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Burial and thermal maturity modeling of the Middle Cretaceous– Early Miocene petroleum system, Iranian sector of the Persian Gulf

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Abstract The Cretaceous Kazhdumi and Gurpi formations, Ahmadi Member of the Sarvak Formation, and Paleogene Pabdeh Formation are important source rock candidates of the Middle Cretaceous-Early Miocene petroleum system in the Persian Gulf. This study characterizes generation potential, type of organic matter, and thermal maturity of 262 cutting samples (marls and argillaceous limestones) from these rock units taken from 16 fields in the Iranian sector of the Persian Gulf. In addition, the burial and thermal histories of these source rocks were analyzed by one-dimensional basin modeling. Based on the total organic carbon and genetic potential values, fair hydrocarbon generation potential is suggested for the studied samples. Based on T_{max} and vitrinite reflectance values, the studied samples are thermally immature to mature for hydrocarbon generation. The generated models indicate that studied source rocks are immature in central wells. The Gurpi and Pabdeh formations are immature and the Ahmadi Member and Kazhdumi Formation are early mature in the western wells. The Pabdeh Formation is within the main oil window and other source rocks are at the late oil window in the eastern wells. The hydrocarbon expulsion from the source rocks began after deposition of related caprocks which ensures entrapment and preservation of migrated hydrocarbon.

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Keywords Persian Gulf · Kazhdumi Formation · Ahmadi Member · Gurpi Formation · Pabdeh Formation · Middle Cretaceous–Early Miocene petroleum system

1 Introduction

The Persian Gulf and its coastal areas (Fig. 1) contain the largest occurrence of crude oil in the world (Haghi et al. 2013) accounting for two-thirds of the world's proven oil reserves and approximately more than one-third of total proven world gas reserves (Rabbani 2007). Existence of repeated and extensive source rock beds, substantial carbonate and some sandstone reservoirs, excellent regional caprocks, huge anticlinal traps, and continuous sedimentation are the major factors making this region a remarkable area for hydrocarbon accumulations (Rabbani 2008).

The Middle Cretaceous-Early Miocene petroleum system is one of the five petroleum systems of the Zagros foldbelt and the Persian Gulf area (Bordenave and Hegre 2010). The Oligo-Miocene Asmari and Cretaceous Bangestan are the main reservoirs and the Cretaceous Kazhdumi Formation, Ahmadi Member of the Sarvak Formation, Gurpi Formation, and Paleogene Pabdeh Formation are important source rock candidates of this petroleum system. The evaporites of the Gachsaran Formation are cap rocks of this petroleum system. The Kazhdumi and Pabdeh formations are excellent source rocks and the Ahmadi Member and Gurpi Formation have been identified as marginal source rocks in the Dezful embayment (Bordenave and Burwood 1990; Bordenave and Huc 1995; Bordenave 2002; Bordenave and Hegre 2010; Rabbani and Tirtashi 2010; Alizadeh et al. 2012; Opera et al. 2013).

Despite the significant hydrocarbon accumulation in the Middle Cretaceous–Early Miocene petroleum system

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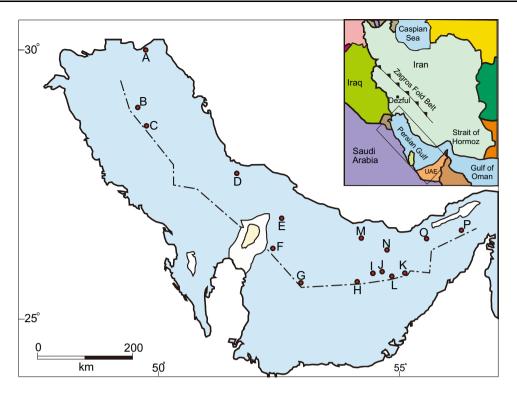


Fig. 1 Location of the studied fields in the Iranian sector of the Persian Gulf

within the Persian Gulf, little is known about the quality and maturity of the potential source rocks of this petroleum system in this area. This study tries to investigate hydrocarbon generation potential, depositional environment, and thermal maturity of the Kazhdumi Formation, Ahmadi Member, Gurpi Formation, and Pabdeh Formation in 16 fields located in the Iranian sector of the Persian Gulf (Fig. 1) by using Rock-Eval pyrolysis, molecular composition, and vitrinite reflectance measurement. Also, 1D basin modeling, a very useful tool in exploration-related studies, was applied to investigate the thermal maturity evolution and timing of hydrocarbon generation of these source rock candidates in the study area. The integration of the results of source rock characterization and basin modeling provides more detailed information to answer exploration questions. Accurate identification of a source rock helps to characterize the petroleum system and predict the location of future prospects charged by that source rock.

2 Geological setting

The Persian Gulf forms the northeast portion of the anticlockwise-moving Arabian Plate and formed during the Late Miocene (Alavi 2004). The Persian Gulf is situated at the junction of the Arabian and Eurasian lithospheric plates. It is structurally a foreland basin filled by terrigenous clastics transported from adjacent regions and carbonate sediments generated across the ramp surface (Ghazban 2009). Figure 2 shows the general lithostratigraphic column for the Iranian sector of the Persian Gulf.

During the Paleozoic, the Arabian Plate including the Persian Gulf region was located in the southern hemisphere with predominantly clastic sedimentation (Konert et al. 2001). Afterwards, during the Mesozoic and Cenozoic, the study area was mainly in tropical regions where carbonate deposition prevailed (Murris 1980; Ziegler 2001). Throughout most of the Mesozoic and up to the Lower Miocene, the area was part of a broad, shallow carbonate platform.

The Mesozoic carbonate systems of the Persian Gulf contain most of the extensive reservoir rocks in this area and form one of the richest hydrocarbon provinces in the world. This was mostly due to their vast-scale deposition and presence of source rocks, reservoir, and seal cap facies within the depositional system (Murris 1980). Within this time interval, in the Jurassic and Early Cretaceous periods, maximum marine transgression led to high production of organic matter, and its deposition under anoxic conditions forming organic rich deposits that were transformed over geologic time into petroleum source rocks (Alsharhan and Nairn 1997). Thick evaporites of the Gachsaran Formation, limestones and marls of the Mishan Formation followed by the sandstone, red marls, and siltstones of the Agha Jari Formation characterize the Mio-Pliocene of the region.

Era	Period	Epoch	Units	Lithology	Tectonic events	
		Pliocene	Bakhtiyari		700000	Legend
			Aghajari		Zagros collision	
<u>.</u>		Miocene	Mishan			Shale
ZO	iary		Gachsaran		Oligocene	
Cenozoic	Tertiary	Oligocene	Asmari		unconformity	
		Eocene	Jahrum		Zagros subduction	Evaporites
		Paleocene	Pabdeh			
			Gurpi	=================	First Alpine	Limestone
	snc	Upper	llam Laffan		orogeny (Ophiolite obduction)	
	Cretaceous	Mistella	Sarvak			
	reta	Middle	Ahmadi Kazhdumi		Turonian unconformity	Dolomitic
	0	Lower	Dariyan		uncomornity	limestone
ĿĊ		20110.	Fahliyan			
Mesozoic		Upper	Hith			Dolomite
Mes	sic.	орры			Passive	
	Jurassic	Middle	Surmeh		margin	
	٦٢				U U	Salt
		Lower				
			Neyriz			
	ic.	Upper				Sandstone
	Triassic	Middle	Dashtak			
	ц	Lower	Kangan			
			Dalan		Neo-Tethys	Siltstone
	F	Permian	Faraghan	555555555	opening	
	Car	boniferous	i aragnari		Hercynian	
<u>ici</u>	D	evonian	Zakeen		orogeny	
Paleozoic		Silurian	Sarchahan		Hercynian unconformity	
ale	0	rdovician	Zardkuh	5-5-5-5-5-	uncomornity	
1 "		aoviolari	Mila		Passive	
	С	ambrian	Lalun Zaigun		margin	
-			Zaigun Barut			
oic			Hormoz		Najid transtension	
Proterozoic			Crystalline	M + M + + M + M * * * * * M * *		
rot			basement	* * * * * *	Arabian shield consolidation	
Ľ					consolidation	<u> </u>

Fig. 2 Generalized stratigraphic column of Iranian sector of the Persian Gulf (modified from Al-Husseini 2008)

Folding accompanied by syntectonic and post-tectonic molasses took place in Plio–Pleistocene (Rabbani 2013).

The morphology of the Persian Gulf is highly affected by the Qatar Arch (Aali et al. 2006). The Qatar Arch is a first-order structure that was created in the central Persian Gulf following the tectonic movements during the Late Precambrian to Early Cambrian in the region (Fig. 3). It is a very large (over 100 km wide and 300 km long) regional gentle anticline (Ziegler 2001). According to offshore seismic data in the study area, this structure has a northeast–southwest direction in the Iranian sector of the Persian Gulf and continues southwards to the Qatar peninsula (Perotti et al. 2011). As Fig. 4 demonstrates, the thicknesses of the Pabdeh, Gurpi and Kazhdumi formations, and Ahmadi Member significantly decrease toward the central parts of the study area with a noticeable thinning which can be due to the effect of the Qatar Arch Paleohigh during depositional time (Alsharhan and Nairn 1997).

Salt diapirism is another significant structural element in the Persian Gulf (<5–20 km in size) piercing the stratigraphic sequences at different levels. Anticlines and domes have been induced by deep-seated salt pillows and salt

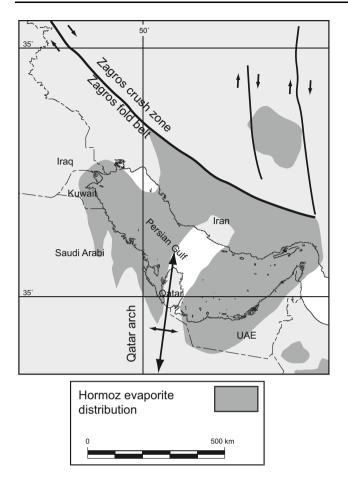


Fig. 3 Location of the Qatar Arch and distribution of Hormoz Salt in the study area (modified after Ghazban and Al-Aasm 2010)

ridges (Kent 1979). Almost all of this salt diapirism originates from the extrusion and remobilization of the Infra-Cambrian Hormoz Salt Series (Edgell 1991). Figure 3 shows the distribution of the Hormoz Salt in the Persian Gulf area.

2.1 Kazhdumi Formation

A transgression and sea-level rise in the Middle Cretaceous resulted in the deposition of the Kazhdumi Formation throughout the Albian (Alsharhan and Kendall 1991). In the Iranian offshore fields, the Kazhdumi Formation consists of calcareous shale and dark bituminous limestone with subordinate argillaceous limestone formed mostly in a neritic environment (Ghazban 2009). In addition, some thin sandstone beds may be present (Ghasemi-Nejad et al. 2009). The Burgan and Nahr Umr formations are the regional equivalents of the Kazhdumi Formation in the Arabian parts of the Persian Gulf (Rahmani et al. 2010). The underlying and overlying formations of the Kazhdumi Formation are the Sarvak and Dariyan formations, respectively (Rahmani et al. 2013).

2.2 Sarvak Formation

The Sarvak Formation is part of the Bangestan Group and deposited as a result of a significant transgressive phase in the Middle Cretaceous, after regional emergence and periods of clastic and deltaic sedimentation (Alsharhan and Nairn 1997). The carbonates in the Sarvak Formation blanket most of the Persian Gulf area. The bituminous shaly limestone of the Mauddud and Khatiyah members (in the central and western parts of the Persian Gulf), the Ahmadi Member with shaly facies in the northern Persian Gulf, and the Mishrif reefal limestone member in the southern Persian Gulf (Ghazban 2009) are four members of the Sarvak Formation. The Laffan shales overlay the Sarvak Formation with an unconformity surface and act as an efficient regional seal for the Sarvak Reservoir. The Kazhdumi Formation underlies the Sarvak Formation with a transitional contact.

2.3 Gurpi Formation

The Gurpi Formation consists of thin bedded, deep-marine marl, and marly limestone deposited when local dysoxic conditions occurred in the northern Persian Gulf region. The Gurpi Formation can act as a seal for the Ilam reservoir which underlies the Gurpi Formation with an erosional disconformity (Homke et al. 2009). The Upper Aruma and Bahrah–Tayarat are the Gurpi equivalents in the coastal Arabia and Kuwait areas (Rabbani 2013).

2.4 Pabdeh Formation

Neritic to basinal marls and argillaceous limestones of the Pabdeh Formation deposited in a Paleocene-Eocene transgression which resulted from the Late Cretaceous tectonic activities. This formation consists of shale, marl, and argillaceous limestones (Soleimani et al. 2013). A monotonous deep-water shale facies with a limestone unit in its middle part is the main lithology of the Pabdeh Formation. Based on lithological characteristics, the Pabdeh represents deposition in a deep-water, anoxic environment in an overall transgressive sequence. In the northern Persian Gulf, the Asmari and Gurpi formations overlay and underlie the Pabdeh Formation, respectively. There appears to be a transition to clean limestones of the Jahrum Formation toward the southwest (Sharland 2001). Regional equivalents of the Pabdeh Formation are the Umm er Radhuma, Rus, and Dammam formations.

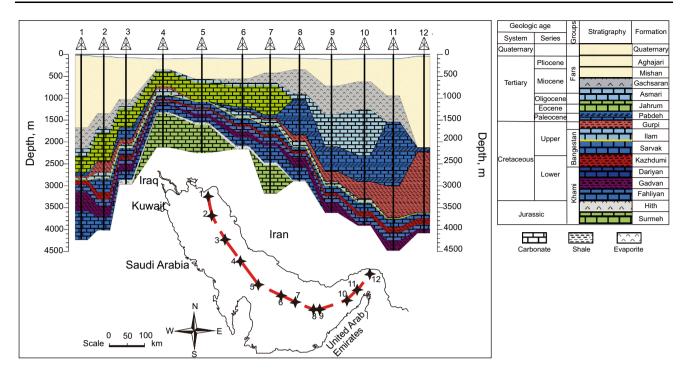


Fig. 4 Lithostratigraphic cross section in the Iranian sector of the Persian Gulf through Jurassic to Quaternary (modified after Rabbani et al. 2014)

3 Materials and methods

A total of 262 samples (including shale, marl, and argillaceous limestone) from the Kazhdumi, Gurpi, and Pabdeh formations and Ahmadi Member were taken from 16 fields within the Iranian sector of the Persian Gulf (Tables 1, 2, 3, 4, 5; Fig. 1). The selected samples were washed with water and detergent several times to remove contaminants from drilling mud additives. Then, the samples were crushed, pulverized, and homogenized. Rock-Eval pyrolysis was performed by a Vinci Rock-Eval 6 instrument in the AGH University of Poland on a 50 mg rock sample. See Espitalié et al. (1977), Lafargue et al. (1998) and Peters et al. (2005) for the details of this method. After completing the pyrolysis, the samples were heated to 850 °C at a rate of 25 °C/min in an oxidation oven and in the presence of air to oxidize (burn) all of the residual carbon. This process generates CO and CO₂ which are measured quantitatively. The parameters measured by this analysis included the total organic carbon (TOC) (wt%), S_1 (mg HC/g rock), S_2 (mg HC/g rock), T_{max} (the temperature at which the S_2 peak is the highest, °C), and S_3 (mg CO_2/g rock). Moreover, the hydrogen index (HI), oxygen index (OI), production index (PI), and migration index (S_1/TOC) were calculated.

Bitumen extractions were performed on approximately 10–15 g of 18 powdered samples from the Kazhdumi, Gurpi, and Pabdeh formations and also Ahmadi Member by using a Soxhlet apparatus for 72 h with an azeotropic mixture of dichloromethane (DCM) and methanol (CH₃OH) (93:7). The analyzed samples were selected based on higher organic matter content. The extracted bitumen in the rock samples was deasphalted by precipitation with *n*-hexane. Aliphatic, aromatic, and polar fractions were separated from the deasphalted samples by liquid column chromatography. The saturated fractions in the extracted bitumens were analyzed by gas chromatography (GC) in the AGH University of Poland. A Hewlett Packard 5890 series II GC held at a temperature of 300 °C equipped with a 50 m \times 0.2 mm Agilent DB1 column $(0.5 \ \mu m \ film \ thickness)$ with a constant flow of 0.3 mL/min of nitrogen as a carrier gas was used for this analysis. The column oven was programmed to hold at 30 °C for 5 min and then increase to 320 °C at a rate of 3 °C/min. The oven stays at 320 °C for 20 min. The components eluting the column were detected by a flame ionization detector (FID) held at 325 °C.

The vitrinite reflectance was measured in a random mode according to Taylor et al. (1998) and reported in $R_0\%$. Based on the amount of organic carbon present within the samples, 50 samples were selected for this analysis. The samples were mounted in resin, and then ground and polished using an alumina–ethanol slurry. The analysis was carried out with a Leitz-MPV-SP microscope

Field	Well	Lith. ^a	MD, m	TOC	$T_{\rm max}$	S_1	S ₁ /TOC	S_2	$S_1 + S_2$	S_3	PI	HI	<i>R</i> _o , %
М	M-1	ShMl	1341	0.37	436	0.52	1.41	0.65	1.17	0.87	0.44	176	
		ShMl	1320	0.43	432	0.40	0.93	0.52	0.92	0.97	0.43	121	0.35
	M-2	ShMl	1159	1.59	430	0.11	0.07	2.92	3.03	1.24	0.04	184	0.35
		ShMl	1195	1.36	429	0.12	0.09	2.61	2.73	1.13	0.04	192	
		ShMl	1242	1.25	430	0.17	0.14	1.11	1.28	1.26	0.13		
		ShMl	1280	1.33	431	0.10	0.08	1.46	1.56	1.37	0.06	110	
[I-1	Ml	2365	0.50	435	0.12	0.24	1.36	1.48	1.65	0.08	272	
		Ml	2355	0.43	433	0.15	0.35	1.37	1.52	1.71	0.10	319	
		Ml	2345	0.31	435	0.10	0.32	0.54	0.64	1.33	0.16	174	
		ShMl	2195	0.33	435	0.10	0.30	0.75	0.85	1.50	0.12	227	
		Ml	2165	0.37	434	0.12	0.32	1.08	1.20	1.54	0.10	292	
		Ml	2135	0.33	434	0.19	0.58	1.39	1.58	1.15	0.12	421	
		Ml	2100	0.39	433	0.12	0.31	1.47	1.59	1.34	0.08	377	
		LimMl	2080	0.34	434	0.10	0.29	0.97	1.07	1.36	0.09	285	
		LimMl	1955	0.55	432	0.12	0.22	2.89	3.01	1.09	0.04	525	
		Ml	1945	0.59	435	0.17	0.29	2.77	2.94	1.08	0.06	469	
		Ml	1895	0.52	433	0.12	0.23	2.24	2.36	0.83	0.05	431	
		Ml	1835	0.58	433	0.16	0.28	2.71	2.87	1.02	0.06	467	
		Ml	1825	0.62	438	0.19	0.31	2.61	2.80	0.96	0.07	421	
		Ml	1805	0.52	431	0.10	0.19	1.87	1.97	1.10	0.05	360	
		Ml	1795	0.46	434	0.13	0.28	1.60	1.73	1.17	0.08	348	
		Ml	1775	0.33	432	0.11	0.33	1.19	1.30	0.92	0.08	361	
		Ml	1765	0.51	434	0.17	0.33	1.57	1.74	0.67	0.10	308	
		Ml	1755	0.63	431	0.18	0.29	1.20	1.38	1.15	0.13	190	
		Ml	1725	0.31	434	0.11	0.35	0.97	1.08	0.86	0.10	313	
		Ml	1705	0.33	430	0.10	0.30	0.96	1.06	0.89	0.09	291	
		Ml	1695	0.45	433	0.08	0.18	0.81	0.89	1.17	0.09	180	
		LimMl	1625	0.33	436	0.14	0.42	0.93	1.07	1.82	0.14	282	
		Ml	1305	0.31	428	0.16	0.52	1.18	1.34	0.78	0.12	381	
	I-2	Ml	1763	0.38	439	0.14	0.37	0.62	0.76	0.70	0.18	163	
Χ	K-1	Ml	2875	0.35	430	0.13	0.37	0.68	0.81	1.79	0.16	194	
		Ml	2865	0.40	433	0.11	0.28	0.86	0.97	1.25	0.12	215	
		Ml	2715	0.96	425	0.21	0.22	5.20	5.41	1.43	0.04	542	
		Ml	2485	0.53	427	0.11	0.21	1.34	1.45	1.71	0.07	253	
		Ml	2485	0.84	433	0.14	0.17	2.38	2.52	1.91	0.06	283	
		Ml	2455	1.31	433	0.23	0.18	4.54	4.77	1.77	0.05	347	
		Ml	2405	1.44	431	0.26	0.18	5.86	6.12	1.57	0.04	407	
		Ml	2375	1.50	430	0.27	0.18	8.91	9.18	1.66	0.03	594	0.45
		Ml	2298	1.04	434	0.26	0.25	4.69	4.95	1.40	0.05	451	
		Ml	2260	0.74	428	0.20	0.27	2.96	3.16	1.27	0.06	400	
	J-1	MlSh	2046	0.39	432	0.14	0.36	0.50	0.64	1.06	0.22		
		MlSh	2040	0.71	434	0.29	0.41	1.89	2.18	1.36	0.13		
	P-3	Ml	2344	0.43	440	0.16	0.37	0.57	0.73	1.46	0.22	132	
		Ml	2329	0.34		0.15	0.44	0.51	0.66	1.45	0.23	149	
		Ml	2316	0.45	433	0.17	0.38	0.55	0.72	1.47	0.24	122	

Table 1 Rock-Eval and vitrinite reflectance data for the Pabdeh samples

Table 1 continued

Field	Well	Lith. ^a	MD, m	TOC	$T_{\rm max}$	S_1	S ₁ /TOC	S_2	$S_1 + S_2$	S_3	PI	HI	<i>R</i> _o , %
		Ml	2310	0.33	431	0.16	0.49	0.50	0.66	0.96	0.24	152	0.75
		Ml	2272	0.37	431	0.17	0.46	1.33	1.50	1.75	0.11	363	
		Ml	2268	0.33	433	0.18	0.55	0.51	0.69	1.16	0.26	155	
		Ml	2264	0.36	434	0.19	0.52	0.87	1.06	1.49	0.18	240	0.68
		Ml	2252	0.36	430	0.17	0.47	0.62	0.79	0.96	0.22	172	
		Ml	2242	0.43	431	0.19	0.44	1.10	1.29	1.18	0.15	257	
		Ml	2212	0.36	436	0.23	0.64	0.78	1.01	0.94	0.23	216	
		Ml	2204	0.37	431	0.18	0.49	0.61	0.79	1.50	0.23	164	
		Ml	2192	0.44	431	0.21	0.48	1.07	1.28	1.13	0.16	242	
		Ml	2182	0.42	432	0.24	0.57	0.72	0.96	1.35	0.25	172	0.70
		Ml	2172	0.48	436	0.21	0.44	0.93	1.14	1.21	0.18	193	
		Ml	2162	0.41	431	0.20	0.49	0.60	0.80	1.06	0.25	147	
		Ml	2152	0.61	430	0.29	0.47	1.45	1.74	1.36	0.17	237	
		Ml	2143	0.50	429	0.21	0.42	0.70	0.91	1.21	0.23	141	
		Ml	2132	0.45	418	0.30	0.67	1.29	1.59	1.40	0.19	287	0.63
		Ml	2113	0.43	429	0.23	0.53	0.56	0.79	0.94	0.29	130	
		Ml	2106	0.35	434	0.21	0.60	0.62	0.83	1.02	0.25	178	
		Ml	2090	0.30	430	0.18	0.60	0.73	0.91	0.62	0.20	242	0.60
		Ml	2070	0.35		0.16	0.46	0.52	0.68	0.65	0.24	149	0.62
		Ml	2054	0.47	438	0.15	0.32	0.57	0.72	0.49	0.21	122	
		Ml	2046	0.65	437	0.23	0.35	0.90	1.13	2.29	0.20	138	
		Ml	2034	0.45	433	0.22	0.49	0.82	1.04	1.88	0.21	184	
		Ml	2026	0.51	428	0.23	0.45	0.72	0.95	1.89	0.24	140	
		Ml	2014	0.71	420	0.30	0.42	1.19	1.49	1.46	0.20	168	
		Ml	2009	0.66	428	0.19	0.29	0.81	1.00	1.40	0.19	123	
		Ml	2005	0.62	426	0.21	0.34	0.90	1.11	1.17	0.19	145	
		Ml	1996	0.38	428	0.20	0.53	0.50	0.70	1.01	0.29	132	
		Ml	1994	0.70	433	0.29	0.41	1.12	1.41	3.35	0.21	160	0.63
		Ml	1986	0.81	428	0.27	0.33	1.32	1.59	1.83	0.17	162	
		Ml	1978	0.87	423	0.30	0.35	1.55	1.85	2.11	0.16	179	
		Ml	1970	0.75	426	0.29	0.39	1.30	1.59	1.77	0.18	174	
		Ml	1960	0.71	421	0.35	0.49	1.17	1.52	2.05	0.23	164	
		Ml	1946	0.56	426	0.27	0.49	0.74	1.01	1.89	0.27	133	
		Ml	1938	0.63	427	0.25	0.40	1.13	1.38	2.07	0.18	180	
		Ml	1930	0.59	428	0.25	0.42	0.91	1.16	2.15	0.22	153	0.58
		Ml	1922	0.75	426	0.22	0.29	1.16	1.38	1.52	0.16	155	
		Ml	1914	0.70	426	0.22	0.31	0.95	1.17	1.46	0.19	135	
		Ml	1898	0.46	433	0.19	0.41	0.73	0.92	1.38	0.21	158	
		Ml	1887	0.74	426	0.27	0.36	1.34	1.61	1.56	0.17	181	
		Ml	1878	0.76	426	0.28	0.37	1.18	1.46	1.61	0.19	156	
		Ml	1872	1.19	425	0.38	0.32	1.75	2.13	1.53	0.18	147	0.56
		Ml	1856	0.99	423	0.31	0.31	2.04	2.35	1.56	0.13	206	
		Ml	1846	0.98	426	0.31	0.32	1.52	1.83	1.97	0.17	155	
		Ml	1838	0.96	426	0.36	0.38	1.64	2.00	1.95	0.18	172	
		Ml	1819	0.93	425	0.34	0.37	1.45	1.79	1.93	0.19	156	0.54
		Ml	1808	0.83	427	0.41	0.50	1.48	1.89	1.63	0.22	179	
		Ml	1798	0.87	428	0.27	0.31	1.31	1.58	1.65	0.17	150	
		Ml	1788	0.87	436	0.26	0.30	1.18	1.44	1.50	0.18	135	

Table 1 continued

Field	Well	Lith. ^a	MD, m	TOC	$T_{\rm max}$	S_1	S ₁ /TOC	S_2	$S_1 + S_2$	S_3	PI	HI	<i>R</i> _o , %
		Ml	1773	1.27	428	0.32	0.25	2.11	2.43	1.86	0.13	166	0.48
		Ml	1736	1.49	431	0.40	0.27	2.66	3.06	1.69	0.13	179	0.46
		Ml	1728	1.47	430	0.38	0.26	2.71	3.09	1.79	0.12	184	
		Ml	1716	1.06	431	0.26	0.24	1.73	1.99	1.51	0.13	163	
		Ml	1702	1.25	429	0.22	0.18	2.16	2.38	1.60	0.09	173	
		Ml	1688	1.37	427	0.41	0.30	4.15	4.56	2.07	0.09	303	
		Ml	1676	0.99	430	0.26	0.26	1.91	2.17	1.80	0.12	194	
		Ml	1668	0.59	429	0.20	0.34	1.25	1.45	1.64	0.14	214	0.36
		Ml	1660	1.01	428	0.27	0.27	2.23	2.50	1.67	0.11	221	
		Ml	1652	0.67	428	0.31	0.46	1.60	1.91	1.71	0.16	239	
		Ml	1642	0.69	428	0.25	0.36	1.21	1.46	1.78	0.17	176	
		Ml	1632	0.83	428	0.32	0.39	1.88	2.20	1.67	0.15	227	
		Ml	1622	1.03	430	0.25	0.24	1.57	1.82	1.86	0.14	152	0.40
		Ml	1608	1.45	428	0.28	0.19	3.09	3.37	1.46	0.08	213	
		Ml	1596	1.56	427	0.27	0.17	3.12	3.39	1.95	0.08	200	0.48
		Ml	1588	1.58	425	0.43	0.27	4.57	5.00	1.65	0.09	290	
		Ml	1580	1.26	425	0.31	0.25	3.15	3.46	1.62	0.09	250	
		Ml	1572	1.66	427	0.68	0.41	5.15	5.83	1.76	0.12	310	
		Ml	1564	0.96	421	0.31	0.32	2.35	2.66	1.77	0.12	245	0.40
		Ml	1556	1.11	426	0.32	0.29	2.59	2.91	1.53	0.11	233	
		Ml	1544	1.07	416	0.40	0.38	2.59	2.99	1.78	0.13	243	
		Ml	1536	1.17	413	0.50	0.43	3.19	3.69	1.47	0.14	273	
		Ml	1528	0.98	404	0.36	0.37	2.33	2.69	1.35	0.13	237	
		Ml	1520	1.25	402	0.67	0.53	3.02	3.69	1.12	0.18	241	
		Ml	1512	1.50	410	0.66	0.44	3.56	4.22	1.59	0.16	238	0.55
0	O-2	Ml	2372	1.11	408	0.80	0.72	3.60	4.40	1.78	0.18	324	
		Ml	2397	1.35	413	0.56	0.41	5.43	5.99	2.05	0.09	402	
		Ml	2423	1.91	423	0.59	0.31	7.66	8.25	2.10	0.07	401	
		Ml	2443	1.98	422	0.46	0.23	7.51	7.97	1.64	0.06	379	
		Ml	2463	1.43	421	0.57	0.40	6.09	6.66	1.92	0.09	426	
		Ml	2488	1.30	427	0.37	0.28	4.00	4.37	3.26	0.08	308	
		Ml	2507	1.99	421	0.40	0.20	8.20	8.60	2.19	0.05	412	
		Ml	2529	2.64	419	0.51	0.19	10.62	11.13	2.69	0.05	402	
		Ml	2550	3.36	419	0.63	0.19	15.55	16.18	1.68	0.04	463	
		Ml	2567	3.00	424	0.49	0.16	13.51	14.00	2.07	0.04	450	
		Ml	2591	2.95	424	0.51	0.17	13.37	13.88	1.56	0.04	453	
		Ml	2608	2.67	425	0.49	0.18	12.12	12.61	1.47	0.04	454	0.56
		Ml	2658	0.86	428	0.19	0.22	2.43	2.62	1.91	0.07	283	
		Ml	2688	1.01	429	0.18	0.18	2.06	2.24	2.40	0.08	204	
		Ml	2702	0.49	429	0.22	0.45	1.22	1.44	1.90	0.15	249	
		Ml	2722	0.55	427	0.15	0.27	0.80	0.95	2.27	0.16	145	0.58
		Ml	2748	3.77	424	0.54	0.14	16.39	16.93	2.00	0.03	435	

The units of the Rock-Eval pyrolysis parameters and indices: TOC wt%, S_1 mg HC/g rock, S_2 mg HC/g rock, $S_1 + S_2$ mg HC/g rock, S_3 mg CO₂/g rock, T_{max} °C, HI mg HC/g TOC, OI mg CO₂/g TOC

^a ShMl is shaly marl, Ml is marl, Sh is shale, MlSh is marly shale, and LiMl is limy marl

 Table 2 Rock-Eval and vitrinite reflectance data for the Gurpi samples

Field	Well	Lith. ^a	MD, m	TOC	$T_{\rm max}$	S_1	S ₁ /TOC	S_2	$S_1 + S_2$	S_3	PI	HI	<i>R</i> _o , %
М	M-1	ShMl	1412	0.39	427	0.32	0.82	1.02	1.34	1.07	0.24	262	0.50
		ShMl	1403	0.53	435	0.57	1.08	1.23	1.80	1.10	0.32	233	
		ShMl	1393	0.47	430	0.54	1.15	2.21	2.75	0.88	0.20	470	
		ShMl	1387	0.39		0.49	1.26	0.58	1.07	1.16	0.46	149	0.47
		ShMl	1381	0.66	435	0.43	0.65	3.73	4.16	2.00	0.10	565	
	M-2	ShMl	1396	0.93	429	0.16	0.17	1.27	1.43	0.72	0.11	137	0.48
		ShMl	1426	0.84	426	0.11	0.13	0.96	1.07	0.66	0.10	114	
Κ	K-1	ShMl	3523	1.91	434	2.07	1.08	7.72	9.79	1.58	0.21	404	
		ShMl	3411	1.95	434	0.58	0.30	8.39	8.97	10.11	0.06	430	
		Ml	3353	1.92	431	0.48	0.25	8.05	8.53	11.56	0.06	419	
		Ml	3250	0.34	428	0.12	0.35	0.70	0.82	1.18	0.15	206	
		ShMl	3238	0.35	429	0.15	0.43	0.89	1.04	1.21	0.14	254	
		Sh	3230	0.74	432	0.12	0.16	1.46	1.58	1.76	0.08	197	
		Sh	3220	0.68	432	0.12	0.18	1.46	1.58	1.50	0.07	215	
		Sh	3210	0.75	431	0.12	0.16	1.48	1.60	1.78	0.08	197	
		Sh	3200	0.76	430	0.11	0.14	1.14	1.25	1.65	0.09	150	
		Sh	3190	0.71	430	0.15	0.21	1.19	1.34	2.04	0.12	168	
		Ml	3175	0.68	431	0.16	0.24	1.35	1.51	2.46	0.10	199	
		Ml	3165	0.67	433	0.13	0.19	1.58	1.71	1.59	0.07	236	
		Ml	3155	0.83	435	0.26	0.31	2.34	2.60	1.71	0.10	282	
		Ml	3145	0.48	434	0.15	0.31	1.39	1.54	1.47	0.10	290	
		Ml	3135	0.38	432	0.12	0.32	1.00	1.12	1.58	0.11	263	
		Ml	3115	0.44	433	0.12	0.27	1.33	1.45	1.53	0.08	302	
		Ml	3105	0.35	435	0.14	0.40	1.27	1.41	1.48	0.10	363	
		Ml	3085	0.60	433	0.16	0.27	1.76	1.92	1.87	0.09	293	
		Ml	3045	0.55	435	0.17	0.31	1.46	1.63	1.50	0.10	265	
J	J-1	MlSh	2348	0.88	435	0.23	0.26	1.07	1.30	1.32	0.18	992	
		MlSh	2345	0.90	433	0.31	0.35	1.14	1.45	1.43	0.21	947	
		MlSh	2340	0.78	430	0.18	0.23	0.75	0.93	1.29	0.19	972	
		MlSh	2335	0.96	433	0.29	0.30	1.31	1.60	1.53	0.18	986	
		MlSh	2140	0.39	426	0.32	0.82	0.60	0.92	1.24	0.35	786	0.55
Ι	I-2	ShMl	1891	1.62	430.0	0.41	0.25	3.46	3.87		0.11	214	
		ShMl	1942	0.62	434.0	0.21	0.34	0.58	0.79		0.27	94	
		Ml	1980	0.82	432.0	0.21	0.26	1.31	1.52		0.14	160	
	I-1	Sh	2696	0.80	426	0.12	0.15	0.70	0.82	1.50	0.14	88	0.56
		ShMl	2660	1.83	424	0.77	0.42	6.92	7.69	3.67	0.10	378	
		ShMl	2650	1.81	403	0.54	0.30	3.55	4.09	8.21	0.13	196	
		Ml	2640	0.68	413	0.35	0.51	1.43	1.78	3.51	0.20	210	
		ShMl	2625	0.50	432	0.12	0.24	1.38	1.50	2.13	0.08	276	
		ShMl	2605	0.62	433	0.12	0.19	1.97	2.09	2.16	0.06	318	
		ShMl	2595	0.47	433	0.11	0.23	1.33	1.44	1.94	0.08	283	
		LimMl	2585	0.63	430	0.10	0.16	1.06	1.16	1.85	0.08	168	
		Ml	2575	0.50	432	0.09	0.18	1.06	1.15	2.00	0.08	212	

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Table 2 continued

Field	Well	Lith. ^a	MD, m	TOC	$T_{\rm max}$	S_1	S_1/TOC	S_2	$S_1 + S_2$	S_3	PI	HI	$R_{\rm o}, \%$
		Ml	2565	0.51	434	0.10	0.20	0.96	1.06	1.86	0.10	188	
		Ml	2555	0.50	433	0.24	0.48	1.22	1.46	1.93	0.17	244	
		Ml	2535	0.77	431	0.24	0.31	2.43	2.67	2.10	0.09	316	
		Ml	2525	0.70	432	0.20	0.29	1.98	2.18	1.98	0.09	283	
		Ml	2515	0.69	431	0.20	0.29	1.88	2.08	1.98	0.09	272	
		Ml	2503	0.54	434	0.15	0.28	1.56	1.71	1.96	0.08	289	
		Ml	2495	0.81	432	0.18	0.22	1.74	1.92	2.58	0.09	215	
		Ml	2475	0.98	431	0.36	0.37	2.25	2.61	2.86	0.14	230	
		Ml	2445	1.17	430	0.50	0.43	3.32	3.82	3.81	0.13	284	
		Ml	2413	0.80	430	0.26	0.33	2.80	3.06	1.75	0.08	350	
		Ml	2405	0.84	432	0.22	0.26	2.71	2.93	1.82	0.08	323	
		Ml	2395	0.90	431	0.14	0.16	3.00	3.14	1.77	0.04	333	
		Ml	2385	0.99	431	0.17	0.17	3.13	3.30	1.79	0.05	316	
		Ml	2375	1.05	430	0.22	0.21	3.67	3.89	1.96	0.06	350	
0	O-2	Ml	2985	0.34		0.23	0.68	0.62	0.85		0.27	182	0.65
		Ml	3165	0.63	433	0.48	0.76	0.92	1.40		0.34	146	0.70
		Ml	3362	1.28	439	0.75	0.59	2.06	2.81		0.27	161	0.75

The units of the Rock-Eval pyrolysis parameters and indices: TOC wt%, S_1 mg HC/g rock, S_2 mg HC/g rock, $S_1 + S_2$ mg HC/g rock, S_3 mg CO₂/g rock, T_{max} °C, HI mg HC/g TOC, OI mg CO₂/g TOC

^a ShMl is shaly marl, Ml is marl, Sh is shale, and MlSh is marly shale

Table 3 Rock-Eval and vitrinite reflectance data for the Ahmadi sam	ples
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Field	Well	Lith. ^a	MD, m	TOC	$T_{\rm max}$	S_1	S ₁ /TOC	S_2	$S_1 + S_2$	S_3	PI	HI	<i>R</i> ₀ , %
J	J-1	Ml	2561	1.60	440	1.83	1.14	4.25	6.08	1.38	0.30	842	
		Ml	2561	0.52	432	0.37	0.71	0.61	0.98	1.24	0.38	750	0.67
Ι	I-2	Ml	2123	1.37	422	0.41	0.30	7.98	8.39	0.53	0.05	582	
		Ml	2138	2.68	424	1.19	0.44	16.94	18.13	0.70	0.07	632	
D	D-1	Ml	1924	0.42		0.08	0.19	0.36	0.44	2.01	0.19	86	0.39
		Ml	1846	0.30		0.11	0.37	0.21	0.32	1.58	0.34	70	0.36
С	C-1	MlSh	1930	1.25	428	0.78	0.62	1.66	2.44	2.15	0.32	133	0.50
		MlSh	1910	1.25	431	0.62	0.50	1.67	2.29	1.93	0.27	134	
		MlSh	1886	1.08	432	0.89	0.82	1.75	2.64	1.48	0.34	162	
		MlSh	1902	1.02	426	0.90	0.88	1.73	2.63	2.23	0.34	170	
В	B-1	ShMl	2520	0.45	432	0.21	0.47	0.64	0.85	0.85	0.25	143	
		ShMl	2510	0.51	433	0.18	0.35	0.70	0.88	0.87	0.20	138	
		MlSh	2490	0.98	435	0.16	0.16	1.11	1.27	1.21	0.13	113	
		MlSh	2450	1.34	425	0.35	0.26	5.42	5.77	1.16	0.06	404	
		Ml	2420	3.86	421	0.96	0.25	14.16	15.12	2.73	0.06	367	
		Ml	2380	0.38	435	0.16	0.42	0.52	0.68	0.73	0.24	138	0.57
0	O-2	Ml	3453	0.44	442	0.46	1.05	1.94	2.40	0.82	0.19	441	0.80
М	M-2	MlSh	1555	0.33	419	0.05	0.15	0.11	0.16	1.09	0.33	33	

The units of the Rock-Eval pyrolysis parameters and indices: TOC wt%, S_1 mg HC/g rock, S_2 mg HC/g rock, $S_1 + S_2$ mg HC/g rock, S_3 mg CO₂/g rock, T_{max} °C, HI mg HC/g TOC, OI mg CO₂/g TOC

^a Ml is marl, MlSh is marly shale, and ShMl is shaly marl

Table 4 Rock-Eval and vitrinite reflectance data for the Kazhdumi samples

Field	Well	Lith. ^a	MD, m	TOC	$T_{\rm max}$	S_1	S ₁ /TOC	S_2	$S_1 + S_2$	S_3	PI	HI	<i>R</i> _o , %
М	M-1	ShMl	1555	0.34	431	0.46	1.35	0.51	0.97	1.11	0.48	150	
	M-2	ShMl	1588	0.30	410	0.05	0.17	0.10	0.15	1.14	0.35	33	
F	F-1	Sh	1468	0.19	416	0.24	1.26	0.16	0.40	0.74	0.60	84	0.50
D	D-1	Ml	2026	0.39	425	0.19	0.49	0.71	0.90	1.77	0.21	182	0.41
С	C-1	LimMl	2108	0.72	422	0.75	1.04	1.00	1.75	2.33	0.43	139	0.47
		Sh	2091	0.86	423	0.85	0.99	1.23	2.08	2.89	0.41	143	
		Sh	2086	1.13	428	0.99	0.88	2.04	3.03	2.94	0.33	181	
		Sh	2070	1.68	428	1.62	0.96	3.47	5.09	1.61	0.32	207	
		Sh	2060	2.83	429	2.24	0.79	6.62	8.86	1.41	0.25	234	0.54
		MlSh	2033	2.79	423	1.48	0.53	6.45	7.93	2.47	0.19	231	0.47
		Sh	2029	2.64	426	1.27	0.48	4.44	5.71	2.84	0.22	168	
		Sh	2022	2.47	426	1.29	0.52	5.08	6.37	2.70	0.20	206	0.51
		Sh	2014	2.19	426	1.26	0.58	4.86	6.12	2.69	0.21	222	
		Sh	2006	1.99	430	1.19	0.60	3.84	5.03	2.34	0.24	193	0.56
		Sh	1998	1.94	430	1.00	0.52	3.49	4.49	2.94	0.22	180	
		Sh	1991	2.13	429	1.20	0.56	3.79	4.99	2.77	0.24	178	0.53
		MlSh	1987	1.33	430	0.86	0.65	2.24	3.10	2.14	0.28	168	0.61
		MlSh	1979	1.01	431	0.79	0.78	1.71	2.50	1.51	0.32	169	0.64
		MlSh	1970	0.70	430	0.75	1.07	1.05	1.80	1.61	0.42	150	
		Ml	1967	0.51	429	0.54	1.06	0.72	1.26	1.78	0.43	141	
		Ml	1952	0.34	425	0.57	1.68	0.62	1.19	1.65	0.48	182	
		Ml	1940	1.16	428	0.72	0.62	1.54	2.26	2.00	0.32	133	
В	B-1	Sh	2970	1.02	437	0.14	0.14	1.08	1.22	0.96	0.11	106	
		Sh	2960	1.17	435	0.20	0.17	1.76	1.96	1.26	0.10	151	0.69
		Sh	2950	1.34	437	0.29	0.22	2.09	2.38	1.32	0.12	157	
		Sh	2930	1.66	438	0.31	0.19	2.21	2.52	1.95	0.12	133	0.68
		Sh	2920	1.38	440	0.18	0.13	1.89	2.07	1.71	0.09	137	
		Sh	2910	0.95	440	0.18	0.19	1.32	1.50	0.92	0.12	138	0.66
		Sh	2900	0.80	461	0.10	0.13	0.50	0.60	1.44	0.17	63	
		MlSh	2890	1.22	434	0.38	0.31	2.25	2.63	1.00	0.14	184	0.66
		MlSh	2880	1.43	427	0.94	0.66	3.67	4.61	0.83	0.20	257	
		MlSh	2870	1.39	425	1.32	0.95	4.62	5.94	0.83	0.22	332	0.65
		MlSh	2860	0.83	435	0.21	0.25	1.34	1.55	1.06	0.14	161	
		MlSh	2850	0.78	434	0.28	0.36	1.27	1.55	1.01	0.18	162	
		MlSh	2840	0.96	433	0.24	0.25	1.74	1.98	1.15	0.12	181	0.64
		MlSh	2830	0.83	437	0.25	0.30	1.12	1.37	1.02	0.18	135	0101
		MlSh	2820	0.93	427	0.71	0.76	3.10	3.81	0.90	0.19	333	0.62
		MlSh	2800	0.74	436	0.15	0.20	0.44	0.59	1.48	0.25	59	0.02
		MlSh	2790	0.82	436	0.15	0.19	0.90	1.06	1.40	0.15	109	
		Sh	2780	1.03	436	0.10	0.19	1.30	1.50	1.14	0.13	126	0.62
		Sh	2770	1.02	437	0.20	0.19	1.10	1.28	1.14	0.13	107	0.02
		MlSh	2760	1.02	434	0.18	0.18	2.43	2.83	1.10	0.14	185	0.61
		MISh	2760 2750	0.78	434 436	0.40	0.31	2.43 0.96	2.83 1.17	1.07	0.14	185	0.01
		ShMl	2730 2740	0.78	436 429	0.21	0.27	1.28	1.17	0.85	0.18	124 178	0.60
0	O-2	ShMl			429 436	0.18					0.12	332	0.80
0	0-2	ShMl	3548 3595	0.57	436 464		0.56	1.89 3.81	2.21	2.11	0.14	552 544	0.80
			3595	0.70		0.18	0.26	3.81	3.99	1.43			
		ShMl	3603	0.54	463	0.17	0.31	3.21	3.38	0.91	0.05	594	0.82

Table 4 continued

Field	Well	Lith. ^a	MD, m	TOC	$T_{\rm max}$	S_1	S_1/TOC	S_2	$S_1 + S_2$	S_3	PI	HI	$R_{\rm o}, \%$
J	J-1	Ml	2612	0.77	423	0.41	0.53	3.97	4.38	0.30	0.09	516	
		Ml	2600	0.60	416	0.60	1.00	0.55	1.15		0.52		0.38
		Ml	2610	0.57	420	0.63	1.11	0.69	1.32		0.48	121	
		Ml	2620	0.71	422	1.06	1.49	1.24	2.30		0.46	175	
		Ml	2670	0.67	412	0.33	0.49	0.56	0.89	2.60	0.37		
		Ml	2676	0.92	429	0.61	0.66	1.31	1.92	2.20	0.32	142	
		Ml	2692	2.41	422	3.39	1.41	8.70	12.09	0.90	0.28	361	
		Ml	2720	1.44	429	0.99	0.69	4.35	5.34	1.20	0.19	302	
Ι	I-2	Ml	2211	0.77	423	0.41	0.53	3.97	4.38	0.31	0.09	516	
		Ml	2242	0.41	429	0.17	0.41	0.54	0.71	0.34	0.24	132	

The units of the Rock-Eval pyrolysis parameters and indices: TOC wt%, S_1 mg HC/g rock, S_2 mg HC/g rock, $S_1 + S_2$ mg HC/g rock, S_3 mg CO₂/g rock, T_{max} °C, HI mg HC/g TOC, OI mg CO₂/g TOC

^a ShMl is shaly marl, Ml is marl, Sh is shale, MlSh is marly shale, and LiMl is limy marl

in the organic petrography laboratory of Research Institute of Petroleum Industry (RIPI) in Iran. A sapphire glass standard with a 0.589 % reflectance value was used for calibration. The measurements were performed under reflected light at a wavelength of 546 nm with an oil immersion objective with $\times 125$ magnifications. At least, 50 readings were performed for each sample.

Basin modeling is a useful method for investigating the burial and thermal evolutions of sedimentary basins. In this study, six selected wells were modeled using PetroMod-1D modeling software (version 11, Schlumberger). The selected wells include B-1 (within the Field B, in the western part of the study area), C-1 (within the Field C, in the western part of the study area), D-1 (within the Field D, in the Central Persian Gulf), F-1 (within the Field F, in the Central Persian Gulf), J-1 (within the Field J, in the eastern part of the study area), and P-1 (within the Field P, in the eastern part of the study area). The location of the fields is shown in Fig. 1. Important 1D model input parameters involve the burial depths, thickness of the strata, erosion thickness and time, lithologies, kerogen types and kinetics, and further geochemical parameters such as the initial %TOC and HI (Table 6). The lithological information was inferred from unpublished well log data from National Iranian Oil Company (NIOC). The absolute ages were obtained from the timescale and regional chronostratigraphic subdivisions of Gradstein et al. (2004). The thermal evolution is modeled based on boundary conditions including the sediment water interface temperature (SWIT, in °C), paleo-water depth (PWD, in meter), and heat flow $(in mW/m^2).$

The upper boundary condition for calculating the temperature development in a sedimentary basin is the SWIT (Yalçin et al. 1997). The PetroMod-1D software

estimates the SWIT values through time based on the approach developed by Wygrala (1989). This estimation is based on the paleogeographical position of the area through geological time, variations in the mean surface paleo temperatures versus latitude and geological time; and water depth during the time of deposition (Yalçin et al. 1997). The SWIT calculations of this study were based on the paleo-latitude of the Northern Arabian plate.

The PWD values are required to calculate the SWIT. The PWD is dependent on combination of tectonic subsidence and changes in global sea levels. The depositional environment of each formation gives information about the PWD. A PWD of 0 m was considered for erosional events or phases of non-deposition and values of 20 m were applied for times of carbonate deposition. Negative PWD values were not used in this study.

The heat flow is the lower boundary condition of heat transfer into a sedimentary basin (Yalçin et al. 1997). It is an important input parameter in basin modeling and usually difficult to define for the geological past. Therefore, thermal history models are commonly calibrated against maturity and temperature profiles. In this study, bottomhole temperature and vitrinite reflectance data were used for the temperature and maturity calibrations, respectively. Recently, heat flow values in the range of 60–68 mW/m² were shown to be in accordance with the vitrinite reflectance measurements in the central Persian Gulf (Mohsenian et al. 2014). The easy $\Re R_0$ kinetic model of Sweeney and Burnham (1990) was applied to calculate the thermal maturity levels of the studied formations. Petroleum generation stages were calculated assuming mainly Type II kerogen and using a reaction kinetic dataset based on Burnham (1989).

Field	Well	Formation	Lith.	Depth, m	Pr/Ph	Pr/ <i>n</i> -C ₁₇	Ph/n-C18
I	I-1	Gurpi	Ml	2445	0.42	0.41	0.51
	I-1	Gurpi	Ml	2395	0.28	0.51	0.60
L	L-1	Gurpi	MlSh	2857	0.36	0.55	0.77
	L-1	Gurpi	ShMl	2767	0.31	0.51	0.73
Н	H-1	Gurpi	Sh	2199	0.12	0.44	0.69
Е	E-1	Kazhdumi	Sh	1631	0.20	0.56	0.94
G	G-1	Kazhdumi	ShMl	1504	0.30	0.36	0.63
N	N-1	Kazhdumi	Sh	2000	0.56	0.84	1.42
А	A-1	Kazhdumi	Sh	3070	0.41	0.44	0.79
С	C-1	Kazhdumi	Sh-MlSh	2048	0.42	0.62	0.94
	C-1	Kazhdumi	Sh	1999	0.59	0.53	0.94
	C-1	Kazhdumi	MlSh	1982	0.34	0.58	0.97
L	L-1	Pabdeh	Ml	2238	0.13	0.23	0.61
	L-1	Pabdeh	Ml	2137	0.39	0.55	0.81
	L-1	Pabdeh	Ml	1968	0.28	0.55	0.82
	L-1	Ahmadi Mbr.	MlSh	3148	0.27	0.38	0.59
Н	H-1	Ahmadi Mbr.	Sh	2574	0.12	0.37	0.80
G	G-1	Ahmadi Mbr.	Ml	1412	0.11	0.54	0.60

Table 5 Isoprenoids ratios measured for the studied samples

Table 6 Examples of input data for burial, thermal maturity, and hydrocarbon generation modeling in well P-1

Formation	Top, m	Base, m	Thickness, m	Eroded, m	Deposition from, Ma	Deposition to, Ma	Eroded from, Ma	Eroded to, Ma	Lithology
Bakhtiyari	70	100	30	0	10.87	1.5	1.5	0	Sandstone and marl
Upper Fars	100	914	814		13.3	10.87			Marl and Lime-Marly
Lower Fars	914	1200	286		20	13.3			Evaporite
Asmari	1200	1514	314		23	20			Limestone
Pabdeh	1514	2972	1458	200	59.33	35	35	23	Limestone and marl
Gurpi	2972	3695	723	200	81.01	64.02	64.02	59.33	Limestone and marl
Ilam	3695	3715	20		83.43	81.01			Limestone
Laffan	3715	3716	2		88.29	83.43			Shale
Sarvak-Mishrif	3716	3799	83	50	92.97	90.71	90.71	88.29	Limestone
Sarvak-Ahmadi	3799	3827	28		98	95			Limestone and marl
Sarvak-Mauddud	3827	3912	85		102.77	98			Limestone
Kazhdumi	3912	4045	133		115	102.77			Limestone and marl

4 Results and discussions

4.1 Rock-Eval data

Tables 1, 2, 3 and 4 and Fig. 5 show the results of the Rock-Eval pyrolysis in the studied wells. The cross-plot of S_1 versus TOC discriminates the nonindigenous and indigenous nature of the hydrocarbons present in the source rock samples (Hunt 1996). A migration index (S_1 /TOC) greater than 1.5 reveals that migrated hydrocarbons

affected the samples, whereas an index less than 1.5 points to an indigenous nature for the hydrocarbons. All of the studied samples have migration indices lower than 1.5, indicating that the analyzed samples were not polluted by migrated hydrocarbons (Fig. 6; Tables 1, 2, 3, 4).

The TOC content and genetic potential (summation of S_1 and S_2 peaks of Rock-Eval pyrolysis) provide important information about hydrocarbon generation potential of the source rocks. The TOC contents of the studied samples are in the range of 0.2–3.86 wt% with values generally lower

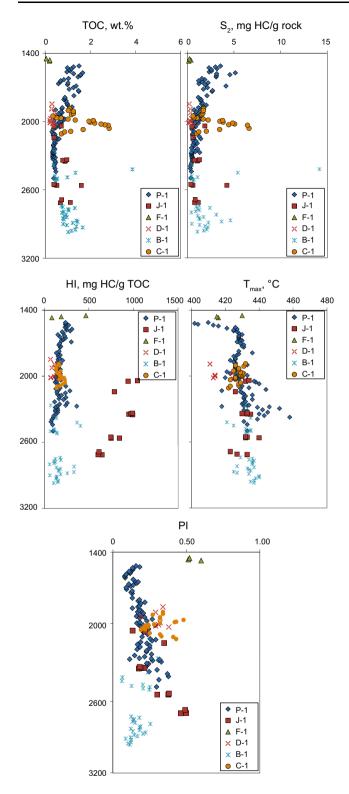


Fig. 5 Geochemical logs showing Rock-Eval data in the studied wells

than 2 wt% (Tables 1, 2, 3, 4). The genetic potential varies between 0.2 and 18.13 mg HC/g rock with values mostly lower than 6 mg HC/g rock (Tables 1, 2, 3, 4). The S_2

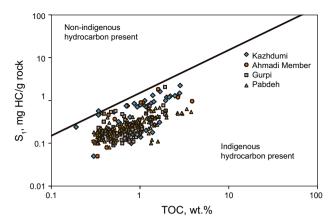


Fig. 6 Plot of S_1 versus TOC for distinguishing between indigenous and nonindigenous hydrocarbons present in the samples in the Kazhdumi (*right*), Gurpi (*middle*), and Pabdeh (*left*) formations (after Hunt 1996). The inclined line represents S_1 /TOC = 1.5

values measured for the studied samples are in the range of 0.5-17 mg HC/g rock with average of 2.5 mg HC/g rock. The cross-plots of genetic potential versus TOC (Fig. 7a) and S_2 versus TOC (Fig. 7b) indicate that the studied samples can be generally regarded as having a fair generative potential of hydrocarbon. HI values are in the range of 33-991 mg HC/g TOC (Tables 1, 2, 3, 4), and fair petroleum generation potential of the studied samples is also evident by cross-plot of HI versus TOC (Fig. 7c). Generally, the studied samples show lower TOC, genetic potential, and HI values in the central wells of the Persian Gulf (Tables 1, 2, 3, 4). This is in agreement with increasing the thickness of the studied source rocks from the Central Persian Gulf toward the eastern and western parts (Fig. 4). The occurrence of uplift in the Central Persian Gulf due to the presence of the Qatar Arch resulted in relatively poor preservation of organic matter in this part of the Persian Gulf compared to the adjacent areas.

The type of organic matter present in the source rocks can be evaluated based on the modified Van Krevelen diagram of HI versus T_{max} . The analyzed samples mainly plotted in the zone of mixed Type II-III kerogens and Type III kerogen of this diagram (Fig. 8a). Moreover, in the S_2 -TOC plot (Fig. 8b), most of the samples fall in the zone of mixed Types II-III kerogens grading to Type III. The local source rock evaluation study of Ghasemi-Nejad et al. (2009) in the South Pars Field also reveals this type of kerogen for the Kazhdumi Formation. This type of kerogen may originate from mixtures of terrigenous and marine organic matter with varying oil and gas generation potential or may also originally be a marine Type II organic matter which has partially been oxidized during deposition. As shown in the following section, the latter interpretation is the more likely.

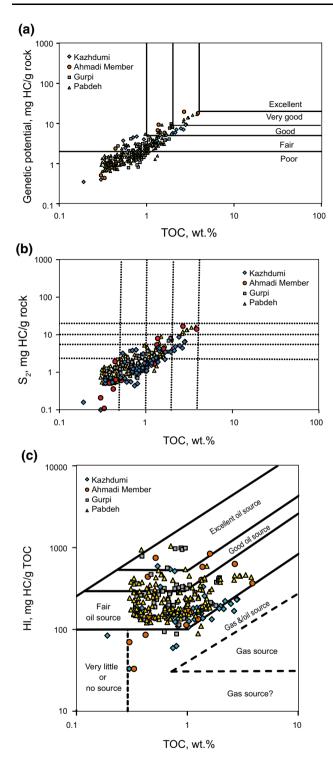


Fig. 7 Plots of **a** the genetic potential versus TOC, **b** S_2 versus TOC, and **c** HI versus TOC diagram showing source rock quality of the studied samples

4.2 Normal alkane and isoprenoids

All of the studied rock samples illustrate unimodal normal alkane distribution patterns, typically ranging from $n-C_{15}$

to $n-C_{34}$ (Fig. 9). Normal alkanes less than C_{14} are absent, probably because of evaporative loss during sample preparation. The maximum peak is generally detected in the range from $n-C_{17}$ to $n-C_{27}$. This normal alkane distribution pattern is characteristic of source rocks with strong input from marine organic matter. The pristane to phytane ratio (Pr/Ph) is considered as an indicator of the redox condition of the depositional environment. Low Pr/Ph ratios (<1) reflect an anoxic depositional environment, while greater values reveal more oxic conditions (Peters et al. 2005). In the studied samples, phytane is dominant over pristane and the Pr/Ph ratio displays values lower than 0.6 (Table 5). These values suggest marine reducing depositional conditions which is verified by $Pr/n-C_{17}$ lower than 1 and $Ph/n-C_{18}$ greater than 0.5 (Table 5). In the crossplot of Pr/n-C₁₇ versus Ph/n-C₁₈, the studied samples fall in the zone of marine Type II kerogen deposited under reducing conditions (Fig. 10).

4.3 Thermal maturity

The evaluation of thermal maturity of organic matter in the studied samples was carried out using vitrinite reflectance $(\% R_0)$, pyrolysis T_{max} , and production index (PI) values. All of the Rock-Eval S_2 values are greater than 0.5 mg HC/g rock (Tables 1, 2, 3, 4), so T_{max} is reliable for thermal maturity evaluation (Tissot and Welte 1984). The measured vitrinite reflectance values are in good agreement with pyrolysis T_{max} data (Fig. 11). The mean vitrinite reflectance values of the studied samples are in the range of 0.35–0.82 % R_0 (Tables 1, 2, 3, 4) showing thermally immature to mature stage of hydrocarbon generation in the analyzed samples. This is supported by T_{max} and PI values in the range of 402–464 °C and 0.03-0.35, respectively (Tables 1, 2, 3, 4). Thermal maturity has also been estimated by HI- T_{max} plot which indicates that the studied samples contain immature to late mature organic matter (Fig. 8a). The lowest vitrinite reflectance, T_{max} , and PI values were recorded in the central wells, while the highest values were measured for the samples from the eastern wells (Tables 1, 2, 3, 4). Generally, the Kazhdumi Formation and Ahmadi Member with older age and deeper burial are more thermally mature than the Pabdeh and Gurpi formations.

4.4 Burial and thermal history modeling

To generate reliable burial and thermal history models, the accurate timing and duration of erosional events should be fully constrained. As a consequence of eustatic sea-level changes and epeirogenic movements, several regional unconformities, erosion, and hiatuses occurred through the sedimentary succession of the study area (Sharland 2001).

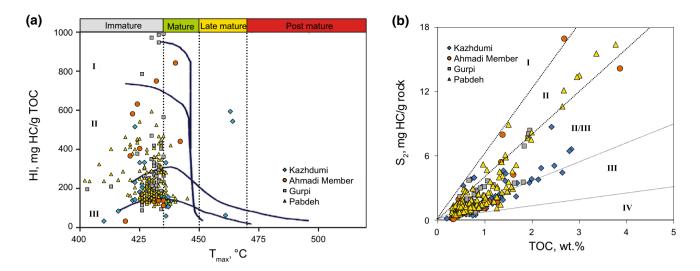


Fig. 8 a Modified Van Krevelen diagram of HI versus T_{max} and b S_2 versus TOC for the analyzed samples

In the Mid-Late Cretaceous and then in the Cenozoic, hiatus and erosion affected the studied area as a consequence of collision of the Arabian and Eurasian plates forming the Zagros Mountains (Zagros orogeny). The widespread Turonian unconformity occurred in the study area as a result of localized uplift (following the initiation of ophiolite obduction along the northeastern plate margin) and a global eustatic fall in sea level. The upper part of the Sarvak Formation is sometimes absent due to erosion during the Turonian unconformity. The pre-Neogene sediments are one of the most important erosional surfaces in the Tertiary sequences in the study area. Tectonics and eustasy integrated to cause a major relative sea-level fall through the Oligocene time resulting in widespread erosion and non-deposition across the entire region (Ghazban 2009). In general, total erosions between 50 and 200 m were considered in our models (Table 6). However, after carrying out a comprehensive sensitivity analysis and trying out several scenarios, the amount of erosions was found to have negligible impact on the present-day maturity and temperature trends. The results of the best fit models are presented here.

4.4.1 Well P-1, within the Field P

In well P-1, a constant heat flow value of 63 mW/m² (from the Late Cretaceous onward) gives the best fit between the measured and calculated vitrinite reflectance and bottomhole temperatures (Fig. 12a). Compared to other wells, the

studied source rocks have attained higher levels of maturity in the well P-1 possibly because of the deeper burial. The Ahmadi Member and Kazhdumi Formation, with maximum burial depths greater than 4000 m, are at the late oil window in this well with maximum burial temperatures of 152 and 160 °C and calculated vitrinite reflectance values of 1.2 $\% R_0$ and 1.25 $\% R_0$, respectively (Figs. 13, 14). The onset of the oil window (0.55–0.7 $\% R_0$) was in the Early Eocene (49 Ma) at a depth of approximately 1800 m (Fig. 12a). Within the Middle-Late Eocene (38 Ma) and at a depth greater than 2500 m, these source rocks entered the main oil window (0.7–1 $\% R_0$), and during the Late Miocene time (8 Ma), they reached the late oil window $(1-1.3 \ \% R_{o})$. The transformation ratio (TR), which is defined as the ratio of generated hydrocarbons to the total generation potential of a source rock (Shalaby et al. 2011), reached 96 % for the Kazhdumi Formation and Ahmadi Member (Figs. 13, 14). The Gurpi Formation has just reached the late oil window with a maximum burial temperature of 143 °C and calculated vitrinite reflectance of $1 \% R_0$ (Figs. 12a, 15). Hydrocarbon generation begun from the late Middle Eocene (40 Ma) at a burial depth of approximately 1800 m. Main oil generation occurred during the Late Oligocene (24 Ma) at a depth of approximately 2400 m. The Pabdeh Formation with a maximum burial temperature of about 124 °C and calculated vitrinite reflectance of 0.8 $\% R_{o}$ is interpreted to be within the main oil generation window (Figs. 12a, 16). Oil generation began in the Oligocene (30 Ma) at a burial depth of

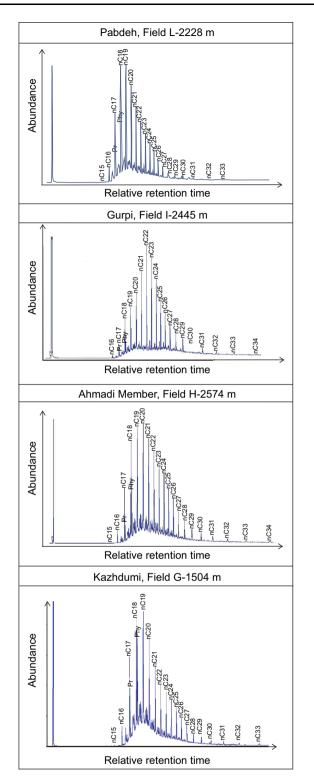


Fig. 9 Examples of gas chromatograms for the selected samples

approximately 1800 m and the main oil window occurred in the Late Miocene (7 Ma). The TR reached 84 % and 40 % for the Gurpi and Pabdeh formations, respectively.

4.4.2 Well J-1, within Field J

In well J-1, a constant heat flow value of 72 mW/m^2 (from the Late Cretaceous onward) gives the best fit between the measured and calculated vitrinite reflectance and bottomhole temperatures (Fig. 12b). Based on the burial/thermal history model, the Kazhdumi Formation and Ahmadi Member are in the main oil window in this well with maximum burial temperatures and calculated vitrinite reflectance values of approximately 115 °C and 0.75 %Ro, respectively (Figs. 12b, 13, 14). These source rocks reached the required levels of thermal maturity for the onset of the oil window from the late Middle Eocene at burial depths greater than 1500 m. At a burial depth greater than 2600 m, the Kazhdumi Formation and Ahmadi Member reached the main oil generation window in the Late Miocene and Pliocene, respectively. The TR of these source rocks reached approximately 20 % in this well. The Gurpi and Pabdeh formations are early mature with estimated maximum burial temperature and calculated vitrinite reflectance values of approximately 90 °C and 0.6 $\% R_{0}$, respectively (Figs. 12b, 15). The oil generation in the Gurpi Formation started from the Late Miocene (9 Ma) at a depth of 2100 m. The Pabdeh Formation has just reached the required thermal maturity for the hydrocarbon generation (Figs. 12b, 16). The generated hydrocarbons by these formations are not significant with TR values lower than 7 %.

4.4.3 Well F-1, within Field F

Well F-1 is modeled with a constant heat flow value of 71 mW/m² from the Late Cretaceous onward. This value led to a good match between the measured and calculated vitrinite reflectance and bottom-hole temperatures (Fig. 17a). All of the studied source rocks are thermally immature in this well, with estimated maximum burial temperatures lower than 80 °C and calculated vitrinite reflectance values lower than 0.5 $\% R_o$ (Figs. 13, 14, 15, 16).

4.4.4 Well D-1, within Field D

In well D-1, the best fit between the calculated and measured vitrinite reflectance and bottom-hole temperatures was obtained assuming a constant heat flow of 60 mW/m² for the Lower Cretaceous onward (Fig. 17b). In this well, the Pabdeh Formation is not present and the facies change to the Jahrum Formation. Other studied rock units are immature in this well with estimated maximum burial temperatures lower than 80 °C and calculated vitrinite reflectance lower than 0.5 % R_o (Figs. 13, 14, 15, 16). 384

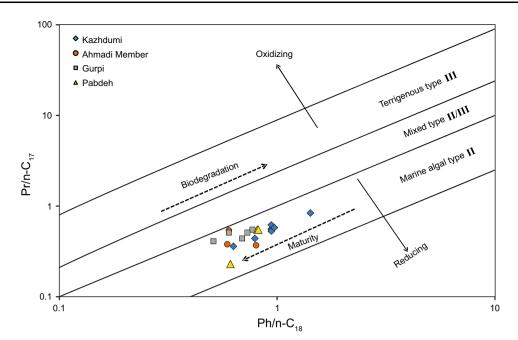


Fig. 10 Cross-plot of Pr/n-C₁₇ versus Ph/n-C₁₈ for the studied crude samples

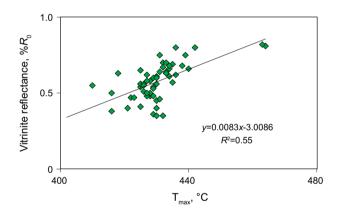


Fig. 11 Plot of T_{max} versus vitrinite reflectance for the studied samples

4.4.5 Well C-1, within Field C

In well C-1, a constant heat flow value of 73 mW/m² (from the Late Cretaceous onward) gives the best fit between the measured and calculated vitrinite reflectance and bottom-hole temperatures (Fig. 18a). The Pabdeh Formation changes to the Jahrum Formation and is not present in this well. The Kazhdumi Formation is at an early mature stage, with a maximum burial temperature of approximately 100 °C (occurred at Late Eocene) and calculated vitrinite reflectance of 0.65 % R_o (Fig. 13). The onset of the oil window was in the Middle Eocene (47 Ma) at a depth greater than 1400 m for this

formation (Fig. 18a). The generated hydrocarbon of the Kazhdumi Formation is not significant in this well location with TR lower than 10 % (Fig. 13). The Ahmadi Member and Gurpi Formation are thermally immature with maximum burial temperatures lower than 85 °C and calculated vitrinite reflectance lower than 0.5 % R_o (Figs. 14, 15).

4.4.6 Well B-1, within Field B

Well B-1 was modeled with a constant heat flow value of 72 mW/m^2 from the Late Cretaceous onward. With this value both the temperature and maturity trends have a reasonably good fit with the observed data (Fig. 18b). The Kazhdumi Formation and Ahmadi Member are interpreted to be early mature, with maximum burial temperatures of 106 and 97 °C and calculated vitrinite reflectance values of 0.67 % R_o and 0.62 % R_o , respectively (Figs. 13, 14). The onset of oil generation from the Kazhdumi Formation and Ahmadi Member occurred in the Oligocene (at a burial depth of 1500 m) and Late Miocene (at a depth of 1788 m), respectively. TR values lower than 10 % indicate that the generated hydrocarbons by these source rocks are not significant. The Gurpi and Pabdeh formations, with maximum burial temperatures lower than 90 °C and calculated vitrinite reflectance lower than 0.5 $\% R_0$, are thermally immature (Figs. 15, 16).

As discussed previously, the studied source rocks have fair hydrocarbon generation potential in the study area. In source rocks with fair generation potential, the necessary

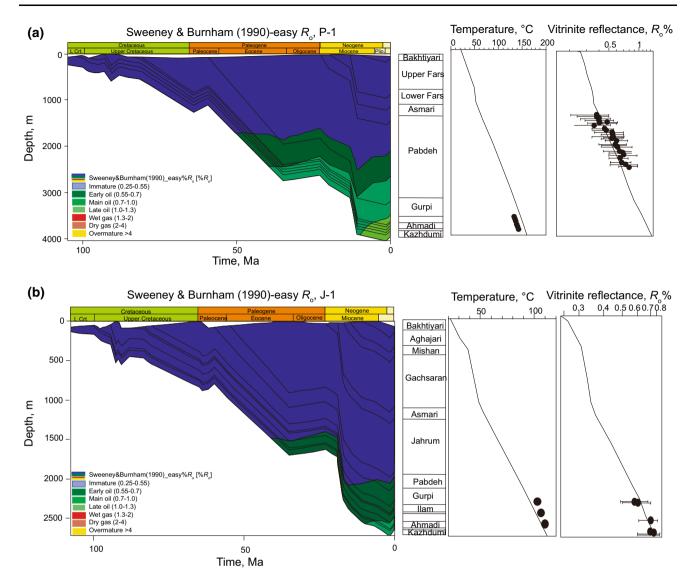
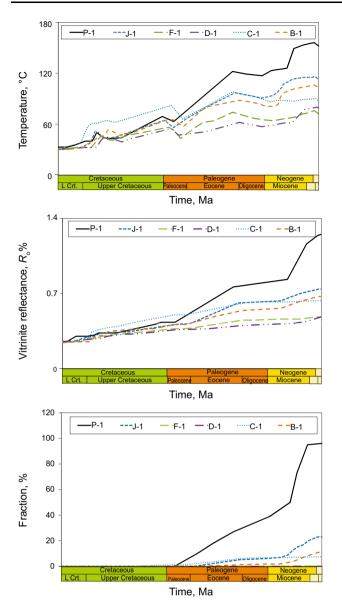


Fig. 12 Burial and thermal modeling of the eastern wells P-1 (top) and J-1 (bottom). The calibrations of the measured and calculated bottomhole temperature and vitrinite reflectance data are also shown

TR required to reach the onset of expulsion was estimated to be in the range of 45 %–55 % (Bordenave and Hegre 2010). Only in the well P-1, are the TR values of the Kazhdumi Formation, Ahmadi Member, and Gurpi Formation in the required range for hydrocarbon expulsion when start of the expulsion is in the Early to Late Miocene.

Due to shallower depth of burial and younger age, the Gurpi and Pabdeh formations generally have lower thermal maturity than the Kazhdumi Formation and Ahmadi Member in the study area. All of the studied source rocks are immature in the central wells. The presence of the Qatar Arch and distribution of Hormoz Salt in the Persian Gulf region can be possible reasons for lower thermal maturity in the central parts. The Qatar Arch has deformed the sedimentary cover by an order of magnitude more than the diapiric structures. A basement high is thus inferred in the core of the Qatar Arch and it has separated the Persian Gulf into northwest and southeast parts (Konert et al. 2001). The presence of this paleohigh at the central part of the Persian Gulf caused different burial depths for the studied rock units in the region such that they have shallower burial depth around the Qatar Arch, while being more deeply buried in the surrounding areas (Alsharhan and Nairn 1997). Lower thermal maturity of the studied formations in the central part can be a result of this lower burial depth in this part of the study area (Fig. 4).

Salt has a thermal conductivity two to four times greater than that of other sedimentary rocks (Bjørlykke 2010). It can have a great impact on the maturity of the organic matter and timing of hydrocarbon generation. Because the



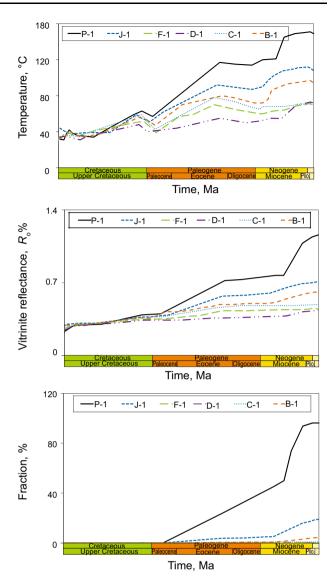


Fig. 13 Evolution of temperature, maturity, and transformation ratio (TR) for the Kazhdumi Formation in the investigated wells of the study area

heat is transferred more easily to the source rock units situated above the salt layers, they become more mature compared to adjacent source rocks not affected by salt. The salt diapirs of the Late Proterozoic Hormoz Formation in the area are present only in southeast and northwest of the Qatar Arch, while they are absent around the crest of this arch (Husseini 2000) (Fig. 3). The absence of salt-related phenomena around the crest of the arch is possibly due to the lack or reduced thickness of the Hormoz Formation in this region (Konert et al. 2001). So the presence of the Hormoz Formation in the southeast and northwest of the Qatar Arch can be considered as another possible reason for higher thermal maturity in these areas relative to the central part influenced by the Qatar Arch.

Fig. 14 Evolution of temperature, maturity, and transformation ratio (TR) for the Ahmadi Member in the investigated wells of the study

5 Summary and conclusions

area

The Kazhdumi Formation, the Ahmadi Member of the Sarvak Formation, and the Gurpi and Pabdeh formations were introduced as source rock candidates of the Middle Cretaceous–Early Miocene petroleum system in the Persian Gulf. In this study, hydrocarbon generation potential, depositional environment, and thermal maturity of 262

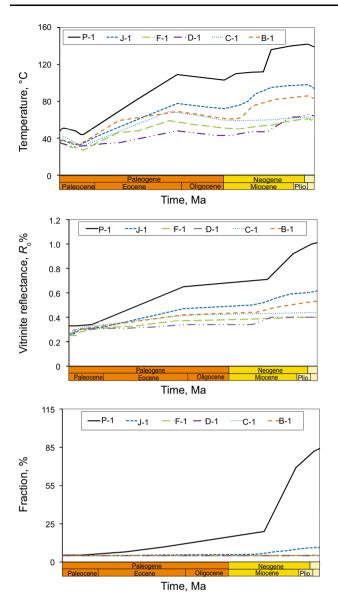


Fig. 15 Evolution of temperature, maturity, and transformation ratio (TR) for the Gurpi Formation in the investigated wells of the study area

cutting samples of these rock units were analyzed in 16 fields located in the Iranian sector of the Persian Gulf. Also, by using PetroMod 1D software, burial and thermal histories were modeled for six selected wells in the study area to analyze the thermal maturity evolution and hydrocarbon generation histories of the Kazhdumi Formation, Ahmadi Member, and the Gurpi and Pabdeh formations. Bottom-hole temperatures and measured vitrinite reflectance values were used for calibration of models.

• Based on Rock-Eval pyrolysis data and normal alkane distribution patterns, the studied source rock candidates

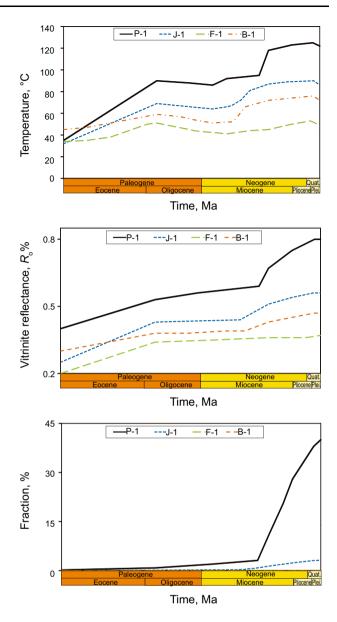


Fig. 16 Evolution of temperature, maturity, and transformation ratio (TR) for the Pabdeh Formation in the investigated wells of the study area

have fair hydrocarbon generation potential and deposited under marine reducing conditions with marine organic matter as the main input.

- Vitrinite reflectance, Rock-Eval pyrolysis *T*_{max}, and PI values indicate a wide range of maturities between thermally immature to mature for the studied rock units. The Kazhdumi Formation is more thermally mature than the other source rocks. The highest level of maturity is observed in the eastern parts of the study area.
- The constant heat flow values in the range of 63–73 (from the Late Cretaceous onward) give the best fit

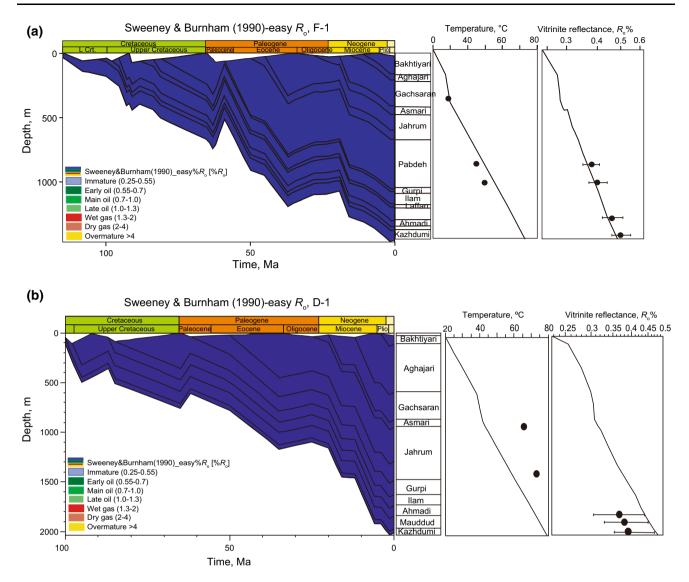


Fig. 17 Burial and thermal modeling of the central wells F-1 (top) and D-1 (bottom). The calibrations of the measured and calculated bottomhole temperature and vitrinite reflectance data are also shown

between the measured and calculated bottom-hole temperatures and vitrinite reflectance values in the studied wells.

- The studied source rock candidates are not sufficiently mature for hydrocarbon generation in the central wells (D-1 and F-1).
- The Kazhdumi Formation is early mature in the western wells (B-1 and C-1) and is in the main oil window in the eastern wells (J-1 and P-1). The hydrocarbon generation from the Kazhdumi Formation started from the Early Eocene, whereas the main phase of generation begun during Late Miocene.
- The results of the burial and thermal modeling indicate that the Ahmadi Member is immature in well C-1 and it is early mature in the B-1 well. In the eastern wells (P-1 and J-1), the Ahmadi Member is in the main oil

window. The oil generation from the Ahmadi Member may have begun from Early Eocene and the main oil window occurred in the Late Miocene.

- The Pabdeh and Gurpi formations are thermally immature for hydrocarbon generation in the western wells. They are early mature in well J-1 and are in the main oil window in well P-1. The hydrocarbon generation from the Gurpi Formation started in the Middle Eocene and the main phase of oil generation was in the Late Oligocene. The onset of oil generation from the Pabdeh Formation was in the Oligocene and the main oil window was within the Late Miocene.
- Due to the higher thermal maturity, the Kazhdumi Formation and Ahmadi Member probably have a more significant role in charging the reservoirs of the study area than the Gurpi and Pabdeh formations.

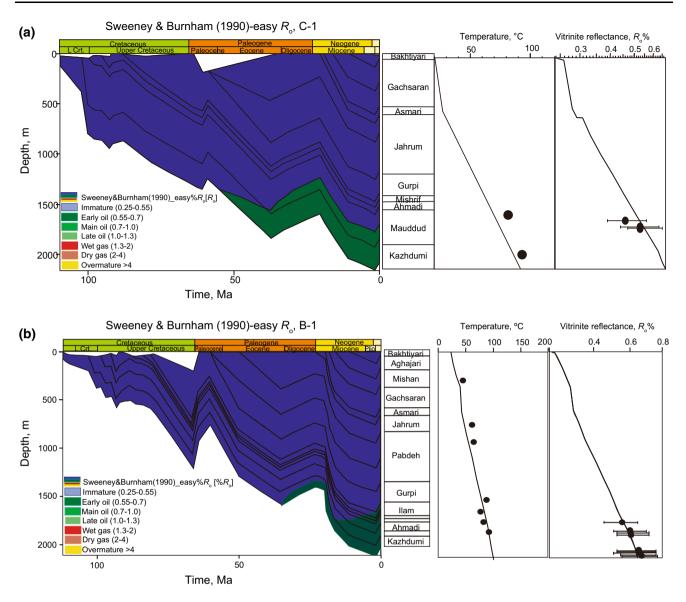


Fig. 18 Burial and thermal modeling of the central wells C-1 (top) and B-1 (bottom). The calibrations of the measured and calculated bottomhole temperature and vitrinite reflectance data are also shown

• The onset of oil expulsion in the studied source rocks was after deposition of related cap rocks, which allows the accumulation of generated hydrocarbons in the available reservoirs.

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