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Inquiring electromagnetic quantum fluctuations about the orientability of space

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Abstract Orientability is an important global topological property of spacetime manifolds. It is often assumed that a test for spatial orientability requires a global journey across the whole 3-space to check for orientation-reversing paths. Since such a global expedition is not feasible, theoretical arguments that combine universality of physical experiments with local arrow of time, CP violation and CPT invariance are usually offered to support the choosing of time- and space-orientable spacetime manifolds. Another theoretical argument also offered to support this choice comes from the impossibility of having globally defined spinor fields on non-orientable spacetime manifolds. In this paper, we argue that it is possible to locally access spatial orientability of Minkowski empty spacetime through physical effects involving quantum vacuum electromagnetic fluctuations. We study the motions of a charged particle and a point electric dipole subject to these electromagnetic fluctuations in Minkowski spacetime with orientable and non-orientable spatial topologies. We derive analytic expressions for a statistical orientability indicator for both of these point-like particles in two inequivalent spatially flat topologies. For the charged particle, we show that it is possible to distinguish the orientable from the non-orientable topology by contrasting the time evolution of the orientability indicators. This result reveals that it is possible to access orientability through electromagnetic quantum vacuum fluctuations. However, the answer to the central question of the paper, namely how to locally probe the orientability of Minkowski 3-space intrinsically, comes about only in the study of the motions of an electric dipole. For this point-like particle, we find that a characteristic inversion pattern exhibited by the curves of the orientability statistical indicator is a signature of non-orientability. This result makes it clear that it is possible to locally unveil spatial non-orientability through

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the inversion pattern of curves of our orientability indicator for a point electric dipole under quantum vacuum electromagnetic fluctuations. Our findings might open the way to a conceivable experiment involving quantum vacuum electromagnetic fluctuations to locally probe the spatial orientability of Minkowski empty spacetime.

1 Introduction

The Universe is modeled as a four-dimensional differentiable manifold, which is a topological space with an additional differential structure that permits to locally define connections, metric and curvature with which the gravitation theories are formulated. Geometry is a local attribute that brings about curvature, whereas topology is a global feature of a manifold related, for example, to compactness and orientability. Geometry constrains but does not specify the topology. So, topologically different manifolds can have a given geometry. I

Since topology antecedes geometry, it is important to determine whether, how and to what extent physical phenomena depend upon or are somehow affected, induced, triggered, or even driven by a nontrivial topology. The net role played by the spatial topology is more clearly ascertained in the static spatially flat Friedmann–Lemaître–Robertson–Walker spacetime, whose dynamical degrees of freedom are frozen. Thus, in this work we focus on Minkowski spacetime, whose spatial geometry is Euclidean.

¹ Despite our present-day inability to predict the spatial topology of the Universe from a fundamental theory, one should be able to probe it through cosmic microwave background radiation (CMBR) or (and) stochastic primordial gravitational waves [1–8], which should follow some basic detectability conditions [9–14]. For recent topological constraints from CMBR data we refer the reader to Refs. [15–17]. For some limits on the circles-in-the-sky method designed for searches of cosmic topology through CMBR see Ref. [18].



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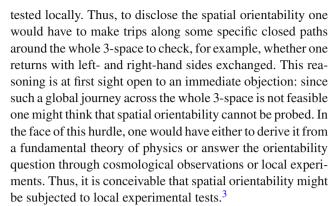
Although the topology of the spatial section, M_3 , of Minkowski spacetime, $\mathcal{M}_4 = \mathbb{R} \times M_3$, is usually taken to be the simply-connected Euclidean space \mathbb{E}^3 , it is a mathematical fact that it can also be any one of the possible 17 topologically distinct quotient (multiply-connected) manifolds $M_3 = \mathbb{E}^3/\Gamma$, where Γ is a discrete group of isometries or holonomies acting freely on the covering space \mathbb{E}^3 [19,20]. The action of Γ tiles or tessellates the covering manifold into identical domains or cells which are copies of what is known as fundamental polyhedron (FP) or fundamental cell or domain (FC or FD). On the covering manifold \mathbb{E}^3 , the multiple connectedness of M_3 is taken into account by imposing periodic boundary conditions (repeated cells) that are determined by the action of the group of discrete isometries Γ on the covering space \mathbb{E}^3 .

In a manifold with periodic boundary conditions only certain modes of fields can exist. Thus, a nontrivial topology may leave its mark on the expectation values of local physical quantities. A case in point is the Casimir effect of topological origin [21–26].

Quantum vacuum fluctuations of the electromagnetic field may give rise to stochastic motions for charged test particles [27–34]. In Minkowski spacetime with the standard simply-connected spatial section, it is unsettled whether such motions of test particles can happen [35,36]. What is theoretically clear is that when changes of a topological nature are made in the background 3-space as, for example, the insertion of perfectly reflecting planes into the three-dimensional spatial section of Minkowski spacetime, the resulting mean squared velocity of a charged test particle is affected [27–30,32–34].

In a recent paper, it was shown that endowing the spatial sections of Minkowski spacetime with a nontrivial topology influences the stochastic motions performed by a test charged particle under vacuum fluctuations of the electromagnetic field [37]. Thus, either by inserting perfectly reflecting boundaries or equipping 3-space with any of the classified nontrivial topologies, the stochastic motions of charged particles under quantum vacuum electromagnetic fluctuations are changed in a way that is measurable in principle.²

Orientability is an important global topological property of a spacetime manifold. It is widely assumed, implicitly or explicitly, that a manifold modeling the physical spacetime is globally orientable in all respects. Namely, that it is spacetime orientable and, additionally, that it is separately time and space orientable. Besides, it is also generally assumed that, being a global property, the 3-space orientability cannot be



Since quantum vacuum fluctuations of the electromagnetic field can be used to disclose a putative nontrivial 3-space topology of Minkowski spacetime through stochastic motions of test charged particles [37], and given that 8 out of the possible 17 quotient flat 3-manifolds are non-orientable [19], a question that naturally arises is whether these quantum vacuum fluctuations could be also used to reveal locally specific topological properties such as orientability of 3-space.⁴

Our chief goal in this article is to address this question by inquiring the electromagnetic quantum fluctuations about the spatial orientability of Minkowski spacetime. To this end, we investigate stochastic motions of a charged particle and of an electric dipole under quantum fluctuations of the electromagnetic field in Minkowski spacetime with two inequivalent spatial topologies, namely the orientable slab (E_{16}) and the non-orientable slab with flip (E_{17}).⁵ These topologies turn out to be suitable to identify orientability or non-orientability signatures by stochastic motions of point-like particles in Minkowski spacetime.

In Sect. 2 we introduce the notation and present some key concepts and results regarding topologies of three-dimensional manifolds, which will be needed in the rest of the paper. In Sect. 3 we present the physical systems along with the background geometry and topology, introduce the orientability statistical indicator and derive its expressions for both a charged particle and an electric dipole under quantum vacuum electromagnetic fluctuations in Minkowski spacetime with E_{16} and E_{17} flat 3-space topologies. For the charged particle, we show that by comparing the time evolution of the orientability indicator for a particle in E_{16} and E_{17} one can discriminate the orientable from the non-orientable topology. This result is significant in that it makes apparent



² In Refs. [27–34] one has a nontrivial inhomogeneous quotient orbifold space topology [20] while in Ref. [37] we have nontrivial flat smooth manifolds. We also note in this regard that the classification of three-dimensional Euclidean spaces was first taken up in the field of crystallography [38–40] and completed in 1934 [41]. For a recent exposition the reader is referred to Refs. [42–45].

³ One can certainly take advantage of gedanken experiments to reach theoretical conclusions, but not as a replacement to actual experimental evidence in physics [46].

⁴ In mathematical terms, this amounts to identifying physical signatures of the flip holonomy through the local motion of point-like particles under electromagnetic vacuum fluctuations.

⁵ In the next section we present a summary of flat three-dimensional topologies and some of their main topological properties. For a more detailed account on these topologies we recommend Refs. [42–45].

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the strength of our approach to access orientability through electromagnetic quantum vacuum fluctuations. However, a possible answer to the central question of the paper, namely how to locally probe the orientability of Minkowski 3-space per se, comes about only in the study of the motions of an electric dipole. In this regard, the most important finding is that spatial non-orientability can be locally unveiled through the inversion pattern of the curves of the orientability statistical indicator for a point electric dipole under quantum vacuum electromagnetic fluctuations. Section 4 is dedicated to our main conclusions and final remarks.

2 Topological prerequisites

Our primary aim in this section is to introduce the notation and give some basic definitions and results concerning the topology of flat three-dimensional manifolds that are used throughout this paper. The spatial section M_3 of the Minkowski spacetime manifold $\mathcal{M}_4 = \mathbb{R} \times M_3$ is usually assumed to be the simply connected Euclidean space \mathbb{E}^3 . But it can also be a multiply-connected quotient 3-manifold of the form $M_3 = \mathbb{E}^3/\Gamma$, where \mathbb{E}^3 is the covering space and Γ is a discrete and fixed-point-free group of discrete isometries (also referred to as the holonomy group [19,20]) acting freely on the covering space \mathbb{E}^3 [19].

Possibly the best known example of three-dimensional quotient Euclidean manifold with nontrivial topology is the 3-torus $T^3=\mathbb{S}^1\times\mathbb{S}^1\times\mathbb{S}^1=\mathbb{E}^3/\Gamma$, whose fundamental polyhedron (FP) is a parallelepiped with sides a,b,c (say), the opposite faces of which are identified through translations. In any multiply-connected quotient flat 3-manifold the fundamental polyhedron tiles (tessellates) the whole infinite simply-connected covering space \mathbb{E}^3 . The group $\Gamma=\mathbb{Z}\times\mathbb{Z}\times\mathbb{Z}$ consists of discrete translations associated with the face identification. The periodicities in the three independent directions are given by the circles \mathbb{S}^1 .

In forming the quotient manifolds M_3 an essential point is that they are obtained from the covering manifold \mathbb{E}^3 by identifying points that are equivalent under the action of the discrete isometry group Γ . Hence, each point on the quotient manifold M_3 represents all the equivalent points on the covering space. The multiple connectedness leads to periodic boundary conditions on the covering manifold \mathbb{E}^3 (repeated cells) that are determined by the action of the group Γ on the covering manifold. Clearly, different isometry groups Γ define different topologies for M_3 , which in turn give rise to different periodicity on the covering manifold (different mosaic of the covering space \mathbb{E}^3).

Another important point in forming the flat quotient manifolds M_3 is that every covering isometry $\gamma \in \Gamma$ can be expressed (in the covering space \mathbb{E}^3) through translation, rotation, reflection (flip) and combinations thereof. A screw motion, for example, is a combination of a rotation $R(\alpha, \widehat{\mathbf{u}})$ by an angle α around an axis $\widehat{\mathbf{u}}$, followed by a translation along a vector $\mathbf{L} = L\widehat{\mathbf{w}}$, say. A general glide reflection is combination of a reflection followed by a translation, as for example $P = (x, y, z) \mapsto \gamma P = (-x, y, z) + (0, 0, c)$, where c is a constant. If c = 0 we have a simple reflection or flip.

In dealing with metric manifolds in mathematical physics two concepts of homogeneity arise. Local homogeneity is a geometrical characteristic of metric manifolds. It is formulated in terms of the action of the group of local isometries. In dealing with topological spaces, we have the global homogeneity of topological nature. A way to characterize global homogeneity of the quotient manifolds is through distance functions. Indeed, for any $\mathbf{x} \in M_3$ the distance function $\ell_\gamma(\mathbf{x})$ for a given discrete isometry $\gamma \in \Gamma$ is defined by

$$\ell_{\gamma}(\mathbf{x}) = d(\mathbf{x}, \gamma \mathbf{x}),\tag{1}$$

where d is the Euclidean metric defined on M_3 . The distance function provides the length of the closed geodesic that passes through \mathbf{x} and is associated with a holonomy γ . In *globally homogeneous* manifolds the distance function for any covering isometry γ is constant. In globally inhomogeneous manifolds, in contrast, the length of the closed geodesic associated with at least one γ is non-translational (screw motion or flip, for example) and depends on the point $\mathbf{x} \in M_3$, and then is not constant.

When the distance between a point \mathbf{x} and its image $\gamma \mathbf{x}$ (in the covering space) is a constant for all points \mathbf{x} then the holonomy γ is a translation, that is, all elements of the covering group Γ in globally homogeneous spaces are translations. This means that in these manifolds the faces of the fundamental cells are identified through independent translations.

In this paper, we shall consider the topologically nontrivial spaces E_{16} and E_{17} . The slab space E_{16} is constructed by tessellating \mathbb{E}^3 by equidistant parallel planes, so it has only one compact dimension associated with a direction perpendicular to those planes. Taking the x-direction as compact, one has that, with $n_x \in \mathbb{Z}$ and a > 0, points (x, y, z) and $(x + n_x a, y, z)$ are identified in the case of the slab space E_{16} . The slab space with flip E_{17} involves an additional inversion of a direction orthogonal to the compact direction, that is, one direction in the tessellating planes is flipped as one moves from one plane to the next. Letting the flip be in the y-direction, the identification of points (x, y, z) and $(x + n_x a, (-1)^{n_x} y, z)$ defines the E_{17} topology. In this way, the slab space E_{16} is globally homogeneous, whereas the slab space with flip, E_{17} , is globally inhomogeneous since



 $^{^6}$ \mathbb{R}^3 is a topological space while \mathbb{E}^3 is a geometrical space, i.e. \mathbb{R}^3 endowed with the Euclidean metric.

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Table 1 Names and symbols of the simply-connected $\mathbb{E}^3=E_{18}$, and two multiply-connected flat orientable and non-orientable Euclidean quotient manifolds $M_3=\mathbb{E}^3/\Gamma$ along with the number of compact dimensions (Comp.), orientability and global (topological) homogeneity

Name	Symb.	Comp	Orientable	Homogeneous
Slab space	E_{16}	1	Yes	Yes
Slab space with flip	E_{17}	1	No	No
Euclidean space	E_{18}	0	Yes	Yes

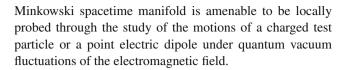
the covering group Γ contains a flip, which clearly is a non-translational discrete isometry.

Orientability is another very important global (topological) property of a manifold that measures whether one can choose consistently a clockwise orientation for loops in the manifold. A closed curve in a manifold M_3 that brings a traveler back to the starting point mirror-reversed is called an orientation-reversing path. Manifolds that do not have an orientation-reversing path are called *orientable*, whereas manifolds that contain an orientation-reversing path are nonorientable [47]. Most surfaces that we encounter, such as cylinders, planes and tori are orientable, whereas the Möbius strip and Klein bottle are non-orientable surfaces. For threedimensional quotient manifolds, when the covering group Γ contains at least one holonomy γ that is a reflection (flip) the corresponding quotient manifold is non-orientable. Thus, for example, the slab space is orientable while the slab space with flip is non-orientable. Clearly non-orientable manifolds are necessarily globally inhomogeneous as the covering group Γ contains a reflection, which obviously is a non-translational covering holonomy.

In Table 1 we collect the names and symbols used to refer to the manifolds together with the number of compact independent dimensions and information concerning their global homogeneity and orientability. In addition to the simply-connected \mathbb{E}^3 topology, these are the three-dimensional manifolds with nontrivial topologies that we shall be concerned with in this paper.

Finally let us briefly mention a few results that are implicity or explicitly used throughout this paper (see Ref. [58] for a detailed discussion). Every simply-connected spacetime manifold is both time- and space-orientable. The product of two manifolds is simply-connected if and only if each factor is. If the spacetime is of the form $\mathcal{M}_4 = \mathbb{R} \times M_3$ then space-orientability of the spacetime reduces to orientability of the spatial section M_3 . This applies to the spacetime endowed with the non-orientable 3-space topology E_{17} that we deal with in this work.

Having set the stage for our investigation, in the next section we proceed to inquire whether the topological (global) non-orientability property of the spatial section of



3 Non-orientability from electromagnetic fluctuations

As noted before, quantum vacuum fluctuations of the electromagnetic field give rise to stochastic motions for charged particles which are sensitive to the topology of the background 3-space. In this section, we address the main underlying question of this paper, which is whether these fluctuations offer a suitable way of discovering a putative non-orientability of Minkowski spatial sections. We take up this question through the study of stochastic motions of a charged particle and an electric dipole under electromagnetic quantum fluctuations in Minkowski spacetime with two inequivalent spatial topologies, namely the orientable slab space (E_{16}) and the non-orientable slab space with flip (E_{17}) . In the following we present the details of our investigation and main results.

3.1 Non-orientability with point charged particle

We first consider a nonrelativistic test particle with charge q and mass m locally subject to vacuum fluctuations of the electric field $\mathbf{E}(\mathbf{x},t)$ in a topologically nontrivial space-time manifold equipped with the Minkowski metric $\eta_{\mu\nu} = \text{diag}(+1,-1,-1,-1)$. The spatial section is usually taken to be \mathbb{E}^3 , but here we take for M_3 each of the two multiply-connected manifolds in Table 1.

Locally, the motion of the charged test particle is determined by the Lorentz force. In the nonrelativistic limit the equation of motion for the point charge is

$$\frac{d\mathbf{v}}{dt} = \frac{q}{m} \mathbf{E}(\mathbf{x}, t) \,, \tag{2}$$

where \mathbf{v} is the particle's velocity and \mathbf{x} its position at time t. We assume that on the time scales of interest the particle practically does not move, i.e. it has a negligible displacement, so we can ignore the time dependence of \mathbf{x} . Thus, the particle's position \mathbf{x} is taken as constant in what follows [27,33]. Assuming that the particle is initially at rest, integration of Eq. (2) gives

$$\mathbf{v}(\mathbf{x},t) = \frac{q}{m} \int_0^t \mathbf{E}(\mathbf{x},t') dt', \qquad (3)$$

and the mean squared velocity, velocity dispersion or simply dispersion in each of the three independent directions i = 1



 $^{^{7}\,}$ The corrections arising from the inexactness of this assumption are negligible in the low velocity regime.

Table 2 Spatial separation in Hadamard function for the simply-connected Euclidean manifold E_{18} , and for the multiply-connected flat orientable (E_{16}) and its non-orientable counterpart (E_{17}) quotient Euclidean manifolds. The topological compact length is denoted by a. The numbers n_x are integers and run from $-\infty$ to ∞ . For each multiply-connected topology, when $n_x = 0$ we recover the spatial separation for the simply-connected Euclidean 3-space

Spatial topology	Spatial separation r^2 for Hadamard function
E ₁₆ -slab space	$(x - x' - n_x a)^2 + (y - y')^2 + (z - z')^2$
E_{17} -slab space with flip	$(x - x' - n_x a)^2 + (y - (-1)^{n_x} y')^2 + (z - z')^2$
E ₁₈ -Euclidean space	$(x - x')^2 + (y - y')^2 + (z - z')^2$

x, y, z is given by⁸

$$\left\langle \Delta v_i^2 \right\rangle = \frac{q^2}{m^2} \int_0^t \int_0^t \left\langle E_i(\mathbf{x}, t') E_i(\mathbf{x}, t'') \right\rangle dt' dt'' \,. \tag{4}$$

Following Yu and Ford [27], we assume that the electric field is a sum of classical \mathbf{E}_c and quantum \mathbf{E}_q parts. Because \mathbf{E}_c is not subject to quantum fluctuations and $\langle \mathbf{E}_q \rangle = 0$, the two-point function $\langle E_i(\mathbf{x},t)E_i(\mathbf{x}',t')\rangle$ in Eq. (4) involves only the quantum part of the electric field [27].

It can be shown [48] that locally

$$\langle E_{i}(\mathbf{x}, t) E_{i}(\mathbf{x}', t') \rangle = \frac{\partial}{\partial x_{i}} \frac{\partial}{\partial x'_{i}} D(\mathbf{x}, t; \mathbf{x}', t') - \frac{\partial}{\partial t} \frac{\partial}{\partial t'} D(\mathbf{x}, t; \mathbf{x}', t')$$
(5)

where in Minkowski spacetime with $M_3 = \mathbb{E}^3 = E_{18}$, the Hadamard function $D(\mathbf{x}, t; \mathbf{x}', t')$ is given by

$$D_0(\mathbf{x}, t; \mathbf{x}', t') = \frac{1}{4\pi^2 (\Delta t^2 - |\Delta \mathbf{x}|^2)}.$$
 (6)

The subscript 0 indicates standard Minkowski spacetime, $\Delta t = t - t'$ and $|\Delta \mathbf{x}| \equiv r$ is the spatial separation for topologically trivial Minkowski spacetime:

$$r^{2} = (x - x')^{2} + (y - y')^{2} + (z - z')^{2}.$$
 (7)

In Minkowski spacetime with a topologically nontrivial spatial section, the spatial separation r^2 takes a different form that captures the periodic boundary conditions imposed on the covering space \mathbb{E}^3 by the covering group Γ , which characterize the spatial topology. In consonance with Ref. [23], in Table 2 we collect the spatial separations for the topologically inequivalent Euclidean spaces we shall address in this paper.

Orientability indicator: slab space with flip E₁₇

For the sake of brevity, we present detailed calculations only for a charged particle in Minkowski spacetime with E_{17} spatial topology. The corresponding results for E_{16} can then be easily obtained from those for E_{17} , as we show below.

To obtain the correlation function for the electric field that is required to compute the velocity dispersion (4) for slab space with flip E_{17} , we replace in Eq. (5) the Hadamard function $D(\mathbf{x}, t; \mathbf{x}', t')$ by its renormalized version given by [37]

$$D_{ren}(\mathbf{x}, t; \mathbf{x}', t') = D(\mathbf{x}, t; \mathbf{x}', t') - D_0(\mathbf{x}, t; \mathbf{x}', t')$$

$$= \sum_{n_x = -\infty}^{\infty} \frac{1}{4\pi^2 (\Delta t^2 - r^2)},$$
(8)

where here and in what follows $\sum_{i=1}^{\infty} f_{i}$ indicates that the Minkowski contribution term $n_{x}=0$ is excluded from the summation, $\Delta t=t-t'$, and, from Table 2, the spatial separation for E_{17} is

$$r^{2} = (x - x' - n_{x}a)^{2} + (y - (-1)^{n_{x}}y')^{2} + (z - z')^{2}.$$
 (9)

The term with $n_x = 0$ in the sum (8) is the Hadamard function $D_0(\mathbf{x}, t; \mathbf{x}', t')$ for Minkowski spacetime with simply-connected spatial section E_{18} . This term has been subtracted out from the sum because it gives rise to an infinite contribution to the velocity dispersion.

Thus, from Eq. (5) the renormalized correlation functions

$$\langle E_{i}(\mathbf{x}, t) E_{i}(\mathbf{x}', t') \rangle_{ren} = \frac{\partial}{\partial x_{i}} \frac{\partial}{\partial x'_{i}} D_{ren}(\mathbf{x}, t; \mathbf{x}', t') - \frac{\partial}{\partial t} \frac{\partial}{\partial t'} D_{ren}(\mathbf{x}, t; \mathbf{x}', t')$$
(10)

are then given by

$$\langle E_x(\mathbf{x}, t) E_x(\mathbf{x}', t') \rangle_{ren}^{E_{17}} = \sum_{n_x = -\infty}^{\infty} \frac{\Delta t^2 + r^2 - 2r_x^2}{\pi^2 [\Delta t^2 - r^2]^3},$$
 (11)

$$\langle E_{y}(\mathbf{x},t)E_{y}(\mathbf{x}',t')\rangle_{ren}^{E_{17}} = \sum_{n_{x}=-\infty}^{\infty} \left\{ \frac{(3-(-1)^{n_{x}}) \Delta t^{2}}{2\pi^{2} [\Delta t^{2}-r^{2}]^{3}} + \frac{(1+(-1)^{n_{x}}) r^{2}-4(-1)^{n_{x}} r_{y}^{2}}{2\pi^{2} [\Delta t^{2}-r^{2}]^{3}} \right\},$$
(12)

$$\langle E_z(\mathbf{x}, t) E_z(\mathbf{x}', t') \rangle_{ren}^{E_{17}} = \sum_{n_x = -\infty}^{\infty} \frac{\Delta t^2 + r^2 - 2r_z^2}{\pi^2 [\Delta t^2 - r^2]^3},$$
 (13)

where $\Delta t = t - t'$ and

$$r_x = x - x' - n_x a,$$
 $r_y = y - (-1)^{n_x} y',$
 $r_z = z - z',$ $r = \sqrt{r_x^2 + r_y^2 + r_z^2}.$ (14)

The orientability indicator $I_{v_i^2}^{E_{17}}$ that we will consider is defined by replacing the electric field correlation functions



⁸ By definition, $\langle \Delta v_i^2(\mathbf{x}, t) \rangle = \langle v_i^2(\mathbf{x}, t) \rangle - \langle v_i(\mathbf{x}, t) \rangle^2$.

⁹ The reader is referred to Refs. [43–45] for pictures of the fundamental cells and further properties of all possible three-dimensional Euclidean topologies.

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in Eq. (4) by their renormalized counterparts (8) in which r is given by (9):

$$I_{v_i^2}^{E_{17}}(\mathbf{x},t) = \frac{q^2}{m^2} \int_0^t \int_0^t \left\langle E_i(\mathbf{x},t') E_i(\mathbf{x},t'') \right\rangle_{ren}^{E_{17}} dt' dt''.$$
(15)

From (8) it is clear that the orientability indicator $I_{v_i^2}^{E_{17}}$ is the difference between the velocity dispersion in E_{17} and the one in Minkowski with trivial (simply connected) topology.

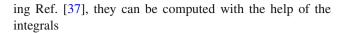
Before proceeding to the calculation of the components of the orientability indicator (15) some words of clarification are in order for the sake of generality and for later use. From Eqs. (4) and (8) a general definition of the orientability indicator can be written in the form

$$I_{v_i^2}^{MC} = \left\langle \Delta v_i^2 \right\rangle^{MC} - \left\langle \Delta v_i^2 \right\rangle^{SC},\tag{16}$$

where $\langle \Delta v_i^2 \rangle$ is the mean square velocity dispersion, and the superscripts MC and SC stand for multiply- and simply-connected manifolds, respectively. The right-hand side of (16) is defined by first taking the difference of the two terms with $\mathbf{x}' \neq \mathbf{x}$ and then setting $\mathbf{x}' = \mathbf{x}$. Since $I_{v_i^2}^{MC}$

is not the velocity dispersion $(\Delta v_i^2)^{MC}$ but the difference (16), the possibility that it takes negative values should not be cause of concern, 10 a point that does not seem to have been appreciated in some previous works in which this indicator was implicity used [27-30,32-34,37,49] together with the particular assumption that the second term vanishes. A similar statistical indicator that measures the departure of a statistical quantity from its values for the simply-connected space comes about in cosmic crystallography, which is an approach to detect cosmic topology from the distribution of discrete cosmic sources [50]. Indeed, a topological signature of any multiply connected 3-manifold of constant curvature is given by a constant times the difference $\Phi_{exp}^{MC}(s_i) - \Phi_{exp}^{SC}(s_i)$ of the expected pair separation histogram (EPSH) corresponding to the multiply connected manifold minus the EPSH for the underlying simply connected covering manifold [51,52], whose expression can be derived in an analytical form [51,53].

We shall discuss experimental features of the orientability indicator (16) in the two closing paragraphs of Sect. 3.2 in this section. But now let us return to the calculation of the components of the orientability indicator $I_{v^2}^{E_{17}}$. Follow-



$$\mathcal{I} = \int_0^t \int_0^t dt' dt'' \frac{1}{(\Delta t^2 - r^2)^3}$$

$$= \frac{t}{16r^5(t^2 - r^2)} \left\{ 4rt - 3(r^2 - t^2) \ln \frac{(r - t)^2}{(r + t)^2} \right\}$$
(17)

and

$$\mathcal{J} = \int_0^t \int_0^t dt' dt'' \frac{\Delta t^2}{(\Delta t^2 - r^2)^3}$$

$$= \frac{t}{16r^3(t^2 - r^2)} \left\{ 4rt + (r^2 - t^2) \ln \frac{(r - t)^2}{(r + t)^2} \right\}, \quad (18)$$

in which $\Delta t = t' - t''$.

By using these integrals and Eqs. (11) to (14) with $\mathbf{x}' = \mathbf{x}$ in Eq. (15) we find

$$I_{v_{x}^{2}}^{E_{17}}(\mathbf{x},t) = \sum_{n_{x}=-\infty}^{\infty} \frac{q^{2}t}{16\pi^{2}m^{2}r^{5}(t^{2}-r^{2})} \left\{ 4rt(\bar{r}_{x}^{2}+r^{2}) + (t^{2}-r^{2})(3\bar{r}_{x}^{2}-r^{2}) \ln\frac{(r-t)^{2}}{(r+t)^{2}} \right\},$$
(19)
$$I_{v_{y}^{2}}^{E_{17}}(\mathbf{x},t) = \sum_{n_{x}=-\infty}^{\infty} \frac{q^{2}t}{32\pi^{2}m^{2}r^{5}(t^{2}-r^{2})} \left\{ 4rt\bar{r}_{y}^{2} + 4rt(3-(-1)^{n_{x}})r^{2} + (t^{2}-r^{2})\left[3\bar{r}_{y}^{2}-(3-(-1)^{n_{x}})r^{2}\right] \ln\frac{(r-t)^{2}}{(r+t)^{2}} \right\},$$
(20)
$$I_{v_{z}^{2}}^{E_{17}}(\mathbf{x},t) = \sum_{n_{x}=-\infty}^{\infty} \frac{q^{2}t}{16\pi^{2}m^{2}r^{5}(t^{2}-r^{2})} \left\{ 4rt(\bar{r}_{z}^{2}+r^{2}) + r^{2} \right\},$$
(20)

$$I_{v_z^2}^{E_{17}}(\mathbf{x},t) = \sum_{n_x = -\infty}^{\infty} \frac{q^2 t}{16\pi^2 m^2 r^5 (t^2 - r^2)} \left\{ 4rt(\bar{r}_z^2 + r^2) + (t^2 - r^2)(3\bar{r}_z^2 - r^2) \ln \frac{(r-t)^2}{(r+t)^2} \right\},\tag{21}$$

where

$$r = r_{n_x} = \sqrt{n_x^2 a^2 + 2(1 - (-1)^{n_x})y^2},$$
(22)

$$\bar{r}_x^2 = r^2 - 2r_x^2 = -n_x^2 a^2 + 2(1 - (-1)^{n_x})y^2,$$

$$\bar{r}_y^2 = (1 + (-1)^{n_x})r^2 - 8(-1)^{n_x}(1 - (-1)^{n_x})y^2$$
(23)

$$= (1 + (-1)^{n_x})n_x^2 a^2 + 8(1 - (-1)^{n_x})y^2, \tag{24}$$

$$\bar{r}_z^2 = r^2 - 2r_z^2 = r^2 \,, (25)$$

as follows from Eq. (14) in the coincidence limit.

Before proceeding to E_{16} topology, as a consistency check of the above calculations we discuss the topological Minkowskian limit for the orientability indicator $I_{v_i^2}^{E_{17}}$. We begin by recalling that compact lengths associated with Euclidean quotient manifolds are not fixed. Different values of a correspond to different 3-manifolds with the same topology. Intuitively, by letting the compact length $a \to \infty$ the role of the compactness (nontrivial topology) should disappear and a topological Minkowskian simply-connected limit



¹⁰ It is neither experimentally nor theoretically settled whether the second term on the right-hand side of Eq. (16) vanishes or not. Here we take the more general view that it is nonzero. It is only when the rather particular assumption that it vanishes is made that one encounters counterintuitive negative values for mean square velocities often found in the literature [27–30,32–34,37,49]. This may be looked upon as an indication that the simply-connected term in Eq. (16) should not vanish.

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for the orientability indicator should be attained, which is zero by definition (16). In order to verify if this is the case, we first note that from Eq. (22) it follows that letting $a \to \infty$ amounts to letting $r \to \infty$. For very large r each term of the sum (19) consists of a fraction whose numerator is dominated by a power of r not bigger than the fourth (the logarithmic term tends to zero as $r \to \infty$) whereas the denominator becomes proportional to r^7 . Therefore each term of the sum vanishes in the limit $a \to \infty$ and $I_{v_x^2}^{E_{17}}$ is zero. The same argument shows that the other components of the orientability indicator also vanish in the limit $a \to \infty$, in agreement with (16).

In the opposite limit of small a the effects of the topology become arbitrarily large: the indicators (19)–(21) grow proportionally to $1/a^2$ as $a \to 0$, and the same is true regarding the indicators for E_{16} given by Eqs. (26) and (27) below. This behavior is intuitively expected since for $a \ll t$ in the time interval of interest an electromagnetic signal would cross an enormous number of fundamental cells, and the effect of the nontrivial topology would be huge.

Orientability indicator: slab space E₁₆

The factors of $(-1)^{n_x}$ that appear in Eqs. (12) and (20) arise from derivatives with respect to y' in Eq. (10) contributed by the separation r given by Eq. (9). Hence, the results for E_{16} are immediately obtained from those for E_{17} by simply replacing $(-1)^{n_x}$ by 1 everywhere in Eqs. (19)–(25). This leads to

$$I_{v_x^2}^{E_{16}}(\mathbf{x},t) = -\frac{q^2 t}{4\pi^2 m^2} \sum_{n_x = -\infty}^{\infty} \frac{1}{n_x^3 a^3} \ln \frac{(n_x a - t)^2}{(n_x a + t)^2},$$
 (26)

$$I_{v_y^2}^{E_{16}}(\mathbf{x},t) = I_{v_z^2}^{E_{16}}(\mathbf{x},t) = \frac{q^2 t}{8\pi^2 m^2} \sum_{n_x = -\infty}^{\infty} \left\{ \frac{4t}{n_x^2 a^2 (t^2 - n_x^2 a^2)} + \frac{1}{n_x^2 a^3} \ln \frac{(n_x a - t)^2}{(n_x a + t)^2} \right\},$$
(27)

in agreement with the results obtained in Refs. [37,49], where the indicator (16) was used.

Non-orientability with a charged particle: conclusions

The components of the orientability indicator are singular at $t = r = r_{n_x}$, where $r_{n_x} = |n_x|a$ for E_{16} while r_{n_x} is given by Eq. (22) for E_{17} . In both cases, the singularities are of topological origin and correspond to each of the infinitely many distances r_{n_x} that arise from the periodic boundary conditions imposed by the topologies on the covering space \mathbb{E}^3 , which are given by the identifications $(x, y, z) \leftrightarrow (x + n_x a, y, z)$ and $(x, y, z) \leftrightarrow (x + n_x a, (-1)^{n_x} y, z)$ for E_{16} and E_{17} , respectively. Incidentally, we note that for a quotient orbifold arising from the action of \mathbb{Z}_2 on \mathbb{E}^3 by reflection [20], which is a topological space resulting from the insertion of a perfectly reflecting plane often used in the literature [27,30,33,34,49], the indicator (16) exhibits only one singularity as should be

expected from its origin, i.e. from the orbifold topology properties [20,37,49]. This singularity has been ascribed to local properties of the physical system, and functions of time to switch on and off the interaction between the particle and the electromagnetic field have been suggested to regularize it [33,34,54]. However, this procedure does not seem appropriate to cope with singularities of topological nature as the one provided by the reflection orbifold. On the other hand, it is also possible that in being of topological origin these singularities are not only unavoidable but also unremovable in the sense that one can remove them only by 'switching off' the topology.

An important question that arises at this point is what then we ultimately learn from the above calculations regarding the local test of spatial orientability by studying the motions of a charged particle under electromagnetic quantum fluctuations. In other words, what these fluctuations are teaching us about orientability, Eqs. (19)–(21) for the non-orientable E_{17} topology and Eqs. (26)–(27) for the orientable E_{16} space topology. Let us now discuss this capital issue. Clearly, conclusions can only be reached by comparisons between the stochastic motions of the charged test particles lying in space manifolds with each of the two topologies. In this regard, a first difficulty one encounters is how to make a proper comparison because E_{16} is globally homogenous whereas E_{17} is not (cf. Table 1). This means that the orientability indicator does not depend on the particle's position for E_{16} , but it does when the particle lies in a space with the globally inhomogeneous topology E_{17} . The functional dependence of the dispersion on the particle's position coordinates in these manifolds makes apparent the first difficulty. Indeed, the components of the orientability indicator (19)-(21) for E_{17} depend on the y-coordinate, while the components (26)– (27) for E_{16} do not. Thus, one has to suitably choose the point P = (x, y, z) in E_{17} for the particle's position in order to make a proper comparison between the orientability indicator curves for the topologically homogeneous E_{16} and the topologically inhomogeneous E_{17} manifolds. From the identification of (x, y, z) and $(x + n_x a, (-1)^{n_x} y, z)$ that defines the E_{17} topology, clearly a suitable way to freeze out the global inhomogeneity degree of freedom, and thus isolate the non-orientability effect, is by choosing as the particle's position the point $P_0 = (x, 0, z)$. Since our chief concern is orientability, in all figures in this paper but one (Fig. 1) we choose this point as the particle's position when dealing with E_{17} topology.

Having circumvented this particle position difficulty related to the topological inhomogeneity of E_{17} , in Fig. 1 we illustrate how global inhomogeneity affects the time evolution of the orientability indicator (19). For two particle's positions, one with y=0 and another with y=1/2 for compact length a=1, in E_{17} there arise different curves. This shows that the indicator (16) is able to capture the non-



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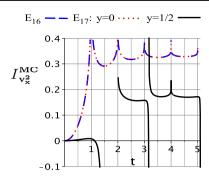


Fig. 1 Time evolution of the *x*-component of the orientability indicator $I_{v_x^2}^{MC}(\mathbf{x},t)$ in units of q^2/m^2 for a test particle with mass m and charge q in Minkowski spacetime with spatial section endowed with the non-orientable and global inhomogeneous E_{17} and orientable E_{16} topologies, both with compact length a=1. We show one curve for the globally homogeneous E_{16} (dashed line) and two curves for E_{17} : dotted and solid lines, for the particle at the positions $P_0=(x,0,z)$ and P=(x,1/2,z), respectively. The figure illustrates the topological inhomogeneity of E_{17} , and shows that when the degree of inhomogeneity is frozen the indicator curves for E_{17} [for the particle at $P_0=(x,0,z)$] and E_{16} [for the particle at generic P=(x,y,z)] coincide

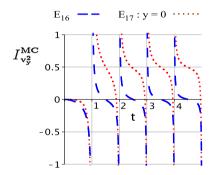


Fig. 2 Time evolution of the *y*-component of the orientability indicator $I_{v_y^2}^{MC}$ in units of q^2/m^2 for a charged test particle in Minkowski spacetime with spatial section endowed with the orientable E_{16} and non-orientable E_{17} topologies, both with compact length a=1. We show curves for E_{16} (dashed line) and for E_{17} with the particle at $P_0=(x,0,z)$ (dotted line). The indicator curves exhibit similar periodic patterns for time intervals of the order of one, but it is possible to distinguish the two topologies by contrasting the *y*-component of their orientability indicator

homogeneity of topological origin as has been pointed out in [37]. To allow reproduction, we mention that Figs. 1 and 2 arise from Eqs. (19)–(25) as well as (26) and (27) with compact length a = 1 and $n_x \neq 0$ ranging from -50 to 50.

Figure 1 also shows that for the particle's position P_0 in E_{17} the orientability indicator curves coincide with those for a generic point in globally homogenous E_{16} . This shows that for $P_0 = (x, 0, z)$, where the global inhomogeneity degree of freedom is frozen, the x-component of the orientability indicator cannot be used to distinguish between the two orientable and non-orientable 3-spaces. With different pattern of curves, the same holds true for the z-components of the orientability indicator for E_{17} and E_{16} , but we do not show this

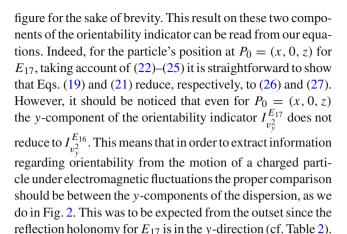


Figure 2 displays the y-components of the orientability indicator for a charged test particle in Minkowski space whose spatial section has E_{16} (orientable) and E_{17} [nonorientable, $P_0 = (x, 0, z)$] topologies. The chief conclusion is that the component along the direction of the flip (cf. Table 2) might be used to find out whether the particle lies in Minkowski spacetime with orientable or non-orientable 3-space. Figure 2 also shows different orientability indicator curves for E_{16} and E_{17} which, in both cases, repeat themselves periodically. For the two topologies the overall patterns of the orientability indicator curves are similar.

Although the above result seems valuable to the extent that it makes clear the strength of our approach to access orientability through electromagnetic vacuum fluctuations, it demands a comparison between the orientability indicator curves for the two given spatial manifolds for a verdict about orientability. Thus, it does not provide a conclusive answer to the central question of this paper, namely how to locally probe the orientability of the spatial section of Minkowski spacetime in itself from electromagnetic vacuum fluctuations. Given the directional properties of a point electric dipole, a pertinent question that emerges here is whether it would be more effective to use it for testing the individual non-orientability of a generic 3-space. If so we could envisage a local experiment to probe the orientability of a 3-space per se. In the next subsection we shall investigate this remarkable possibility.

3.2 Non-orientability with point electric dipole

A noteworthy outcome of the previous section is that the time evolution of the velocity dispersion for a charged particle can be used to locally differentiate an orientable (E_{16}) from a non-orientable (E_{17}) spatial section of Minkowski spacetime. However, it cannot be used to decide whether a given 3-space manifold is or not orientable. In this way, it cannot be taken as a definite answer to our central question about the spatial orientability of Minkowski spacetime. So, a question that naturally arises here is whether the velocity dispersion



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of a different type of point-like particle could provide a suitable local signature of non-orientability. As a point electric dipole has directional properties, one might expect that its orientability indicator would potentially be more effective in providing information regarding non-orientability associated with a particular direction. To examine this issue we now turn our attention to topologically induced motions of an electric dipole under quantum vacuum electromagnetic fluctuations.

Newton's second law for a point electric dipole of mass *m* in an external electric field reads

$$m\frac{d\mathbf{v}}{dt} = \mathbf{p} \cdot \nabla \mathbf{E}(\mathbf{x}, t) , \qquad (28)$$

where \mathbf{p} is the electric dipole moment. With the same hypotheses as for the point charge and assuming the dipole is initially at rest, integration of Eq. (28) yields

$$\mathbf{v}(\mathbf{x},t) = \frac{1}{m} p_j \int_0^t \partial_j \mathbf{E}(\mathbf{x},t') dt'$$
 (29)

with $\partial_j = \partial/\partial x_j$ and summation over repeated indices implied.

The mean squared speed in each of the three independent directions i = x, y, z is given by

$$\left\langle \Delta v_i^2 \right\rangle = \frac{p_j p_k}{m^2} \int_0^t \left\langle \left(\partial_j E_i(\mathbf{x}, t') \right) \left(\partial_k E_i(\mathbf{x}, t'') \right) \right\rangle dt' dt'' , \tag{30}$$

which can be conveniently rewritten as

$$\left\langle \Delta v_i^2 \right\rangle = \lim_{\mathbf{x}' \to \mathbf{x}} \frac{p_j p_k}{m^2} \int_0^t \int_0^t \partial_j \partial_k' \left\langle E_i(\mathbf{x}, t') E_i(\mathbf{x}', t'') \right\rangle dt' dt''$$
(31)

where $\partial_k' = \partial/\partial x_k'$.

Now we proceed to the computation of the velocity dispersion for a point dipole in spaces E_{17} and E_{16} . The space E_{17} has two topologically conspicuous directions: the compact x-direction and the flip y-direction associated with the non-orientability of E_{17} . To probe the non-orientability of E_{17} by means of stochastic motions, it seems most promising to choose a dipole oriented in the y-direction, since the orientation of the dipole would also be flipped upon every displacement by the topological length a along the compact direction. Indeed, it is for a dipole oriented in the flip direction that the effect of the non-orientability is most noticeable, as we show in the following.

For a dipole oriented along the y-axis the dipole moment is $\mathbf{p}=(0,p,0)$ and we again use the indicator in which the electric field correlation functions are replaced by their renormalized counterparts, as in Eq. (15) for the charged point particle. We have

$$= \lim_{\mathbf{x}' \to \mathbf{x}} \frac{p^2}{m^2} \int_0^t \int_0^t \partial_y \partial_{y'} \left\langle E_i(\mathbf{x}, t') E_i(\mathbf{x}', t'') \right\rangle_{ren}^{E_{17}} dt' dt'',$$
(32)

where the left superscript within parentheses indicates the dipole's orientation. With the help of Eq. (11) the x-component of the orientability indicator for the slab space with flip E_{17} takes the form

$$I_{v_x^2}^{(y)}(\mathbf{x},t) = \lim_{\mathbf{x}' \to \mathbf{x}} \frac{p^2}{\pi^2 m^2} \sum_{n_x = -\infty}^{\infty} \int_0^t \int_0^t dt' dt'' \partial_y \partial_{y'} \frac{\Delta t^2 + r^2 - 2r_x^2}{(\Delta t^2 - r^2)^3}.$$
(33)

with r defined by Eq. (9) and $\Delta t = t' - t''$, while r_x is given by Eq. (14). Making use of

$$\partial_{y}\partial_{y'}\frac{\Delta t^{2}+r^{2}-2r_{x}^{2}}{(\Delta t^{2}-r^{2})^{3}} = -4(-1)^{n_{x}} \left\{ \frac{2}{(\Delta t^{2}-r^{2})^{3}} + 3\frac{r^{2}-r_{x}^{2}+6r_{y}^{2}}{(\Delta t^{2}-r^{2})^{4}} + 24\frac{(r^{2}-r_{x}^{2})r_{y}^{2}}{(\Delta t^{2}-r^{2})^{5}} \right\}.$$
(34)

we find

$$I_{v_x^2}^{E_{17}}(\mathbf{x},t) = -\frac{4p^2}{\pi^2 m^2} \sum_{n_x = -\infty}^{\infty} (-1)^{n_x} \left\{ 2I_1 + 3(r^2 - r_x^2 + 6r_y^2)I_2 + 24(r^2 - r_x^2)r_y^2 I_3 \right\},$$
(35)

where, with $\Delta t = t' - t''$,

$$I_{1} = \mathcal{I} = \int_{0}^{t} \int_{0}^{t} \frac{dt'dt''}{(\Delta t^{2} - r^{2})^{3}} = \frac{t}{16} \left[\frac{4t}{r^{4}(t^{2} - r^{2})} + \frac{3}{r^{5}} \ln \frac{(r - t)^{2}}{(r + t)^{2}} \right],$$
(36)

$$I_{2} = \int_{0}^{t} \int_{0}^{t} \frac{dt'dt''}{(\Delta t^{2} - r^{2})^{4}} = \frac{1}{6r} \frac{\partial I_{1}}{\partial r} = \frac{t}{96} \left[\frac{4t(9r^{2} - 7t^{2})}{r^{6}(t^{2} - r^{2})^{2}} - \frac{15}{r^{7}} \ln \frac{(r - t)^{2}}{(r + t)^{2}} \right],$$
(37)

$$I_{3} = \int_{0}^{t} \int_{0}^{t} \frac{dt'dt''}{(\Delta t^{2} - r^{2})^{5}} = \frac{1}{8r} \frac{\partial I_{2}}{\partial r}$$

$$= \frac{t}{768} \left[\frac{4t(57t^{4} - 136r^{2}t^{2} + 87r^{4})}{r^{8}(t^{2} - r^{2})^{3}} + \frac{105}{r^{9}} \ln \frac{(r - t)^{2}}{(r + t)^{2}} \right].$$
(38)



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Similar calculations lead to

$$I_{v_y^2}^{(y)} I_{v_y^2}^{E_{17}}(\mathbf{x}, t) = -\frac{2p^2}{\pi^2 m^2} \sum_{n_x = -\infty}^{\infty'} (-1)^{n_x} \left\{ (5 - 3(-1)^{n_x}) I_1 + 6[r^2 + (7 - 6(-1)^{n_x}) r_y^2] I_2 + 48[r^2 - (-1)^{n_x} r_y^2] r_y^2 I_3 \right\}$$
(39)

and

$$I_{v_z^2}^{(y)} I_{v_z^2}^{E_{17}}(\mathbf{x}, t) = -\frac{4p^2}{\pi^2 m^2} \sum_{n_x = -\infty}^{\infty'} (-1)^{n_x} \left\{ 2I_1 + 3(r^2 + 6r_y^2)I_2 + 24r^2 r_y^2 I_3 \right\}.$$
(40)

Since the coincidence limit $\mathbf{x}' \to \mathbf{x}$ has been taken, it follows from Eq. (14) that in Eqs. (35)–(40) one must put

$$r = \sqrt{n_x^2 a^2 + 2(1 - (-1)^{n_x})y^2},$$

$$r_x^2 = n_x^2 a^2, \qquad r_y^2 = 2(1 - (-1)^{n_x})y^2.$$
(41)

It can be immediately checked that, as for the point charge, in the Minkowskian limit $(a \to \infty)$ the orientability indicator for a dipole is zero.

For the slab space E_{16} the components of the dipole orientability indicator are obtained from those for E_{17} by setting $r_x^2 = r^2$, $r_y = 0$, and replacing $(-1)^{n_x}$ by 1 everywhere. Therefore, we have

$$I_{v_x^2}^{(y)} I_{v_x^2}^{E_{16}}(\mathbf{x}, t) = -\frac{8p^2}{\pi^2 m^2} \sum_{n_x = -\infty}^{\infty} I_1, \tag{42}$$

$$I_{v_y^2}^{E_{16}}(\mathbf{x},t) = -\frac{4p^2}{\pi^2 m^2} \sum_{n_x = -\infty}^{\infty} (I_1 + 3r^2 I_2), \tag{43}$$

$$I_{v_z^2}^{E_{16}}(\mathbf{x},t) = -\frac{4p^2}{\pi^2 m^2} \sum_{n=-\infty}^{\infty} (2I_1 + 3r^2 I_2), \tag{44}$$

in which $r = |n_x|a$.

Non-orientability with a dipole: conclusions

We begin by noting that the expressions (35)–(40) and (42)–(44) for the components of the orientability indicator for E_{17} and E_{16} topologies, respectively, are too involved to lend themselves to a straightforward interpretation. Nevertheless, something significant can be said: for a dipole located at $P_0 = (x, 0, z)$ all components of the orientability indicator for E_{17} are different from those for E_{16} because each summand in Eqs. (35), (39) and (40) contains the prefactor $(-1)^{n_x}$ which is absent from the corresponding Eqs. (42)–(44) for E_{16} . Since not much further can be read from our equations, in order to demonstrate our main result, which is ultimately stated in terms of patterns of curves for the orientability indicator, we begin by plotting figures for the

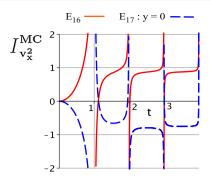


Fig. 3 Time evolution of the x-component of the orientability indicator in units of p^2/m^2 for a point electric dipole oriented in the flip y-direction in Minkowski spacetime with orientable E_{16} and nonorientable E_{17} spatial topologies, both with compact length a=1. The solid and dashed lines stand, respectively, for the indicator curves for a dipole in 3-space with E_{16} and E_{17} topologies. For the globally inhomogeneous topology E_{17} the dipole is at $P_0=(x,0,z)$, thus freezing out the topological inhomogeneity. Both orientability indicator curves show a periodicity, but the curve for E_{17} exhibits a different kind of periodicity characterized by a distinctive inversion pattern. Non-orientability is responsible for this pattern of successive inversions, which is absent in the indicator curve for the orientable E_{16}

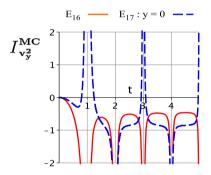


Fig. 4 Time evolution of the *y*-component of the orientability indicator for a point electric dipole oriented in the *y*-direction under the same conditions as those of Fig. 3. The orientability indicator curve for E_{17} also displays a characteristic inversion pattern but which is different from the one for the *x*-component shown in Fig. 3. For the *y*-component of the orientability indicator the signature of non-orientability can be recognized in the pattern of successive upward and downward "horns" formed by the dashed curve

components of the orientability indicator. Figures 3, 4 and 5 come from Eqs. (35)–(40) as well as (42)–(44), with the topological length a=1 and $n_x \neq 0$ ranging from -50 to 50. In the three figures the solid lines stand for the orientability indicator curves for the dipole in Minkowski spacetime with E_{16} orientable spatial topology, whereas the dashed lines represent orientability indicator curves for the dipole located at $P_0 = (x, 0, z)$ in a 3-space with E_{17} non-orientable topology.

In the case of the *x*-component, the time evolution curves of the orientability indicator for E_{16} and E_{17} , shown in Fig. 3, present a common periodicity but with clearly distinguishable patterns. The orientability indicator curve for E_{17} displays a distinctive sort of periodicity characterized by an inversion pattern. Non-orientability gives rise to this pattern



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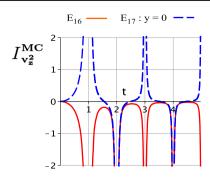


Fig. 5 Time evolution of the z-component of the orientability indicator for a point electric dipole oriented in the y-direction under the same conditions as those of Fig. 3. For the z-component of the orientability indicator the non-orientability of E_{17} manifests itself by an inversion pattern similar to the one for the y-component shown in Fig. 4, namely the pattern of alternating upward and downward "horns" formed by the dashed curve

of consecutive inversions, which is not present for the orientable E_{16} .

The differences become more salient when one considers the other two components of the orientability indicator, shown in Figs. 4 and 5. For both of these components, the nonorientability of E_{17} is disclosed by an inversion pattern whose structure is more striking than the one for the x-component. The orientability indicator curves for E_{17} form a pattern of alternating upward and downward "horns", making the nonorientability of E_{17} unmistakably identifiable.

From the above analysis of Figs. 3, 4 and 5 as compared with the corresponding analysis of Fig. 2, we see that the point dipole is potentially a much more efficient non-orientability probe than the point particle. Furthermore, the characteristic inversion pattern exhibited by the dipole indicator curves makes it possible to identify the non-orientability of E_{17} by itself, without having to make a comparison with the indicator curves for its orientable counterpart. We have checked, although we do not show the calculations, that for a dipole located at $P_0 = (x, 0, z)$ and oriented either in the x-direction or in the z-direction, only the y-component of the orientability indicator for E_{17} is different from the one for E_{16} . Thus, it is for a dipole oriented in the flip direction that the non-orientability of E_{17} is most sharply exposed.

This discussion suggests that it may be possible to unveil a presumed spatial non-orientability by local means, namely by the stochastic motions of point particles caused by quantum electromagnetic vacuum fluctuations. If the motion of a point electric dipole is taken as probe, non-orientability can be intrinsically discerned by the inversion pattern of the dipole's orientability indicator curves.

To close this section, some additional words of clarification regarding the use of the orientability indicator $I_{v_i^2}^{MC}(\mathbf{x},t)$ are fitting. We first note that, from a theoretical viewpoint, similarly to the way we calculated the orientability indica-

tor for E_{17} and E_{16} topologies [(35)–(40) and (42)–(44)] one can use the indicator's definition (16) to derive analytic expressions and plot corresponding curves for all eight non-orientable Euclidean topologies for a dipole. This set of theoretical curves for the orientability indicator would form a collection of curve patterns that are signatures of nonorientability, and which can be compared with experimental curves to find out about orientability of 3-space. From the experimental standpoint there are two main approaches to the use of $I_{n^2}^{MC}$. One is to contrast theory and observation through engineered experiments. This would be useful to check the different expressions we have derived for $I_{1,2}^{MC}$ against experimental data. In this case, one would need to manufacture devices that simulate non-trivial topologies. However, experimental apparatuses to mimic non-trivial topologies are not easy to prepare. 11

Another experimental approach is to use the indicator $I_{v_i^2}^{MC}$ to directly access orientability of the spatial section of Minkowski spacetime. First one would measure the velocity correlation functions $\langle \Delta v_i(\mathbf{x},t)\Delta v_i(\mathbf{x}',t)\rangle^{exp}$ for $\mathbf{x}\neq\mathbf{x}'$. Then one would subtract out the theoretically-computed correlation functions $\langle \Delta v_i(\mathbf{x},t)\Delta v_i(\mathbf{x}',t)\rangle^{SC}$ for topologically trivial (simply-connected) Minkowski spacetime, which are given in the Appendix. The corresponding curves for this difference would be plotted in the coincidence limit $\mathbf{x}=\mathbf{x}'$ and contrasted with each of the eight theoretical curves mentioned in the previous paragraph.

4 Conclusions and final remarks

In general relativity and quantum field theory spacetime is modeled as a differentiable manifold, which is a topological space equipped with an additional differential structure. Orientability is an important topological property of spacetime manifolds. It is often assumed that the spacetime manifold is orientable and, additionally, that it is separately time and space orientable. The theoretical arguments usually offered to assume orientability combine the space-and-time universality of local physical experiments ¹² with phys-

¹² The space universality can be looked upon as a topological assumption of global spatial homogeneity, which in turn rules out spatial non-orientability of 3-space.



 $^{^{11}}$ For a very simple nontrivial topology such as the 3-space reflection orbifold a perfectly reflecting plane should simulate the topology. Along these lines, one might wonder whether many equally spaced perfectly reflecting planes would simulate E_{16} . Simulation of nontrivial topologies in laboratory is indeed a thorny issue for experimentalists, for it depends on the physical phenomena involved. This issue is beyond the scope of the present work, but we mention that in a different context examples of simulations of non-trivial topologies such as the cylinder, the torus and the non-orientable Möbius strip have been recently discussed [55].

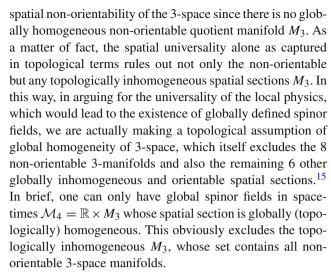
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ically well-defined (thermodynamically, for example) local arrow of time, violation of charge conjugation and parity (CP violation) and CPT invariance [56–58]. Another theoretical argument in favor of a time- and space-orientable spacetime comes from the impossibility of having globally defined spinor fields on non-orientable spacetime manifolds. One can certainly use such reasonings in support of the standard assumptions on the global structure of spacetime. Nevertheless, it is reasonable to expect that the ultimate answer to questions regarding the orientability of spacetime should rely on cosmological observations or local experiments, or might come from a fundamental theory of physics.

In the physics at daily and even astrophysical length and time scales, we do not find any sign or hint of nonorientability. This being true, the remaining open question is whether the physically well-defined local orientations can be extended continuously to cosmological scales.

At the cosmological scale, one would think at first sight that to disclose spatial orientability one would have to make a trip around the whole 3-space to check for orientation-reversing paths. Since such a global journey across the Universe is not feasible one might think that spatial orientability cannot be probed globally. However, a determination of the spatial topology through the so-called circles in the sky [61], for example, would bring out as a bonus an answer to the 3-space orientability problem at the cosmological scale. ¹⁴

There is no reason in the classical physics of point-like particles to force a spacetime manifold $\mathcal{M}_4 = \mathbb{R} \times M_3$ to be orientable. However, the situation changes once fermions, for example, are considered at the quantum level. Indeed, in quantum field theory one uses spinors to describe fermions. This fact inevitably leads to the inclusion of spinors as desirable objects (locally) in the spacetime \mathcal{M}_4 on physical grounds. However, the global question, which ultimately is not a physical requirement, is left open and the assumption of existence of globally defined spinor fields, which rules out non-orientable spacetimes [62–65], comes from theoretical arguments that combine this local physics requisite with the space-and-time universality of the basic local rules of physics. For spacetimes of the form $\mathcal{M}_4 = \mathbb{R} \times M_3$, the spatial universality (see related comment on p.288 of Visser [66]), can be mathematically translated in terms of global (topological) homogeneity of M_3 , which in turn rules out



In this paper we have addressed the question as to whether electromagnetic quantum vacuum fluctuations can be used to bring out the spatial orientability of Minkowski spacetime. To this end, we have studied the stochastic motions of point-like particles under quantum electromagnetic fluctuations in Minkowski spacetime with the orientable slab space (E_{16}) and the non-orientable slab space with flip (E_{17}) topologies (cf. Tables 1, 2).

For a point charged particle, we have derived analytic expressions for the orientability indicator, namely Eqs. (19)–(25) for E_{17} space topology, and Eqs. (26)–(27) for E_{16} space topology. From these equations we have made Figs. 1 and 2. Using these equations and figures we have shown that it is possible to distinguish the orientable from the non-orientable topology by contrasting the time evolution of the respective orientability indicators along the flip direction of E_{17} .

In spite of being a significant result in that it makes apparent the power of our approach to access orientability through electromagnetic quantum vacuum fluctuations, it is desirable, however, to be able to decide about the orientability of a given spatial manifold in itself. To tackle this question, motivated by a dipole's directional properties, we have then examined whether the study of stochastic motions of a point-like electric dipole would be more effective for testing the non-orientability of a generic 3-space individually, i.e. without having to make a comparison of the results for an orientable space with those for its non-orientable counterpart.

To this end, we have derived the expressions for the orientability indicators given by Eqs. (35)–(40) for the dipole oriented in the flip direction in the non-orientable 3-space



¹³ See Ref. [59] for a dissenting point of view, and also the related Ref. [60].

¹⁴ In the searches for these circles so far undertaken, including the ones carried out by the Planck Collaboration [16,17], no statistically significant pairs of matching circles have been found (see Ref. [15] for the most extensive search yet, and also references therein for the other searches). These negative observational results, however, are not sufficient to exclude the possibility that the Universe has a detectable (orientable or non-orientable) nontrivial topology (see Ref. [18] for some limits of these searches).

¹⁵ In technical terms, the ultimate mathematical point that leads to the fact that we cannot have globally defined spinor fields in non-orientable manifolds lies in bringing together the two-valuedness of spinors with the flip holonomy. Actually, this mismatching occurs even in some orientable, but inhomogeneous, quotient manifolds M_3 , whose covering or holonomy group contains at least one $\gamma \in \Gamma$ which is a proper screw motion $R(\alpha, \widehat{\mathbf{u}})$ with $\alpha \neq 0$ (cf. Section II).

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with E_{17} topology, and Eqs. (42)–(44) for the dipole in the orientable 3-space with E_{16} topology. From these equations we have calculated and plotted Figures 3, 4 and 5. As a result of the analysis of these equations and figures we have found that there exists a characteristic inversion pattern exhibited by the orientability indicator curves in the case of E_{17} , signaling that the non-orientability of E_{17} might be identified per se. The inversion pattern of the orientability indicator curves for the dipole is a signature of the reflection holonomy, and ought to be present in the orientability indicator curves for the dipole in all remaining seven non-orientable topologies with flip, namely the four Klein spaces (E_7 to E_{10}) and those in the chimney-with-flip class (E_{13} to E_{15}). ¹⁶ Clearly the inversion patterns for the electric dipole change with the associated topology: different topologies give rise to curves for the orientability indicator with distinct inversion patterns.

A short comment on how our results extend to the quantum level is fitting. We have tacitly assumed that the charged object of interest is described by a well-localized wave packet. Our classical analysis is justified because it has been argued [16] that under suitable conditions the dispersion due to wave packet spreading is negligible as compared with the dispersion brought about by vacuum fluctuations of the electromagnetic field.

Observation of physical phenomena and experiments are fundamental to our understanding of the physical world. Our results indicate that it is potentially possible to locally unveil spatial non-orientability through stochastic motions of point-like particles under electromagnetic quantum vacuum fluctuations. In this way, the present paper may be looked upon as an intimation of a conceivable way to locally probe the spatial orientability of Minkowski empty spacetime. A different statistical indicator of the stochastic motions can certainly be conceived, but perhaps the most important point raised in this work is that the spatial orientability of spacetime can be locally probed by inquiring quantum vacuum electromagnetic fluctuations about the motion of point-like "charged" particles.

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Data Availability Statement This manuscript has associated data in a data repository. [Authors' comment: The data that support the findings of this study are available from the corresponding author upon reasonable request.]

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Appendix A: Velocity correlation functions for the simplyconnected 3-space

In this appendix we give the velocity correlations functions for simply-connected spatial section of Minkowski spacetime that are referred to on the last paragraph of Sect. 3, for both the point particle and the point dipole. These are calculated as follows. First replace D in Eq. (5) by D_0 as given by (6), thus obtaining the electric field correlation functions for the topologically trivial (simply connected) Minkowski spacetime. Then insert these electric field correlation functions with $\mathbf{x} \neq \mathbf{x}'$ into the right-hand side of Eq. (4), for the particle, and Eq. (30), for the dipole. With

$$r_x = x - x', \quad r_y = y - y', \quad r_z = z - z',$$
 (A.1)

and

$$r = \sqrt{(x - x')^2 + (y - y')^2 + (z - z')^2},$$
(A.2)

we have, for the particle:

$$\left\langle \Delta v_{x}(\mathbf{x}, t) \Delta v_{x}(\mathbf{x}', t) \right\rangle^{SC} = \frac{q^{2}t}{8\pi^{2}m^{2}r^{5}(t^{2} - r^{2})} \left\{ 4r(r_{y}^{2} + r_{z}^{2})t + (r^{2} - 3r_{x}^{2})(t^{2} - r^{2}) \ln \frac{(r - t)^{2}}{(r + t)^{2}} \right\}, \tag{A.3}$$

$$\left\langle \Delta v_{y}(\mathbf{x}, t) \Delta v_{y}(\mathbf{x}', t) \right\rangle^{SC} = \frac{q^{2}t}{8\pi^{2}m^{2}r^{5}(t^{2} - r^{2})} \left\{ 4r(r_{x}^{2} + r_{z}^{2})t + (r^{2} - 3r_{y}^{2})(t^{2} - r^{2}) \ln \frac{(r - t)^{2}}{(r + t)^{2}} \right\}, \tag{A.4}$$

$$\left\langle \Delta v_{z}(\mathbf{x}, t) \Delta v_{z}(\mathbf{x}', t) \right\rangle^{SC} = \frac{q^{2}t}{8\pi^{2}m^{2}r^{5}(t^{2} - r^{2})} \left\{ 4r(r_{x}^{2} + r_{y}^{2})t + (r^{2} - 3r_{z}^{2})(t^{2} - r^{2}) \ln \frac{(r - t)^{2}}{(r + t)^{2}} \right\}. \tag{A.5}$$

The corresponding results for the dipole oriented in the *y*-direction are



¹⁶ See Refs. [44,45] for the symbols, names and properties of Klein spaces and chimney-with-flip families of non-orientable topologies.

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$$(y) \left\langle \Delta v_{x}(\mathbf{x}, t) \Delta v_{x}(\mathbf{x}', t) \right\rangle^{SC} = -\frac{4p^{2}}{\pi^{2}m^{2}} \left[2I_{1} + 3(r^{2} - r_{x}^{2} + 6r_{y}^{2})I_{2} + 24(r^{2} - r_{x}^{2})r_{y}^{2}I_{3} \right], \quad (A.6)$$

$$(y) \left\langle \Delta v_{y}(\mathbf{x}, t) \Delta v_{y}(\mathbf{x}', t) \right\rangle^{SC} = -\frac{4p^{2}}{\pi^{2}m^{2}} \left[I_{1} + 3(r^{2} + r_{y}^{2})I_{2} + 24(r^{2} - r_{y}^{2})r_{y}^{2}I_{3} \right], \quad (A.7)$$

$$(y) \left\langle \Delta v_{z}(\mathbf{x}, t) \Delta v_{z}(\mathbf{x}', t) \right\rangle^{SC} = -\frac{4p^{2}}{\pi^{2}m^{2}} \left[2I_{1} + 3(r^{2} - r_{z}^{2} + 6r_{y}^{2})I_{2} + 24(r^{2} - r_{z}^{2})r_{y}^{2}I_{3} \right], \quad (A.8)$$

where the integrals I_1 , I_2 , I_3 are given by Eqs. (36)–(38).

References

- 1. G.F.R. Ellis, Gen. Relativ. Gravit. 2, 7 (1971)
- 2. M. Lachièze-Rey, J.P. Luminet, Phys. Rep. 254, 135 (1995)
- 3. G.D. Starkman, Class. Quantum Gravity 15, 2529 (1998)
- 4. J. Levin, Phys. Rep. 365, 251 (2002)
- 5. M.J. Rebouças, G.I. Gomero, Braz. J. Phys. 34, 1358 (2004)
- M.J. Rebouças, A Brief Introduction to Cosmic Topology, in *Proc. XIth Brazilian School of Cosmology and Gravitation, AIP Conference Proceedings*, vol. 782, ed. by M. Novello, S.E. Perez Bergliaffa (Americal Institute of Physics, Melville, 2005), p. 188, also: arXiv:astro-ph/0504365
- 7. J.P. Luminet, Universe 2(1), 1 (2016)
- 8. M.J. Rebouças, Detecting cosmic topology with primordial gravitational waves (**In preparation**) (2021)
- G.I. Gomero, M.J. Rebouças, R. Tavakol, Class. Quantum Gravity 18, 4461 (2001)
- G.I. Gomero, M.J. Rebouças, R.K. Tavakol, Class. Quantum Gravity 18, L145 (2001)
- G.I. Gomero, M.J. Rebouças, R. Tavakol, Class. Quantum Gravity 18, L145 (2001)
- G.I. Gomero, M.J. Rebouças, R. Tavakol, Int. J. Mod. Phys. A 17, 4261 (2002)
- 13. J. Weeks, Mod. Phys. Lett. A 18, 2099 (2003)
- B. Mota, M.J. Rebouças, R. Tavakol, Class. Quantum Gravity 20, 4837 (2003)
- P.M. Vaudrevange, G.D. Starkman, N.J. Cornish, D.N. Spergel, Phys. Rev. D 86, 083526 (2012)
- P.A.R. Ade et al., Planck Collaboration 2013, Astron. Astrophys. 571, A26 (2014)
- P.A.R. Ade et al., Planck Collaboration 2015, Astron. Astrophys. 594, A18 (2016)
- G. Gomero, B. Mota, M.J. Rebouças, Phys. Rev. D 94, 043501 (2016)
- J.A. Wolf, Spaces of Constant Curvature (McGraw-Hill, New York, 1967)
- W.P. Thurston, in *Three-Dimensional Geometry and Topology*, vol. 1, ed. by S. Levy (Princeton University Press, Princeton, 1997)
- 21. B.S. DeWitt, C.F. Hart, C.J. Isham, Physica **96A**, 197 (1979)
- 22. J.S. Dowker, R. Critchley, J. Phys. A 9, 535 (1976)
- 23. P.M. Sutter, T. Tanaka, Phys. Rev. D 74, 024023 (2006)
- 24. M.P. Lima, D. Muller, Class. Quantum Gravity 24, 897 (2007)
- D. Muller, H.V. Fagundes, R. Opher, Phys. Rev. D 66, 083507 (2002)

- D. Muller, H.V. Fagundes, R. Opher, Phys. Rev. D 63, 123508 (2001)
- 27. H. Yu, L.H. Ford, Phys. Rev. D 70, 065009 (2004)
- 28. H. Yu, J. Chen, Phys. Rev. D 70, 125006 (2004)
- 29. L.H. Ford, Int. J. Theor. Phys. 44, 1753 (2005)
- 30. H.W. Yu, J. Chen, P.X. Wu, JHEP 0602, 058 (2006)
- C.H.G. Bessa, V.B. Bezerra, L.H. Ford, J. Math. Phys. 50, 062501 (2009)
- 32. V. Parkinson, L.H. Ford, Phys. Rev. A 84, 06210 (2011)
- V.A. De Lorenci, C.C.H. Ribeiro, M.M. Silva, Phys. Rev. D 94, 105017 (2016)
- 34. M. Seriu, C.H. Wu, Phys. Rev. A 77, 022107 (2008)
- 35. G. Gour, L. Sriramkumar, Found. Phys. 29, 1917 (1999)
- 36. M.T. Jaekel, S. Reynaud, Quantum Opt. 4, 39 (1992)
- C.H.G. Bessa, M.J. Rebouças, Class. Quantum Gravity 37, 125006 (2020)
- 38. E. Feodoroff, Russ. J. Crystallogr. Mineral. 21, 1 (1885)
- 39. L. Bieberbach, Math. Ann. 70, 297 (1911)
- 40. L. Bieberbach, Math. Ann. 72, 400 (1912)
- 41. W. Novacki, Comment. Math. Helv. 7, 81 (1934)
- 42. C. Adams, J. Shapiro, Am. Sci. 89, 443 (2001)
- 43. B. Cipra, What's Happening in the Mathematical Sciences (American Mathematical Society, Providence, 2002)
- A. Riazuelo, J. Weeks, J.P. Uzan, R. Lehoucq, J.P. Luminet, Phys. Rev. D 69, 103518 (2004)
- 45. H. Fujii, Y. Yoshii, Astron. Astrophys. 529, A121 (2011)
- 46. J. Anandan, Phys. Rev. Lett. 81, 1363 (1998)
- 47. J.R. Weeks, *The Shape of Space*, 3rd edn. (CRC Press, Boca Raton, 2020)
- 48. N.D. Birrel, P.C.W. Davies, *Quantum Fields in Curved Space* (Cambridge University Press, Cambridge, 1982)
- 49. J. Chen, H.W. Yu, Chin. Phys. Lett. 37, 2362 (2004)
- R. Lehoucq, M. Lachièze-Rey, J.-P. Luminet, Astron. Astrophys. 313, 339 (1996)
- G.I. Gomero, M.J. Rebouças, A.F.F. Teixeira, Class. Quantum Gravity 18, 1885 (2001)
- G.I. Gomero, A.F.F. Teixeira, M.J. Rebouças, A. Bernui, Int. J. Mod. Phys. D 11, 869 (2002)
- 53. M.J. Rebouças, Int. J. Mod. Phys. D 9, 561 (2000)
- 54. V.A. De Lorenci, C.C.H. Ribeiro, JHEP **1904**, 072 (2019)
- O. Boada, A. Celi, J. Rodríguez-Laguna, J. Latorre, M. Lewenstein, New J. Phys. 17, 045007 (2015)
- 56. Y.B. Zeldovich, I.D. Novikov, JETP Lett. 6, 236 (1967)
- S.W. Hawking, G.F.R. Ellis, The Large Scale Structure of Space-Time (Cambridge University Press, Cambridge, 1973)
- R. Geroch, G.T. Horowitz, Global structure of spacetimes, in *General Relativity: An Einstein Centenary Survey*, ed. by S. Hawking, W. Israel (Cambridge University Press, Cambridge, 1979), pp. 212–293
- M. Hadley, Testing the orientability of time. Preprints, 2018040240 (2018)
- 60. M. Hadley, Class. Quantum Gravity 19, 4565 (2002)
- N.J. Cornish, D. Spergel, G. Starkman, Class. Quantum Gravity 15, 2657 (1998)
- R. Penrose, The structure of space-time, in *Battelle Rencontres in Mathematics and Physics: Seattle, 1967*, ed. by C.M. DeWitt, J.A. Wheeler (Benjamin, New York, 1968)
- R. Penrose, W. Rindler, Spinors and Space-Time, Two-Spinor Calculus and Relativistic Fields, vol. 1 (Cambridge University Press, Cambridge, 1986)
- 64. R. Geroch, J. Math. Phys. **9**, 1739 (1968)
- 65. R. Geroch, J. Math. Phys. 9, 343 (1970)
- M. Visser, Lorentzian Wormholes from Einstein to Hawking (AIP, New York, 1996)

