result.

We represent the spectral function $S_{s}(y)$ by the substitution

$$
\begin{equation*}
S_{s}(y)=\sum_{n=0}^{\infty} \frac{M_{n}}{n!} \delta^{(n)}(1-z-y) \rightarrow S_{s}(y)=\sum_{n=0}^{\infty} \frac{M_{n}}{n!}(-1)^{n} \frac{d^{n}}{d y^{n}} \frac{1}{\sqrt{2 \pi} \sigma_{y}} e^{-\frac{(b-y)^{2}}{2 \sigma_{y}^{2}}} \tag{5.23}
\end{equation*}
$$

Using the representation of the Hermite polynomials

$$
\begin{equation*}
H_{n}(x)=(-1)^{n} e^{x^{2}} \frac{d^{n}}{d x^{n}} e^{-x^{2}} \tag{5.24}
\end{equation*}
$$

the substitution gives

$$
\begin{equation*}
S_{s}(y)=\frac{1}{\sqrt{2 \pi} \sigma_{y}} \sum_{n=0}^{\infty} \frac{M_{n}}{n!}\left(-\frac{1}{\sqrt{2} \sigma_{y}}\right)^{n} e^{-\frac{(b-y)^{2}}{2 \sigma_{y}^{2}}} H_{n}\left(\frac{b-y}{\sqrt{2} \sigma_{y}}\right) \tag{5.25}
\end{equation*}
$$

Notice that, denoting by $\left\langle y^{k}\right\rangle_{\mathcal{N}}$ the moments computed using this expression for the spectral function $S_{s}(y)$,

$$
\begin{equation*}
\left\langle y^{k}\right\rangle_{\mathcal{N}}=\int_{0}^{y_{\max }} d y y^{k} S_{s}(y) \tag{5.26}
\end{equation*}
$$

in the limit $\sigma_{y} \rightarrow 0$ one obtains

$$
\begin{equation*}
\lim _{\sigma_{y} \rightarrow 0}\left\langle y^{k}\right\rangle_{\mathcal{N}}=\left\langle y^{k}\right\rangle \tag{5.27}
\end{equation*}
$$

with $\left\langle y^{k}\right\rangle$ in (5.15). This is shown in appendix C.
The ansatz eq. (5.25) has many advantages with respect to other representations of the shape function or of the spectral function, based on a choice of a functional representation able to reproduce the photon spectrum, with parameters set by the first computed moments $M_{n}$. Indeed, such representations generally do not guarantee that higher moments are reproduced. Moreover, in such models the moments $M_{n}$ generally increase with the order $n$ [73]. To cure such features, in [74] the shape function is expressed using a complete set of orthonormal functions. In particular, the normalized Legendre polynomials are considered in the range $[-1,+1]$, and a function mapping the range $[-1,1]$ into $[0,+\infty)$ is chosen to represent the shape function $f(\omega)$ in the definition having support $\omega \in[0,+\infty)$.

Remarkably, with the ansatz in eq. (5.25), by construction $S_{s}(y)$ can include all moments $M_{n}$ once they are computed. Moreover, less singular terms in the expansion of
the moments are not discarded. Each $M_{n}$ starts at $\mathcal{O}\left(1 / m_{b}^{n}\right)$ and depends on the matrix elements of the HQET operators of increasing dimension,

$$
\begin{equation*}
\mathcal{M}_{\mu_{1} \ldots \mu_{n}}=\left\langle H_{b}(v, s)\right|\left(\bar{b}_{v}\right)_{a}\left(i D_{\mu_{1}}\right) \ldots\left(i D_{\mu_{n}}\right)\left(b_{v}\right)_{b}\left|H_{b}(v, s)\right\rangle \tag{5.28}
\end{equation*}
$$

of $\mathcal{O}\left(\Lambda_{\mathrm{QCD}}^{n}\right)$, maintaining a hierarchical ordering.
We point out the main differences with the method proposed in [74] and adopted in [30]. In that approaches it is assumed that the leading shape function is positive, hence it is expressed as the square of the sum of orthonormal functions, choosing in particular the normalized Legendre polynomials. The positivity assumption is not necessary in the ansatz (5.25). In such expression the expansion in Hermite polynomials is not arbitrary, since it comes from the replacement of the Dirac delta by the normal distribution with standard deviation $\sigma_{y}$ (5.22). This produces a different result from the single gaussian function in e.g. [67, 75]. Indeed, the derivatives of the exponential in eq. (5.23) produce coefficients with a non trivial dependence on $y$, which are resummed giving the Hermite polynomials in (5.25). This produces the asymmetry of the shape function with respect to the point $y=b$. While $b$ is fixed to the LO result for $\langle y\rangle, \sigma_{y}$ can be determined at an arbitrary order in the $1 / m_{b}$ expansion; eq. (5.21) satisfies by construction the condition $\lim _{m_{b} \rightarrow \infty} \sigma_{y}=0$, recovering the monochromatic spectrum in the limit. In (5.25) the Hermite polynomials are not weighted by new unkown coefficients to be fitted, but by the computed moments $M_{n}$.

For an analysis based on the ansatz (5.25), in figure 3 we show the spectral function obtained at LO, $\mathcal{O}\left(1 / m_{b}^{2}\right)$ and $\mathcal{O}\left(1 / m_{b}^{3}\right)$. In the same figure we plot the shape function obtained from (5.4) at $\mathcal{O}\left(1 / m_{b}^{2}\right)$ and $\mathcal{O}\left(1 / m_{b}^{3}\right)$. As a consequence of broadening the spectrum through the substitution in (5.23), there is a tail exceeding the physical endpoint $\bar{y}_{\text {max }} \simeq \frac{m_{\Lambda_{b}}}{m_{b}}(1-z)$, and a tail in the shape function exceeding $k_{+}^{\max }=m_{\Lambda_{b}}-m_{b}$. This is a spurious effect of the truncation. When higher orders in the HQE are included, the area below such tails approaches to zero. Indeed, denoting this area by $\Delta\left\langle y^{0}\right\rangle_{n}=\int_{\bar{y}_{\text {max }}}^{\infty} d y \frac{1}{\Gamma} d \Gamma$ computed at $\mathcal{O}\left(1 / m_{b}^{n}\right)$, we numerically find $\frac{\Delta\left\langle y^{0}\right\rangle_{3}}{\Delta\left\langle y^{0}\right\rangle_{2}} \simeq \mathcal{O}\left(\frac{1}{m_{b}}\right)$. Increasing the order in the HQ expansion $\Delta\left\langle y^{0}\right\rangle_{n}$ reduces to zero, so that the physical endpoint is reached.

## 6 Conclusions

The HQE has been exploited to compute the inclusive decay width induced by the $b \rightarrow s \gamma$ transition for a beauty baryon, in particular $\Lambda_{b}$. The differential width in the rescaled photon energy $y=\frac{2 E_{\gamma}}{m_{b}}$ and in $\cos \theta_{P}$ allows to construct new observables with respect to mesons. The calculation has been carried out at $\mathcal{O}\left(1 / m_{b}^{3}\right)$ for non-vanishing strange quark mass, using the baryon matrix elements determined in [41]. Physics beyond the Standard Model represented by the operator $O_{7}^{\prime}$ is found to affect the photon polarization asymmetry. For the singular terms appearing in the spectrum as $\delta$-distribution and its derivatives we have proposed a treatment that can be systematically improved including higher order terms in the expansion.

Progress in the next studies will be achieved considering the full Hamiltonian (2.1), the resolved photon contribution and the subleading shape functions for $b$-baryons. This is


Figure 3. Spectral function $S_{s}(y)$ (top panel) and shape function $f\left(k_{+}\right)$(bottom panel) obtained using the ansatz eq. (5.25) up to $n=3$.
an important step forward, in view of the wealth of new information which can be gained on SM and on the possible extensions analyzing the beauty baryon rare radiative decay modes together with the polarization effects.

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## A Derivation of the OPE

The OPE for the expression (3.5) can be constructed expressing the hadron momentum $p=m_{H} v$, with $v$ the four-velocity, in terms of $m_{b}$ and of a residual momentum $k: p=$ $m_{b} v+k$. The QCD $b$ quark field is rescaled

$$
\begin{equation*}
b(x)=e^{-i m_{b} v \cdot x} b_{v}(x), \tag{A.1}
\end{equation*}
$$

and the QCD field $b_{v}(x)$ satisfies the equation of motion

$$
\begin{equation*}
b_{v}(x)=\left(P_{+}+\frac{i \not D}{2 m_{b}}\right) b_{v}(x), \tag{A.2}
\end{equation*}
$$

with velocity projector $P_{+}=\frac{1+\not p}{2}$. Expressed in terms of $b_{v}(x)$ eq. (3.5) becomes:

$$
\begin{equation*}
T_{M N}^{i j}=i \int d^{4} x e^{i\left(m_{b} v-q\right) \cdot x}\left\langle H_{b}(v, s)\right| T\left[\hat{J}_{M}^{i \dagger}(x) \hat{J}_{N}^{j}(0)\right]\left|H_{b}(v, s)\right\rangle \tag{A.3}
\end{equation*}
$$

$\hat{J}^{i}$ contains the field $b_{v}$. The heavy quark expansion is obtained from

$$
\begin{equation*}
T_{M N}^{i j}=\left\langle H_{b}(v, s)\right| \bar{b}_{v}(0) \bar{\Gamma}_{M}^{i} S_{s}\left(p_{X}\right) \Gamma_{N}^{j} b_{v}(0)\left|H_{b}(v, s)\right\rangle, \tag{A.4}
\end{equation*}
$$

with $\bar{\Gamma}_{M}^{i}=\gamma^{0} \Gamma_{M}^{i \dagger} \gamma^{0}$ and $\Gamma_{M}^{7}=\sigma_{\mu \nu}\left(m_{b} P_{R}+m_{s} P_{L}\right), \Gamma_{M}^{7^{\prime}}=\sigma_{\mu \nu}\left(m_{b} P_{L}+m_{s} P_{R}\right) . S_{s}\left(p_{X}\right)$ is the $s$ quark propagator. Replacing $k \rightarrow i D$, with $D$ the QCD covariant derivative, the $s$ quark propagator can be expanded:

$$
\begin{equation*}
S_{s}\left(p_{X}\right)=S_{s}^{(0)}-S_{s}^{(0)}(i \not D) S_{s}^{(0)}+S_{s}^{(0)}(i \not D) S_{s}^{(0)}(i \not D) S_{s}^{(0)}+\ldots \tag{A.5}
\end{equation*}
$$

where $S_{s}^{(0)}=\frac{1}{m_{b} \not \subset-\phi-m_{s}}$. Defining $p_{s}=m_{b} v-q, \mathcal{P}=\not \phi_{s}+m_{s}$ and $\Delta_{0}=p_{s}^{2}-m_{s}^{2}$, the expansion at order $1 / m_{b}^{3}$ is given by:

$$
\begin{align*}
\frac{1}{\pi} \operatorname{Im} T_{M N}^{i j}= & \frac{1}{\pi} \operatorname{Im} \frac{1}{\Delta_{0}}\left\langle H_{b}(v, s)\right| \bar{b}_{v}\left[\bar{\Gamma}_{M}^{i} \mathcal{P} \Gamma_{N}^{j}\right] b_{v}\left|H_{b}(v, s)\right\rangle \\
& -\frac{1}{\pi} \operatorname{Im} \frac{1}{\Delta_{0}^{2}}\left\langle H_{b}(v, s)\right| \bar{b}_{v}\left[\bar{\Gamma}_{M}^{i} \mathcal{P} \gamma^{\mu_{1}} \mathcal{P} \Gamma_{N}^{j}\right]\left(i D_{\mu_{1}}\right) b_{v}\left|H_{b}(v, s)\right\rangle \\
& +\frac{1}{\pi} \operatorname{Im} \frac{1}{\Delta_{0}^{3}}\left\langle H_{b}(v, s)\right| \bar{b}_{v}\left[\bar{\Gamma}_{M}^{i} \mathcal{P} \gamma^{\mu_{1}} \mathcal{P} \gamma^{\mu_{2}} \mathcal{P} \Gamma_{N}^{j}\right]\left(i D_{\mu_{1}}\right)\left(i D_{\mu_{2}}\right) b_{v}\left|H_{b}(v, s)\right\rangle  \tag{A.6}\\
& -\frac{1}{\pi} \operatorname{Im} \frac{1}{\Delta_{0}^{4}}\left\langle H_{b}(v, s)\right| \bar{b}_{v}\left[\bar{\Gamma}_{M}^{i} \mathcal{P} \gamma^{\mu_{1}} \mathcal{P} \gamma^{\mu_{2}} \mathcal{P} \gamma^{\mu_{3}} \mathcal{P} \Gamma_{N}^{j}\right]\left(i D_{\mu_{1}}\right)\left(i D_{\mu_{2}}\right)\left(i D_{\mu_{3}}\right) b_{v}\left|H_{b}(v, s)\right\rangle .
\end{align*}
$$

This expression involves the $H_{b}$ matrix elements of QCD operators of increasing dimension,

$$
\begin{align*}
& \left\langle H_{b}(v, s)\right| \bar{b}_{v}\left[\bar{\Gamma}_{M}^{i} \mathcal{P} \gamma^{\mu_{1}} \ldots \gamma^{\mu_{n}} \mathcal{P} \Gamma_{N}^{j}\right]\left(i D_{\mu_{1}}\right) \ldots\left(i D_{\mu_{n}}\right) b_{v}\left|H_{b}(v, s)\right\rangle= \\
& \quad \operatorname{Tr}\left[\left(\bar{\Gamma}_{M}^{i} \mathcal{P} \gamma^{\mu_{1}} \ldots \gamma^{\mu_{n}} \mathcal{P} \Gamma_{N}^{j}\right)_{b a}\left\langle H_{b}(v, s)\right|\left(\bar{b}_{v}\right)_{a}\left(i D_{\mu_{1}}\right) \ldots\left(i D_{\mu_{n}}\right)\left(b_{v}\right)_{b}\left|H_{b}(v, s)\right\rangle\right] \tag{A.7}
\end{align*}
$$

with $a, b$ Dirac indices. The hadronic matrix elements

$$
\begin{equation*}
\left(\mathcal{M}_{\mu_{1} \ldots \mu_{n}}\right)_{a b}=\left\langle H_{b}(v, s)\right|\left(\bar{b}_{v}\right)_{a}\left(i D_{\mu_{1}}\right) \ldots\left(i D_{\mu_{n}}\right)\left(b_{v}\right)_{b}\left|H_{b}(v, s)\right\rangle \tag{A.8}
\end{equation*}
$$

can be expressed in terms of a set of nonperturbative parameters, the number of which increases with the operator dimension. At $\mathcal{O}\left(1 / m_{b}^{3}\right)$ the following matrix elements are required:

$$
\begin{align*}
\left\langle H_{b}(v, s)\right| \bar{b}_{v}(i D)^{2} b_{v}\left|H_{b}(v, s)\right\rangle & =-2 m_{H} \hat{\mu}_{\pi}^{2}  \tag{A.9}\\
\left\langle H_{b}(v, s)\right| \bar{b}_{v}\left(i D_{\mu}\right)\left(i D_{\nu}\right)\left(-i \sigma^{\mu \nu}\right) b_{v}\left|H_{b}(v, s)\right\rangle & =2 m_{H} \hat{\mu}_{G}^{2}  \tag{A.10}\\
\left\langle H_{b}(v, s)\right| \bar{b}_{v}\left(i D_{\mu}\right)(i v \cdot D)\left(i D^{\mu}\right) b_{v}\left|H_{b}(v, s)\right\rangle & =2 m_{H} \hat{\rho}_{D}^{3}  \tag{A.11}\\
\left\langle H_{b}(v, s)\right| \bar{b}_{v}\left(i D_{\mu}\right)(i v \cdot D)\left(i D_{\nu}\right)\left(-i \sigma^{\mu \nu}\right) b_{v}\left|H_{b}(v, s)\right\rangle & =2 m_{H} \hat{\rho}_{L S}^{3} . \tag{A.12}
\end{align*}
$$

A procedure to compute $\mathcal{M}_{\mu_{1} \ldots \mu_{n}}$ has been exploited for $B$ meson for $n=4$ [76] and $n=5$ [77], introducing additional parameters with respect to those in (A.9)-(A.12). For a heavy baryon, the dependence on the spin four-vector $s_{\mu}$ is specified in (A.8). The matrix elements in the expansion at $\mathcal{O}\left(1 / m_{b}^{3}\right)$ keeping the $s_{\mu}$ dependence have been parametrized in [41] and are used in the present study.

## B Factorization formula in the endpoint region and shape functions

The local OPE exploited in this paper holds in the kinematic region where the hadronic invariant mass $p_{X}^{2}$ is $\mathcal{O}\left(m_{b}^{2}\right)$. In the region where $p_{X}^{2} \sim \mathcal{O}\left(m_{b} \Lambda_{\mathrm{QCD}}\right)$, the endpoint or shape function region, the theoretical treatment is different according to which of the operators in the effective Hamiltonian one considers. In our analysis we have focused on the dipole operators $O_{7}^{(1)}$ since they are the only terms contributing to lowest order in QCD. In the SM the contribution of $O_{7}$ to the correlation function (3.5) at the endpoint obeys a factorization formula

$$
\begin{equation*}
d \Gamma^{77} \sim H \cdot J \otimes S+\frac{1}{m_{b}} \sum_{i} H \cdot J \otimes s_{i}+\frac{1}{m_{b}} \sum_{i} H \cdot J_{i} \otimes S+\mathcal{O}\left(\frac{\Lambda_{\mathrm{QCD}}^{2}}{m_{b}^{2}}\right) . \tag{B.1}
\end{equation*}
$$

$H$ denote hard functions and $J$ and $j_{i}$ are jet functions computed perturbatively, with $H$ of $\mathcal{O}(1) . S$ and $s_{i}$ are the shape functions, which are nonperturbative. The function $S$ entering in (B.1) at leading order is the shape function defined in HQET.

When the other operators in (3.5) are considered, a more involved factorization formula holds at $\mathcal{O}\left(1 / m_{b}\right)$ [31]. The most important operators are $O_{2}$ and $O_{8}$, and the pairing of the operators in (3.5) produces different effects. In particular, the resolved photon contribution (RPC) mentioned in the Introduction appears as single RPC from the pairing of $O_{8}$ and $O_{2}$ among themselves and with $O_{7}$. Double RPC arise from the pairing of $O_{8}$ and $O_{2}$ among themselves [31]. For such contributions the leading term is $\mathcal{O}\left(\alpha_{s}\right)$ : this justifies their neglect in the present analysis.

## C Moments of the photon energy spectrum

To obtain eq. (5.27) we use the representation of the Hermite polynomials

$$
\begin{equation*}
H_{n}(x)=n!\sum_{m=0}^{\left\lfloor\frac{n}{2}\right\rfloor} \frac{(-1)^{m}}{m!(n-2 m)!}(2 x)^{n-2 m} . \tag{C.1}
\end{equation*}
$$

The moments $\left\langle y^{k}\right\rangle_{\mathcal{N}}$ are given by:

$$
\begin{equation*}
\left\langle y^{k}\right\rangle_{\mathcal{N}}=\sum_{j=0}^{k}\binom{k}{j} b^{k-j} \sum_{n=0}^{\infty} M_{n}\left(-\sqrt{2} \sigma_{y}\right)^{j-n} \Phi_{j, n} \tag{C.2}
\end{equation*}
$$

where

$$
\begin{align*}
\Phi_{j, n}= & \frac{1}{\sqrt{\pi}} \sum_{m=0}^{\left\lfloor\frac{n}{2}\right\rfloor} \frac{(-1)^{m}}{m!(n-2 m)!} 2^{n-2 m} \\
& \times \frac{1}{2}\left[\gamma\left(\frac{j+n-2 m+1}{2}, x_{\max }^{2}\right)-(-1)^{j+n-2 m+1} \gamma\left(\frac{j+n-2 m+1}{2}, x_{\min }^{2}\right)\right] . \tag{C.3}
\end{align*}
$$

The parameters $b$ and $\sigma_{y}$ are defined in section 5, and $x_{\max (\min )}$ are $x_{\max }=\frac{b}{\sqrt{2} \sigma_{y}}>0$, $x_{\text {min }}=\frac{b-y_{\text {max }}}{\sqrt{2} \sigma_{y}}<0 . \gamma(n, z)$ is the lower incomplete Euler function with the condition
$\operatorname{Re}(j+n-2 m)>-1$ always satisfied in our case. $\Phi_{j, n}$ depends on $\sigma_{y}$ only through $x_{\max (\min )}$. Since $\lim _{\sigma_{y} \rightarrow 0} x_{\max (\min )}=+(-) \infty$, we have $\lim _{\sigma_{y} \rightarrow 0} \Phi_{j, n}=0$ for $n \neq j$. Consequently, for $\sigma_{y} \rightarrow 0$ we obtain that $\left\langle y^{k}\right\rangle_{\mathcal{N}}$ is given by the $n=j$ terms in (C.2),

$$
\begin{equation*}
\left\langle y^{k}\right\rangle_{\mathcal{N}}=\sum_{j=0}^{k}\binom{k}{j} b^{k-j} M_{j} \Phi_{j, j} . \tag{C.4}
\end{equation*}
$$

Since $\lim _{\sigma_{y} \rightarrow 0} \Phi_{j, j}=1$, eq. (5.15) is recovered.
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[^0]:    ${ }^{1} \mathrm{~A}$ detailed discussion and a list of references can be found in [14].

[^1]:    ${ }^{2}$ Sudakov terms (referred to as Sudakov shoulders) located in the middle of the phase-space can also be present when singularities from real and virtual perturbative corrections do not compensate each other $[25,26]$.

[^2]:    ${ }^{3}$ The literature on the shape function, on its properties and on the RGE evolution is wide, a list of references is in [66].
    ${ }^{4}$ Perturbative corrections to the shape function and to its moments are discussed in [60, 65].

